

Problematic connection between UHECRs and LLGRBs

Filip Samuelsson, Damien Bégué, Felix Ryde, Asaf Pe'er, Kohta Murase Samuelsson et al. (2019) ApJ, 876:93, Samuelsson et al. (2020) ApJ, 902:148

Idea: Use the synchrotron emission from the primary electrons as an additional messenger and compare the emission with observations of GRB 060218.

Results prompt phase:

The high magnetic field required for UHECR acceleration lead to immense optical emission from the electrons



Results afterglow phase:

The high energy budget required for observed UHECR flux lead to immense radio emission from the electrons

CTA

10

 $\log(\nu/[Hz])$

22

26

18

sensitivity V

