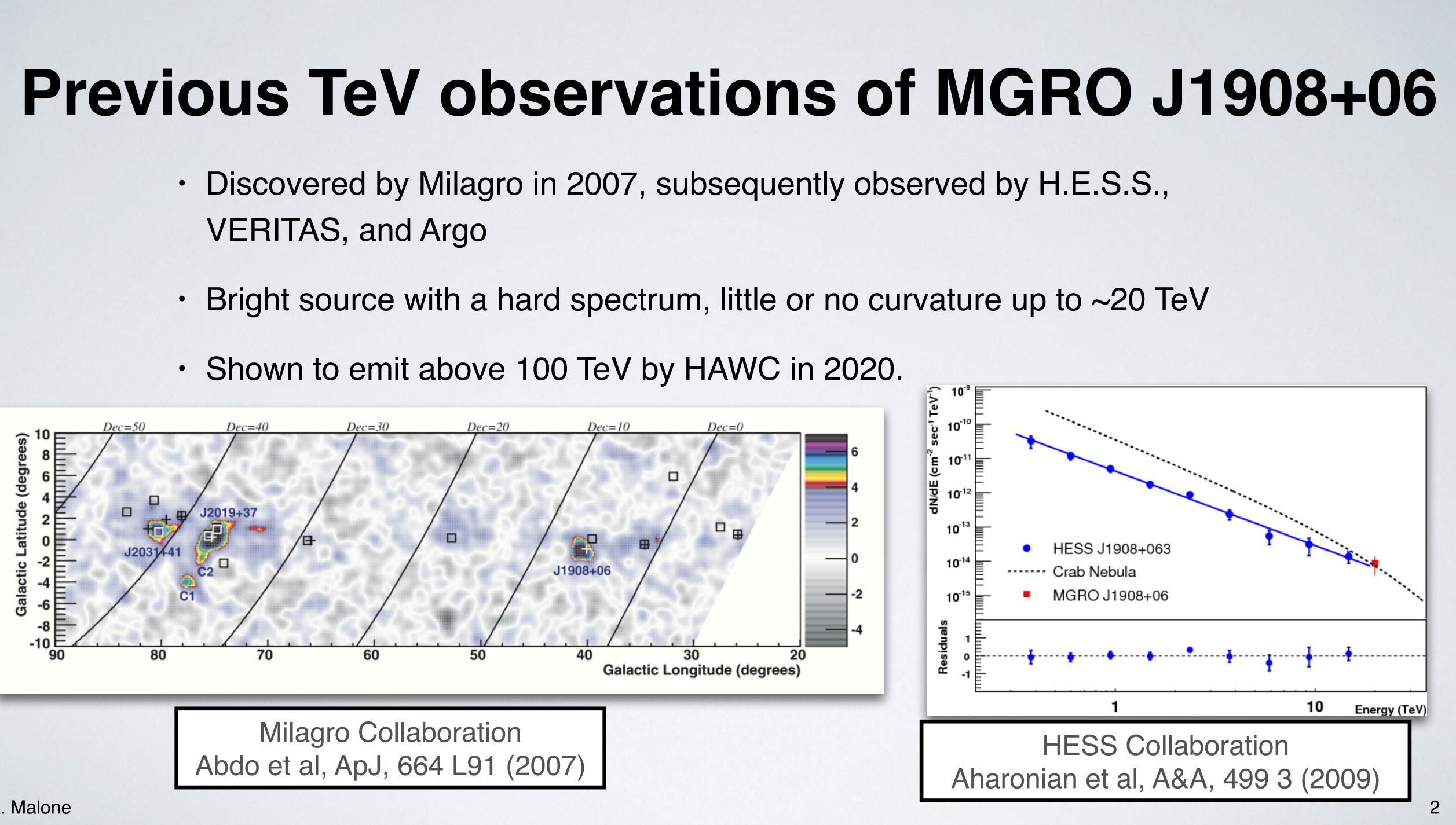
The Ultra-high-energy source MGRO J1908+06

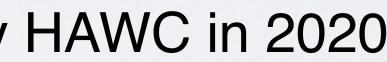


Kelly Malone Los Alamos National Laboratory ICRC 2021



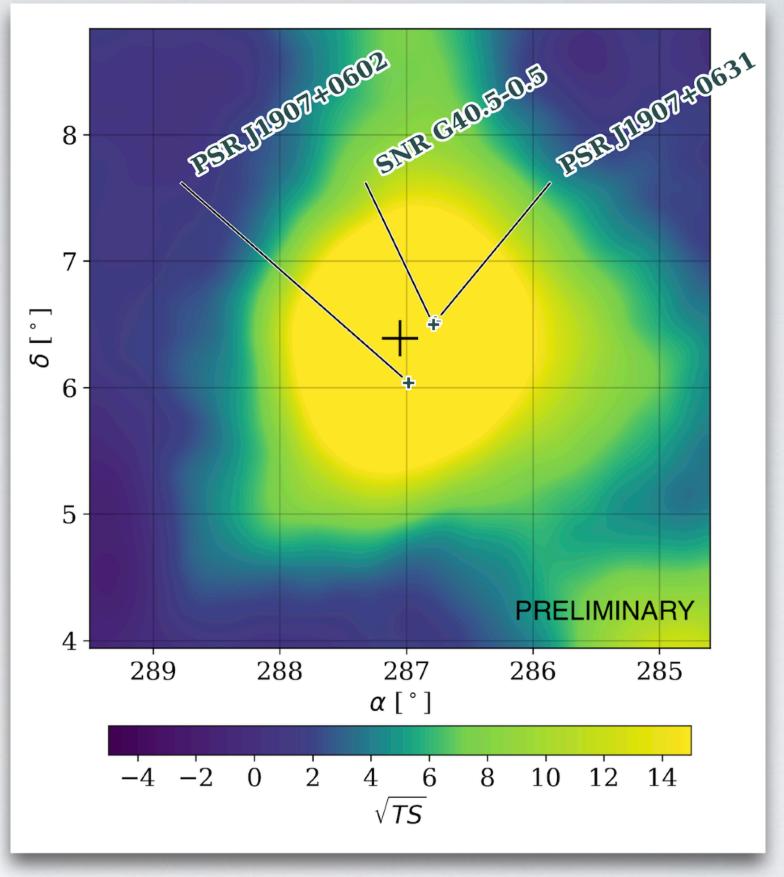
- **VERITAS**, and Argo
- Bright source with a hard spectrum, little or no curvature up to ~20 TeV



