

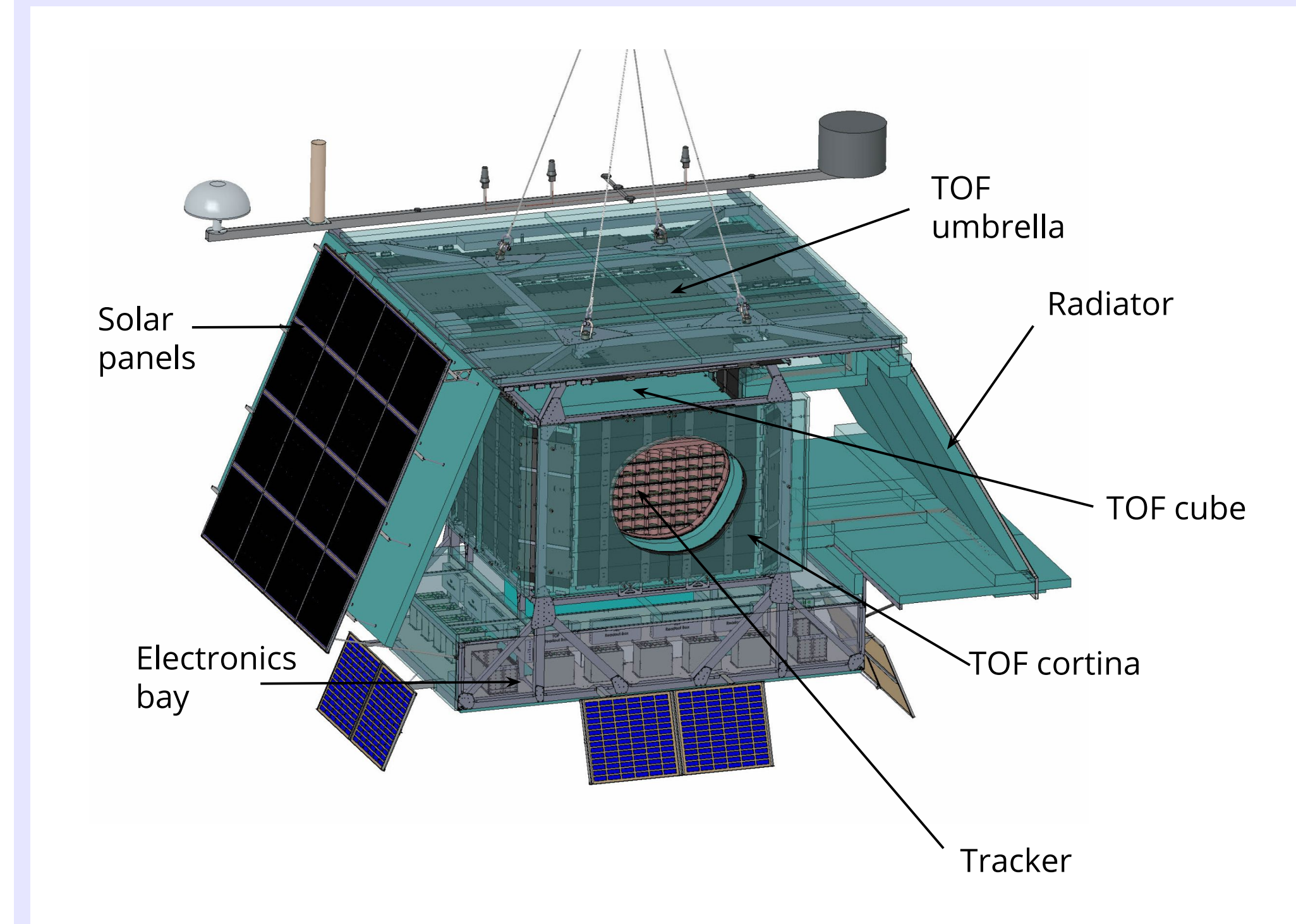
Searching for cosmic antihelium nuclei with the GAPS experiment

A. Stoessl

Department of Physics and Astronomy, University of Hawaii at Manoa, 2505 Correa Rd, Honolulu, HI 96822, USA.



The GAPS experiment



Two subsystems:

TOF time of flight system, plastic scintillator paddles, β measurement, trigger
Si(Li) tracker 1000 detectors in 10 planes, energy resolution at 4 keV for 20 – 100 keV

Event selection

Likelihood technique for particle identification. Details in [1]. Analysis in 3 stages:

Preselection Ensure that the reconstructed stopping/annihilation vertex is inside the tracker volume and there are enough hits

