



Introduction

Magnetic storms represent major signatures of variability in the Sun-Earth interaction and can severely impact infrastructures at the ground level and in space, also posing a hazard to human health. On 20th August 2018, a large-scale filament gradually erupted from a quiet region of the Sun into an interplanetary coronal mass ejection that affected the Earth's environment a few days later, starting **on** late 25th August 2018 and giving rise to the third largest geomagnetic storm of Solar Cycle 24 [1]. The magnetospheric disturbance was strong enough to trigger a response in the High-Energy Particle Detector (HEPD-01) on board the China Seismo-Electromagnetic Satellite (CSES-01) [2].

The HEPD-01 detector



Fig.1: Schematic of the HEPD-01 detector.

The HEPD-01 (Fig.1) is made up of a silicon **tracking** system; a trigger system that includes one plastic scintillator layer segmented into six paddles; a range calorimeter comprising a tower of 16 plastic scintillator planes, a matrix of 3×3 LYSO (lutetium-yttrium oxyorthosilicate) scintillator crystals, and an **anti-coincidence (VETO) system** equipped with 5 plastic scintillator planes, out of which 4 are placed at the lateral sides of the apparatus and 1 at the bottom (for further details, see [3]).

The High-Energy Particle Detector (HEPD-01) as a space weather monitoring instrument on board the CSES-01 satellite

Francesco Palma^{1,2} on behalf of the LIMADOU-HEPD Collaboration

¹INFN-Sezione di Roma Tor Vergata, V. della Ricerca Scientifica 1, I-00133 Rome, Italy ²At ASI Space Science Data Center (SSDC) also, V. del Politecnico, I-00133 Rome, Italy E-mail: francesco.palma@roma2.infn.it

> HEPD-01 response to the August 2018 storm

> Fig.2 illustrates a comparison between the HEPD-01 count rate maps before (20th-23rd August; upper panel) and after the impact of the storm $(25^{\text{th}}-27^{\text{th}} \text{ August; lower panel}).$



Fig.2: HEPD-01 trigger rate map before (upper panel) and after (lower panel) the impact of the storm.

In the bottom panel, an increase in the count rate is evident at both northern and southern latitudes—especially in the southern region—as a consequence of the storm's arrival. The increased particle rate, during the storm time, is also visible as a function of the L-shell and time in Fig.3.





The enhancement of HEPD-01 trigger rates at L-shells > 3 during the storm's recovery phase suggested a phenomenon of acceleration of energetic electrons, which lasted several days.

Conclusions

Considering the sky-rocketing focus on space weather studies in this last decade, HEPD-01's results prove promising, especially in view of the already-planned constellation of CSES satellites in the next few years (CSES-02 is currently under construction).

References

- [1] F. Palma and et al., *Appl. Sci.*, vol. 11, p. 5680, 2021.
- [2] X. Shen and et al., Sci China Tech Sci, vol. 61, no. 5, p. 634, 2018.

Fig.3: Top three panels: Trigger rates for three different HEPD-01 configurations over the period August-September 2018. Bottom panel: Time evolution of the Disturbance storm-time index.

[3] P. Picozza and et al., Astrophy. J. Suppl., vol. 243, no. 1, 16, p. 16, Jul. 2019.