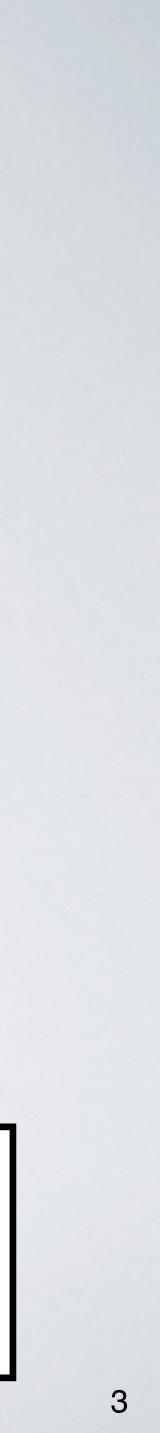


- Crowded region: Two pulsars, an SNR remnant, and molecular clouds are all present.
- PSR J1907+0602 is a young (~20 kyr) radio-quiet pulsar that is extremely high-Edot (2.8 x 10³⁶ ergs/s)
 - Historically the TeV emission has been attributed to this pulsar, but other objects could also contribute to the emission

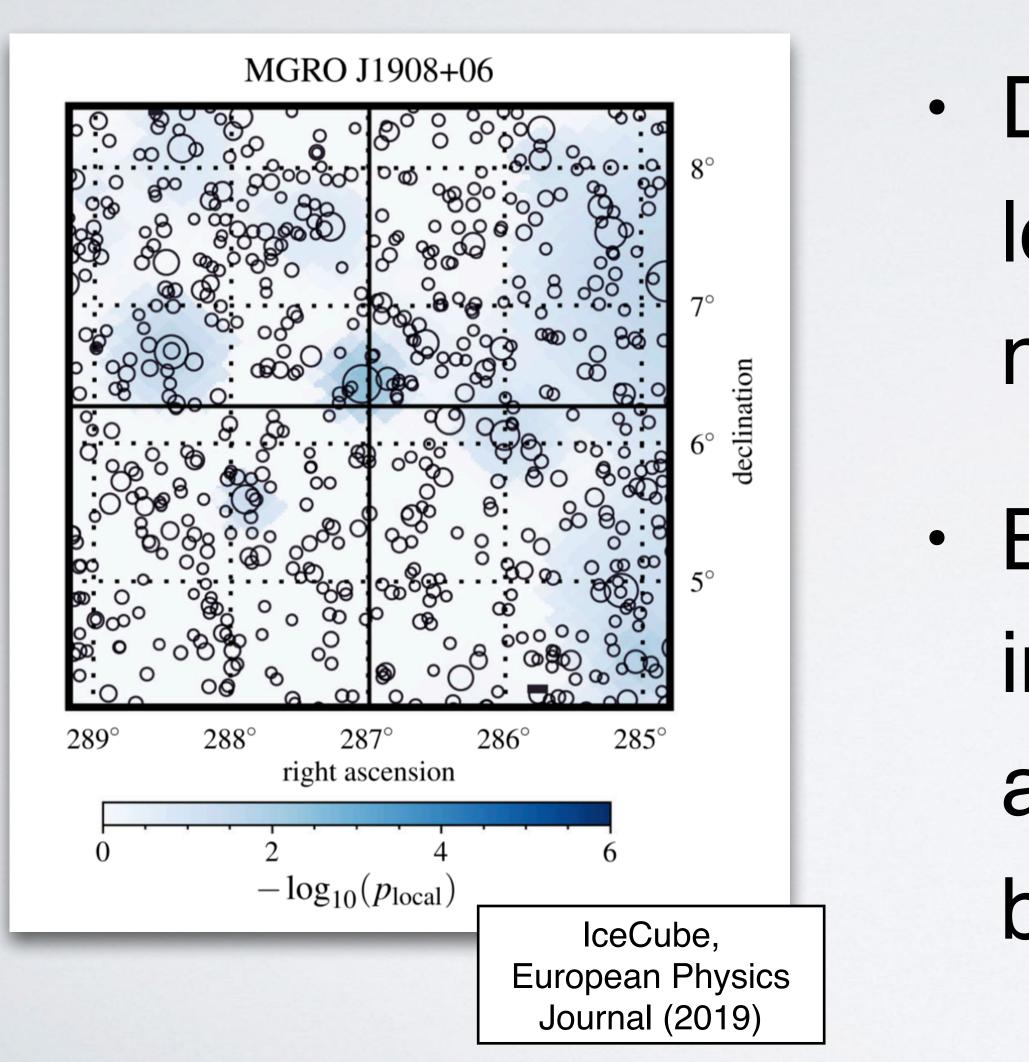
Possible counterparts



HAWC significance map with potential counterparts labeled. The cross is the center of the HAWC source



Multi-messenger observations



- Due to it's hard spectrum, it has long been considered a potential neutrino source.
- Best p-value for a Galactic source in IceCube catalog searches, although still consistent with background.



Introduction to HAWC



HAWC with Pico de Orizaba in the background

Area

Location

Altitude

Duty Cycle

Coverage

Sensitivity

Angular resolution 300 (4 PMTs/200,000 L of water in each)

22,000 m²

Puebla, Mexico (19° North)