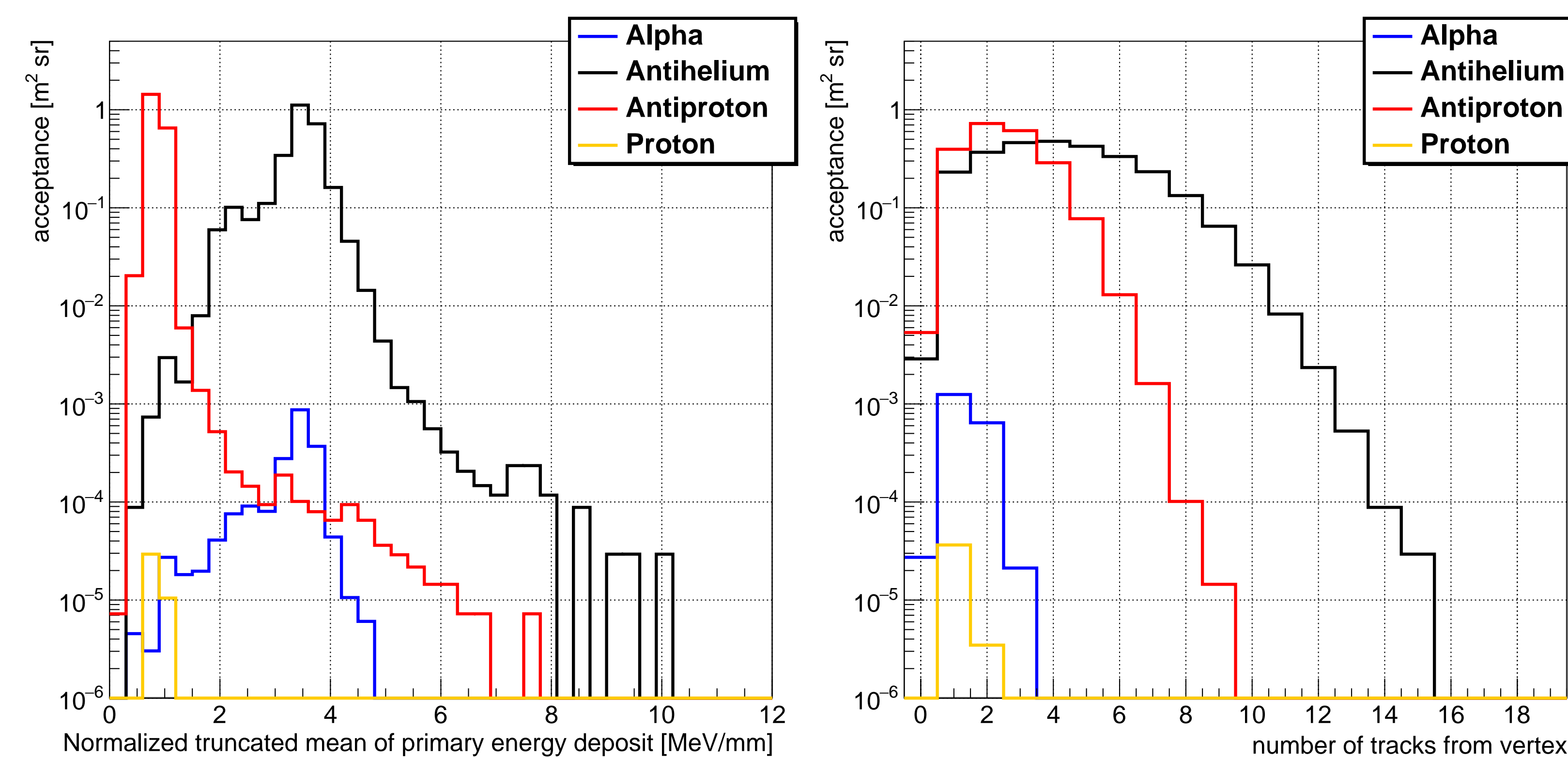
Llh construction Use seven variables to construct 2d probability distributions together with the reconstructed velocity.

Final cuts Cuts on the calculated likelihood ratio optimized individually for three $\cos(\theta)$ bins together with a cut on the mean truncated energy to ensure the reconstruction is compatible with $|Z| = 2$ and a cut on reconstructed $0.3 < \beta < 0.6$

References

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Variables



Truncated energy mean - separates primary particles with charge $|Z| = 1$ and $|Z| = 2$

Number of tracks from vertex - identifies antiparticles (annihilation exhibits "star" pattern)

