



LUDWIG-MAXIMILIANS-UNIVERSITÄT MÜNCHEN

# Decelerated sub-relativistic material with energy Injection

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#### **Motivation**

- ➤ Kilonovae (KNe) are produced due the coalescence of binary compact objects (NS-NS or BH-NS).
- ➤ Short duration gamma-ray bursts (sGRBs) are also linked with this phenomenon.

. Kilonovae may be studied from the formalism of a GRB

#### **Model: Considerations**

- ➤ Afterglow consisting of two components:
  - ) One produced by the non-relativistic ejecta mass  $(\Gamma \sim 1)$  . Energy distributed as a power law  $E \propto (\Gamma \beta)^{-\alpha}$  . Tan, J. C., Matzner, C. D. and McKee, C. F., 2001.
  - The other taking into account energy injection  $E \propto \tilde{E} t^{1-q}$  . Zhang, B., et al., 2006.
- Non-relativistic, adiabatic evolution of the forward shock described by the Sedov-Taylor solution.

  Sironi, L. and Giannios, D., 2013.

Evolution of the shock front in a stratified medium  $(n = A_k R^{-k})$ . Electrons of the medium cool by synchrotron process.

Dai, Z. G. and Lu, T., 1998.

#### Model

The velocity of the shock front and blast wave radius for both components may be expressed by the same equations:

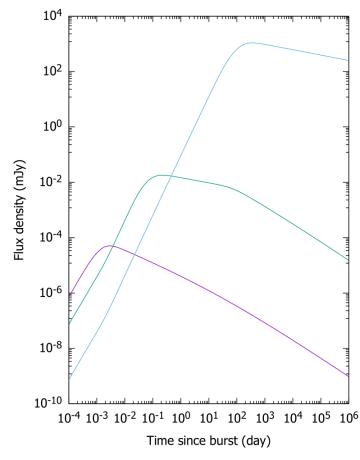
$$\beta = \beta^0 \left(\frac{1+z}{1.022}\right)^{-\frac{k-3}{\alpha+5-k}} A_{\mathbf{k}}^{-\frac{1}{\alpha+5-k}} \tilde{E}^{\frac{1}{\alpha+5-k}} t^{\frac{k-(q+2)}{\alpha+5-k}}$$

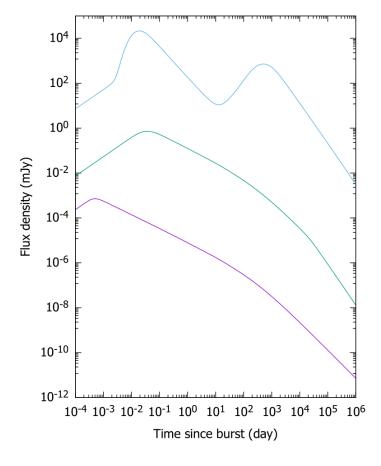
$$r = r^0 \left(\frac{1+z}{1.022}\right)^{-\frac{\alpha+2}{\alpha+5-k}} A_{\mathbf{k}}^{-\frac{1}{\alpha+5-k}} \tilde{E}^{\frac{1}{\alpha+5-k}} t^{\frac{\alpha+3-q}{\alpha+5-k}}$$

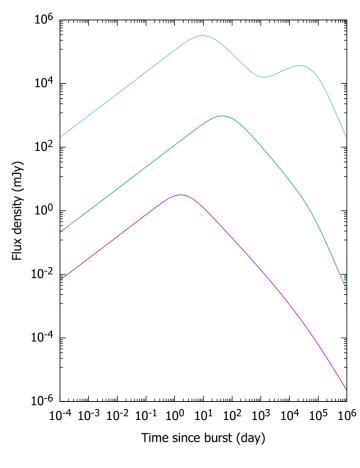
- For the energy injection component:  $\alpha = 0$
- For the component without energy injection: q=1

Fraija, N., Betancourt Kamenetskaia, B., et al., 2021.

### Results







Synchrotron light curves

- ➤ Purple X-ray (1 keV).
- ➤ Green optical (1 eV).
- ➤ Blue radio (1.6 GHz).

Left and middle panels: no energy injection.

- Left: ISM (k = 0).
- $\triangleright$  Right: Wind-like (k=2).

Right panel:

Energy injection in wind-like medium.

## Conclusion and summary

- ➤ A model to describe the afterglow emission of non-relativistic ejecta has been derived.
- >It considers a stratified medium  $n \propto R^{-k}$  and a phase with energy injection  $E \propto \tilde{E}t^{1-q}$  .
- Synchrotron light curves in the fast- and slow-cooling regimes have been obtained.