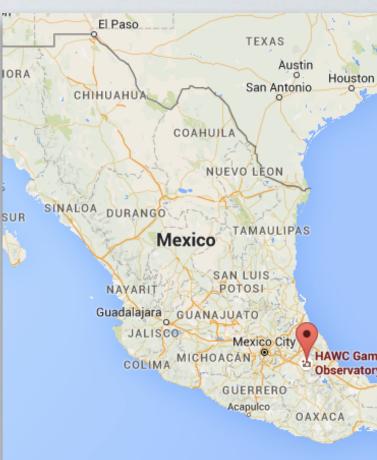
4100 m

> 95%

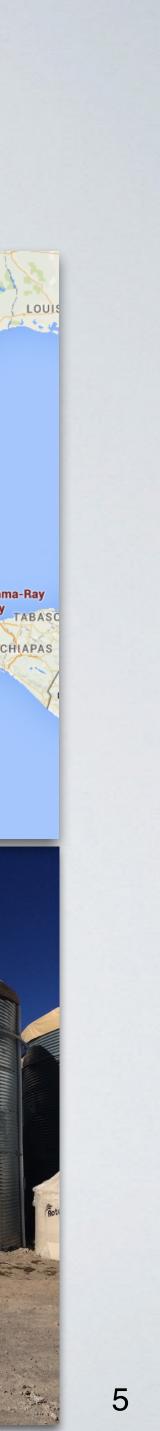
2/3 of sky per day

300 GeV to > 100 TeV

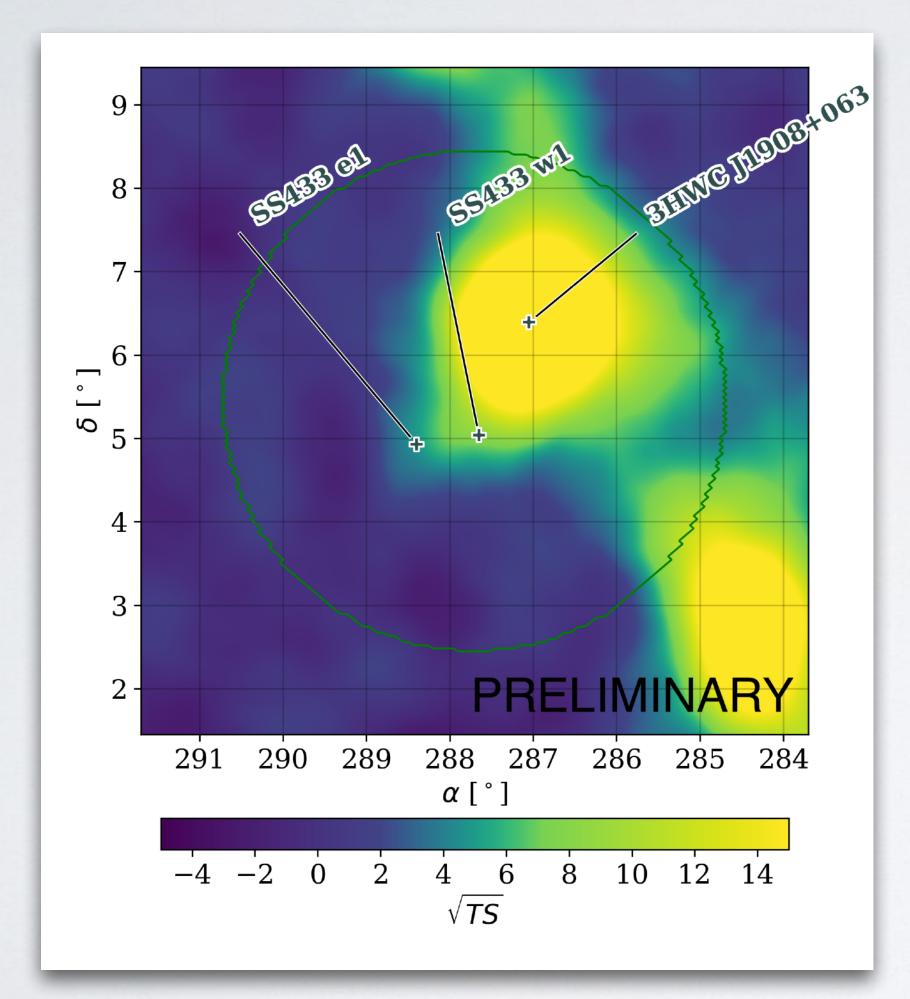
> 0.1 degrees







Significance map of the region

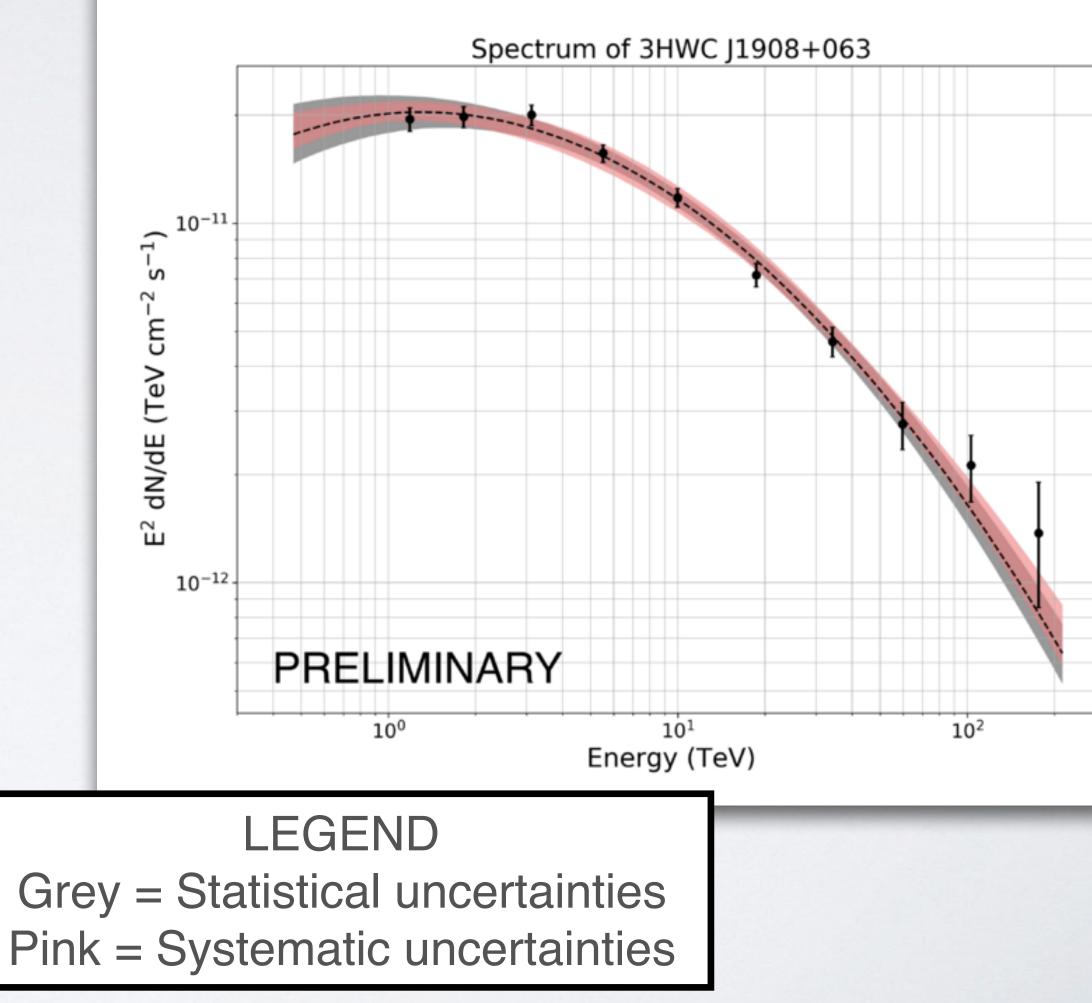


- 1343 days of data with reconstructed energies above 1 TeV
- Likelihood fit containing 3HWC J1908+063 plus the nearby lobes of SS433
- Performed using the HAWC Accelerated Likelihood (HAL) plugin to 3ML (Multimission maximum likelihood)
 - Publicly available: <u>https://github.com/</u> <u>threeML/threeML</u>



- Best-fit spectrum is a logparabola
- Best-fit morphology is a diffusion morphology with particles continuously injected from the center of the source
- Spectrum extends past 200 TeV