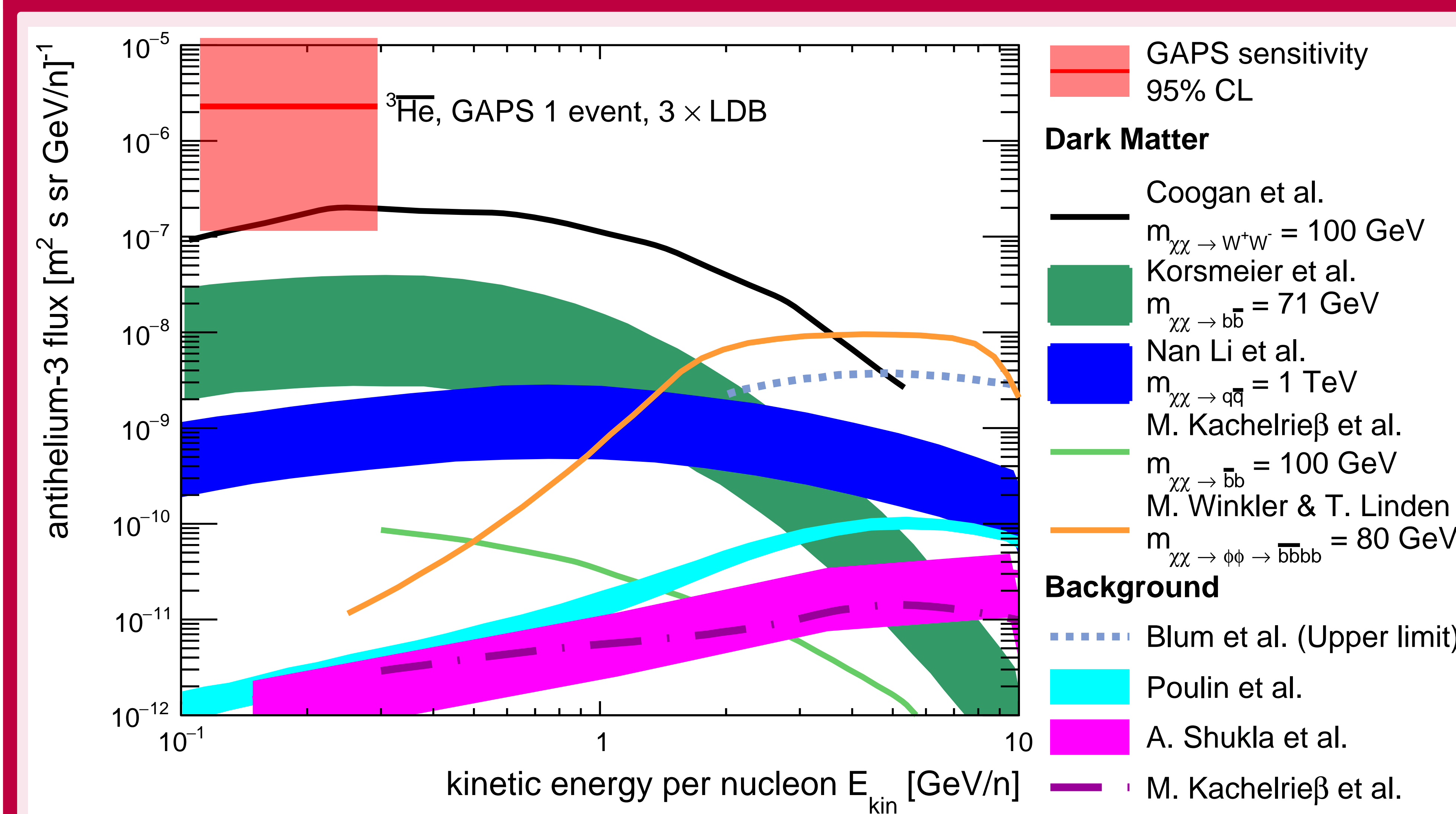
Simulation & Reconstruction

- Simulation using Geant4.10.7p02
- FTFP_BERT_HP physics list
- Detailed model of GAPS geometry
- 10^{11} protons, $4 \cdot 10^9$ α -particles and $7 \cdot 10^8$ antiprotons generated for this study.
- Reconstruction of primary track + annihilation star specifically developed for GAPS [2].

Sensitivity

- Bayesian approach as in [11]
- Sensitivity for one detected event
- Sensitivity calculated for three 35days LDB flights
- Remaining background on the order of 10^{-3}
- $2.29 \cdot 10^{-6} \text{ m}^{-2} \text{ sr}^{-1} \text{ s}^{-1} (\text{GeV}/n)^{-1}$

GAPS anti He3 sensitivity



GAPS antihelium-3 sensitivity for three LDB flights of 35 days each (95% confidence level).
 Antihelium-3 flux predicted from dark matter taken from [3, 4, 5, 6, 7], standard astrophysical background predictions taken from [8, 9, 10].

$$S = \frac{n - b}{A_{id} T \Delta E \epsilon_{geo} \epsilon_s}$$

one expected event detected $\rightarrow n - b$ (number of background events)
 atmospheric survival probability $\rightarrow \epsilon_s$
 geomagnetic cutoff efficiency $\rightarrow \epsilon_{geo}$
 kinetic energy range 0.11 - 0.33 GeV/n $\rightarrow \Delta E$
 observation time for 3 flights (105 days) $\rightarrow T$
 average antihelium-3 acceptance after cuts $\rightarrow A_{id}$

Summary & Outlook

- Initial MC study of GAPS antihelium-3 capabilities
- LLh analysis exploiting primary event characteristics and annihilation star
- Unprecedented sensitivity to antihelium-3 flux in an energy region inaccessible to AMS-02

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