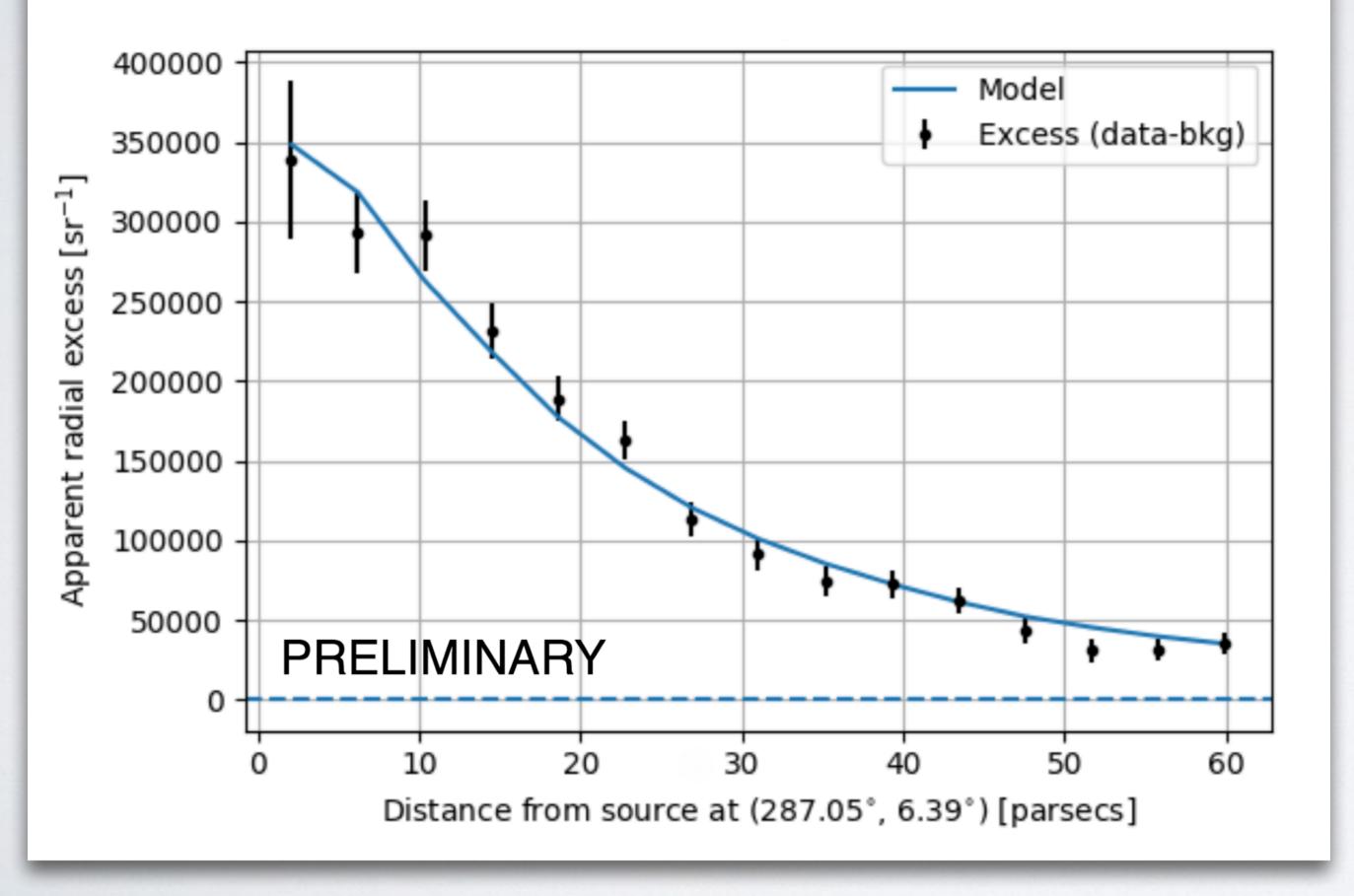
HAWC spectrum



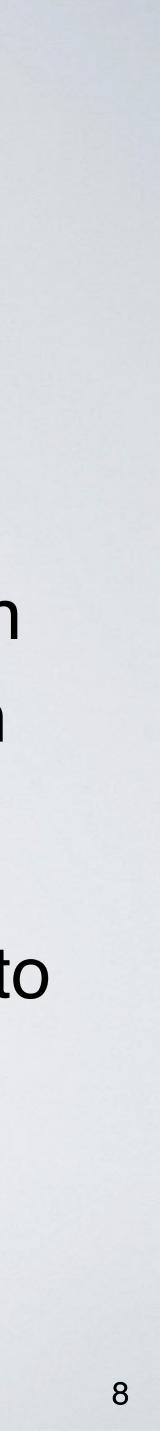


Radial profile

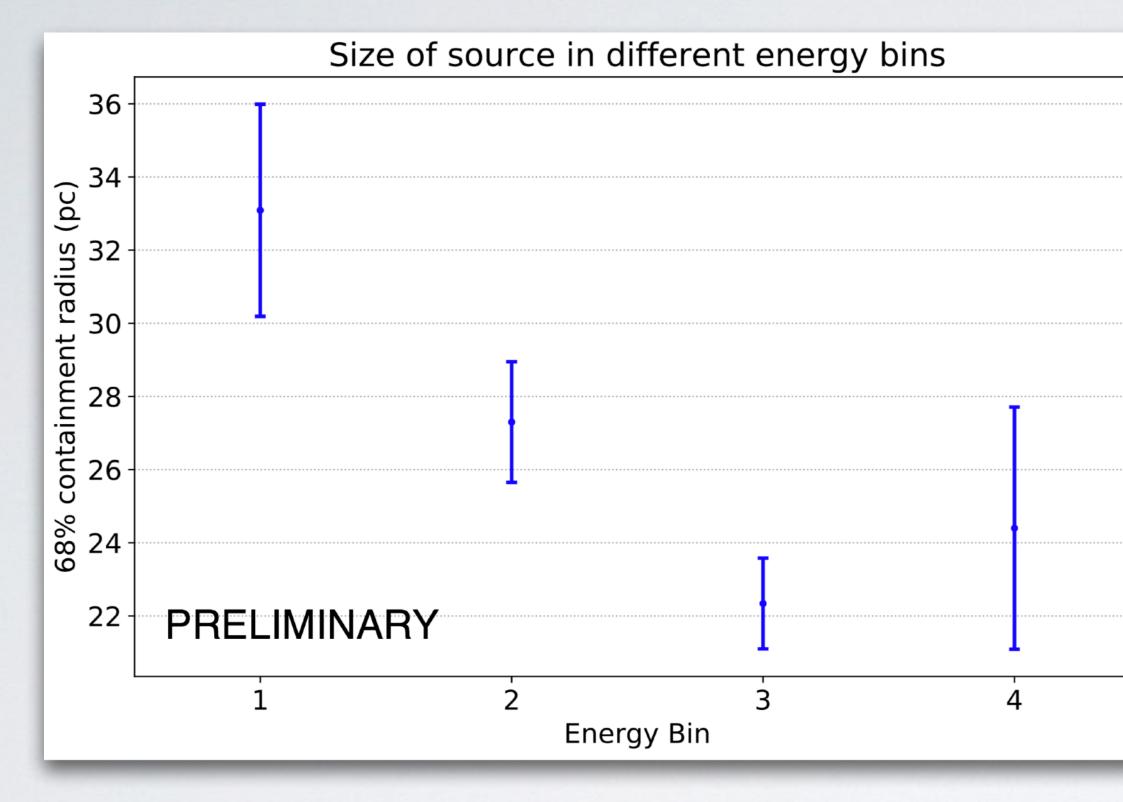
Radial profile of 3HWC J1908+063



- The morphology of the source matches
 expectations for diffusion
 with continuous injection
 from the central point.
- The source is assumed to be 2.37 kpc away (the distance to PSR J1907+0602)



The size of the source decreases with energy



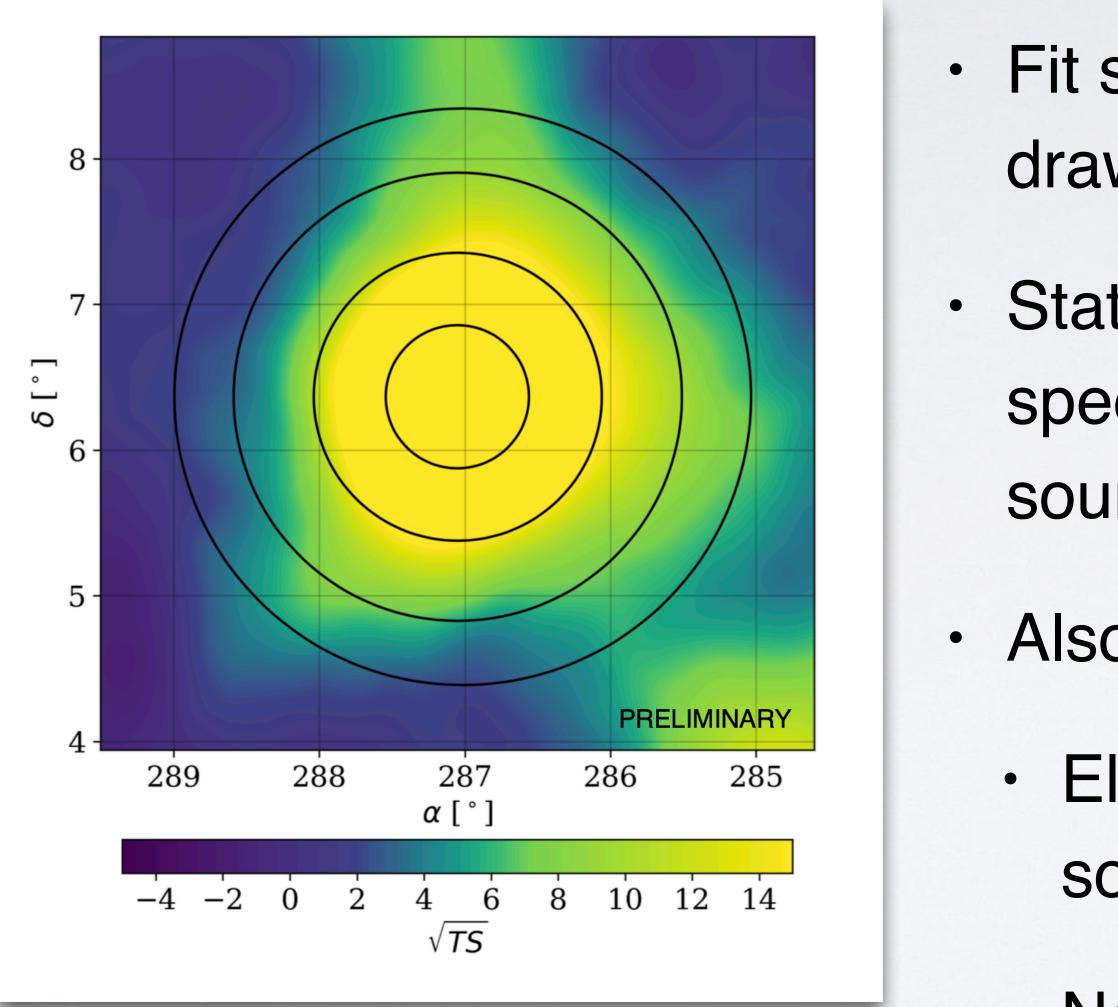
Gaussian width shrinks as energy increases

Indicative of leptonic emission

Energy bin	Minimum reco energy (TeV)	
1	0.5	1.7
 2	1.7	10
3	10	56
4	56	316



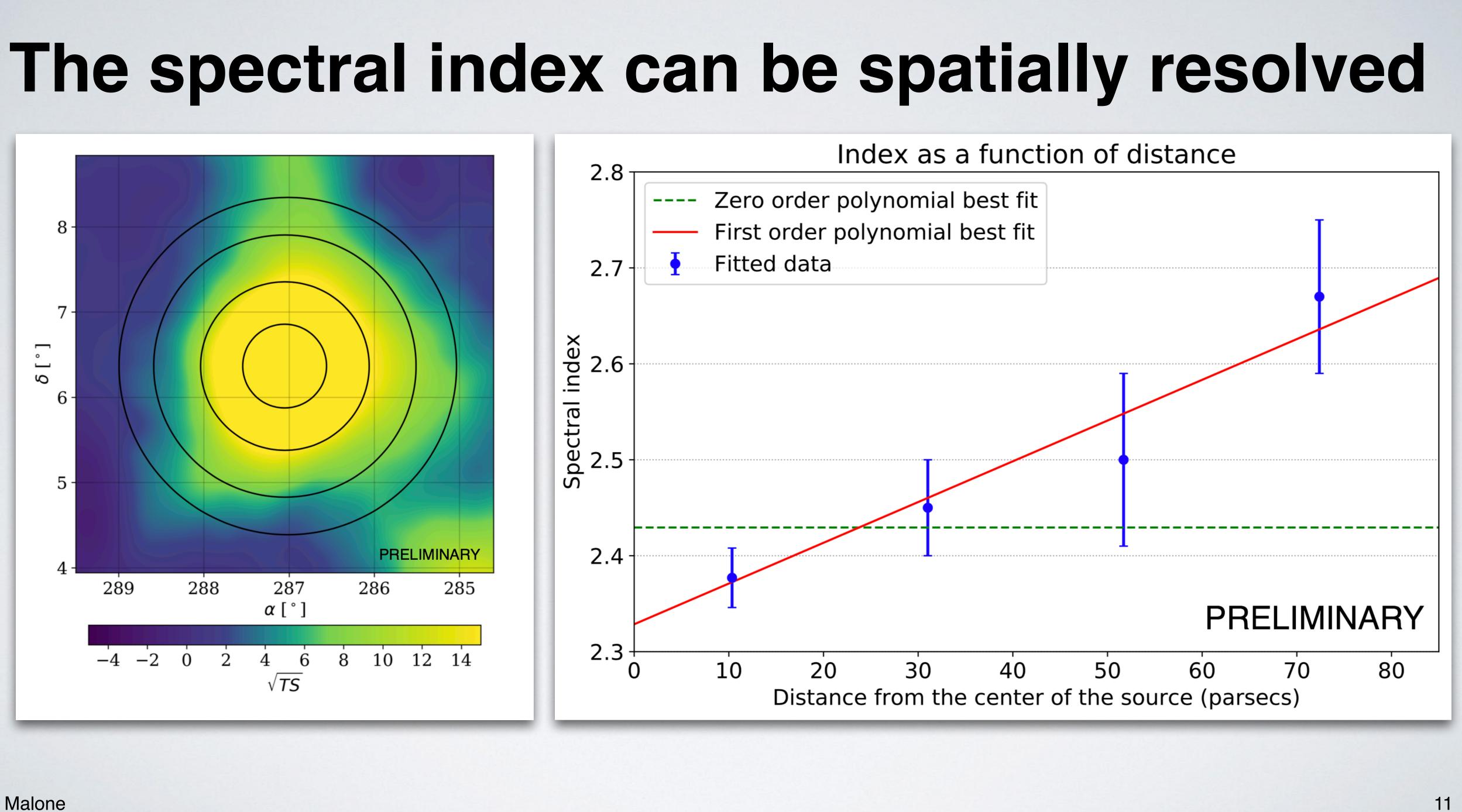
The spectral index can be spatially resolved



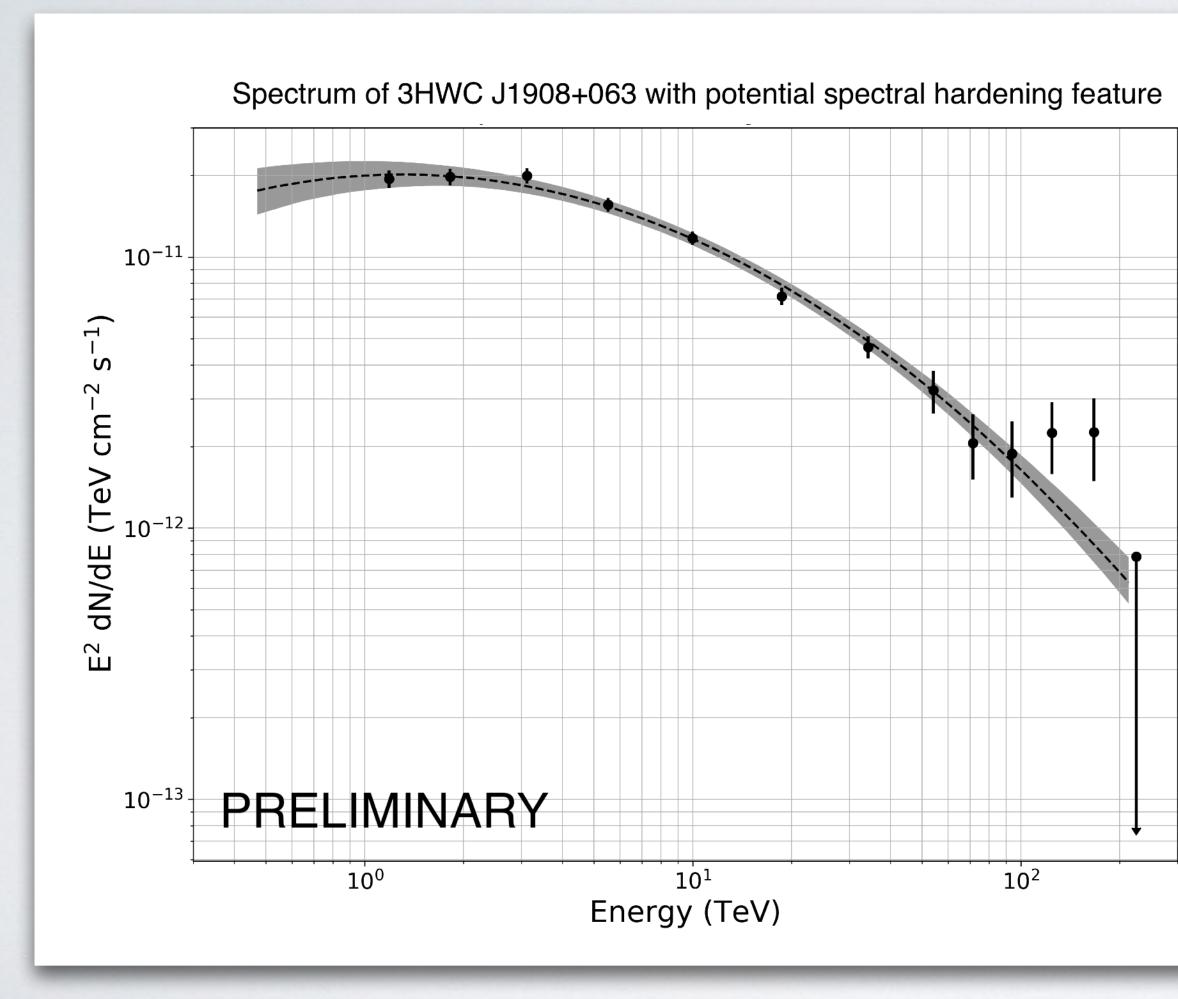
- Fit spectral index in four concentric rings drawn around the center of the source
- Statistically significant softening of the spectrum as distance from the center of the source increases
- Also indicative of leptonic emission
 - Electrons far away from the center of the source are older and have cooled.
 - No spectral change expected for hadrons





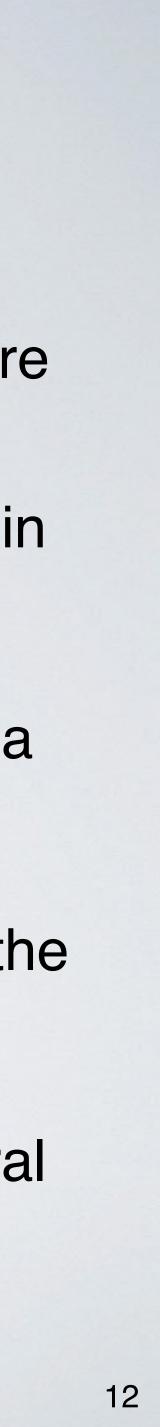


Potential spectral hardening feature



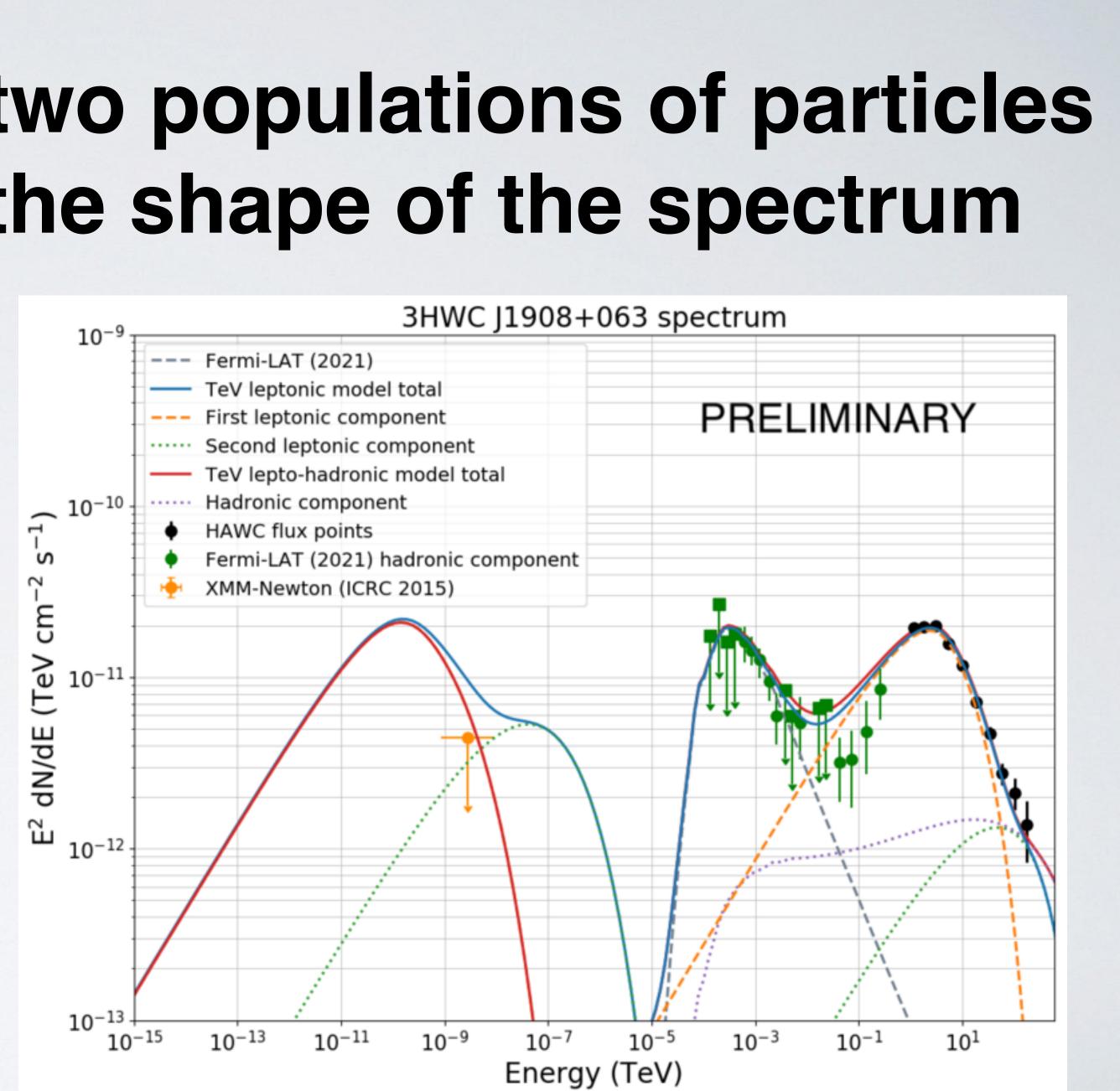
K. Malone

- When the last three energy bins are subdivided into six smaller bins of equal size, an apparent flattening in the spectrum can be seen by eye.
- Deviation from best-fit log-parabola
 by ~2 sigma.
- Second population of particles at the highest energies?
- See poster on searches for spectral hardening in HAWC sources



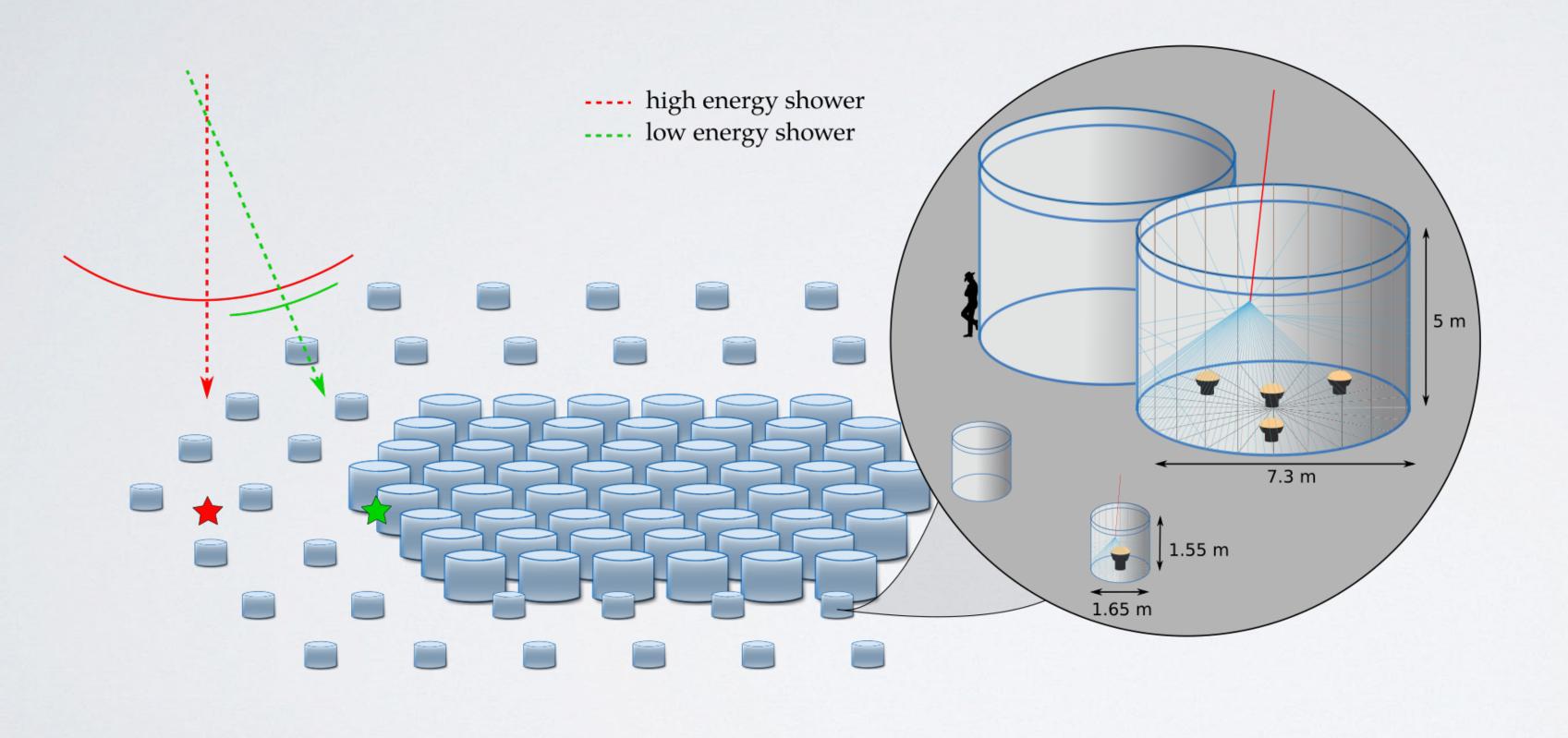
Modeling indicates that two populations of particles are needed to explain the shape of the spectrum

- HAWC's uncertainties are currently too large to determine if this second component is leptonic or hadronic in origin.
- Interesting target for upcoming gamma-ray experiments with better sensitivity at the highest energies
- Implications for detections by multimessenger experiments

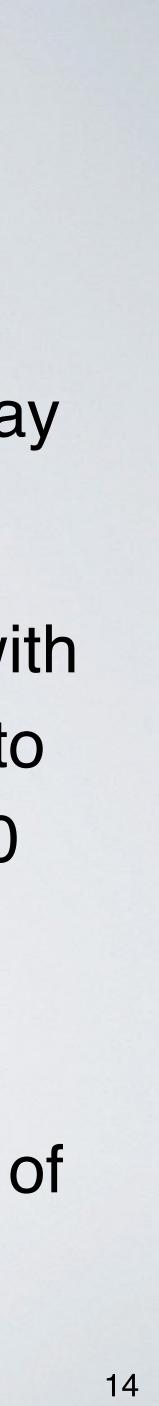




Future outlook



- HAWC's outrigger array is now operational
- Will soon have data with at least 2x sensitivity to gamma rays above 10 TeV
- Will allow for more detailed observations of 3HWC J1908+063



Conclusions

- 3HWC J1908+063 is one of the highest energy gamma-ray sources ever detected
- Modeling indicates that two populations of particles are needed to fit this source.
 - The first component is leptonic and is responsible for most of the emission.
 The second component may be either hadronic or leptonic.
 - This has implications for detection by multi-messenger experiments.

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