

BL Lac object 1ES 0647+250, a decade of MWL observations

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on behalf of the MAGIC and *Fermi*-LAT Collaborations and MWL partners

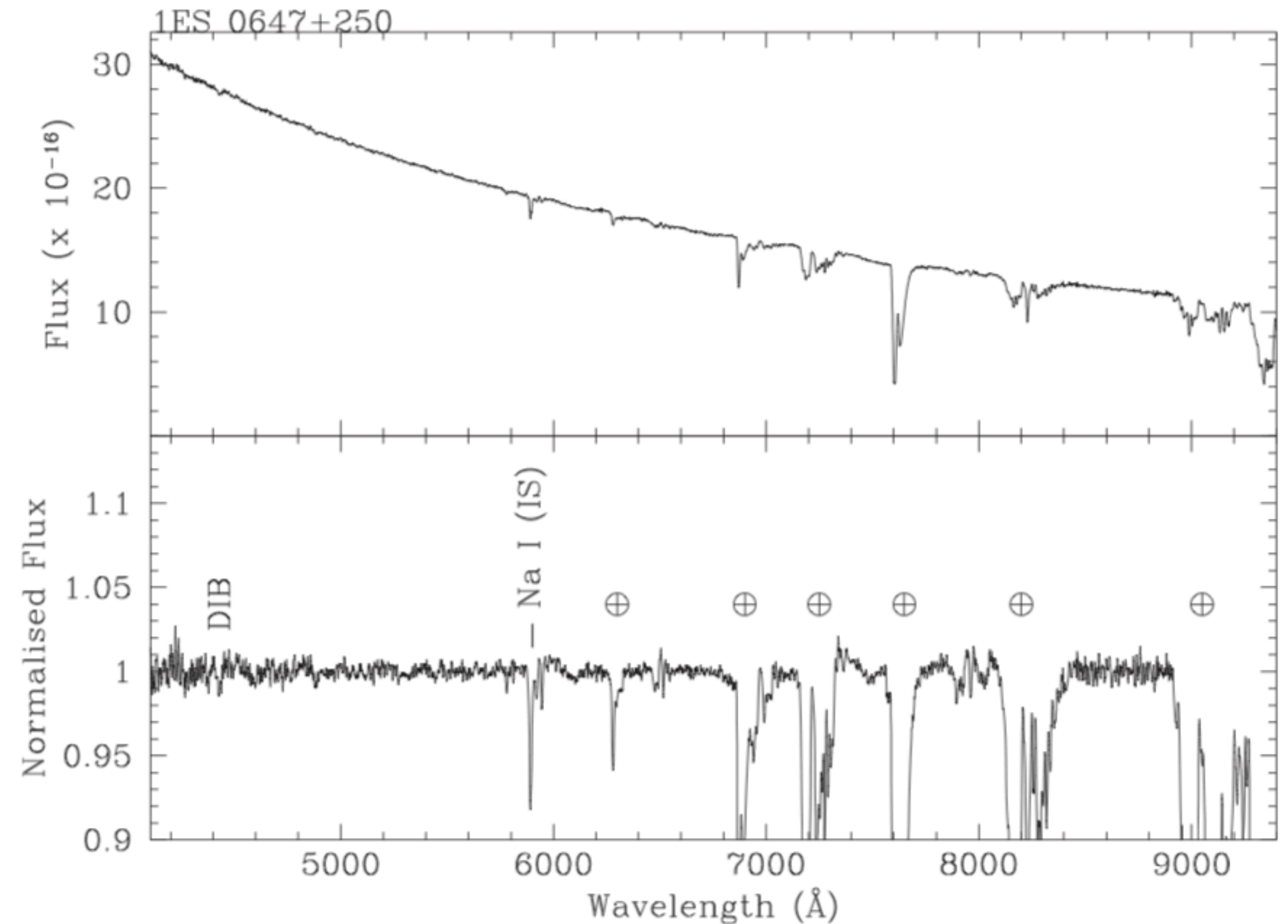


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1ES 0647+250

- High synchrotron peak (HBL) BL Lac object
- Unknown redshift
 - Lower limit of $z > 0.29$
 - Tentative value of $z = 0.41$
- Detected by MAGIC in 2009-2011 in non-flaring state
- Detected in VHE γ -rays in several enhanced states (2014 after optical high state, 2019 and 2020 after X-ray high states)
- 11 years of MWL data collected



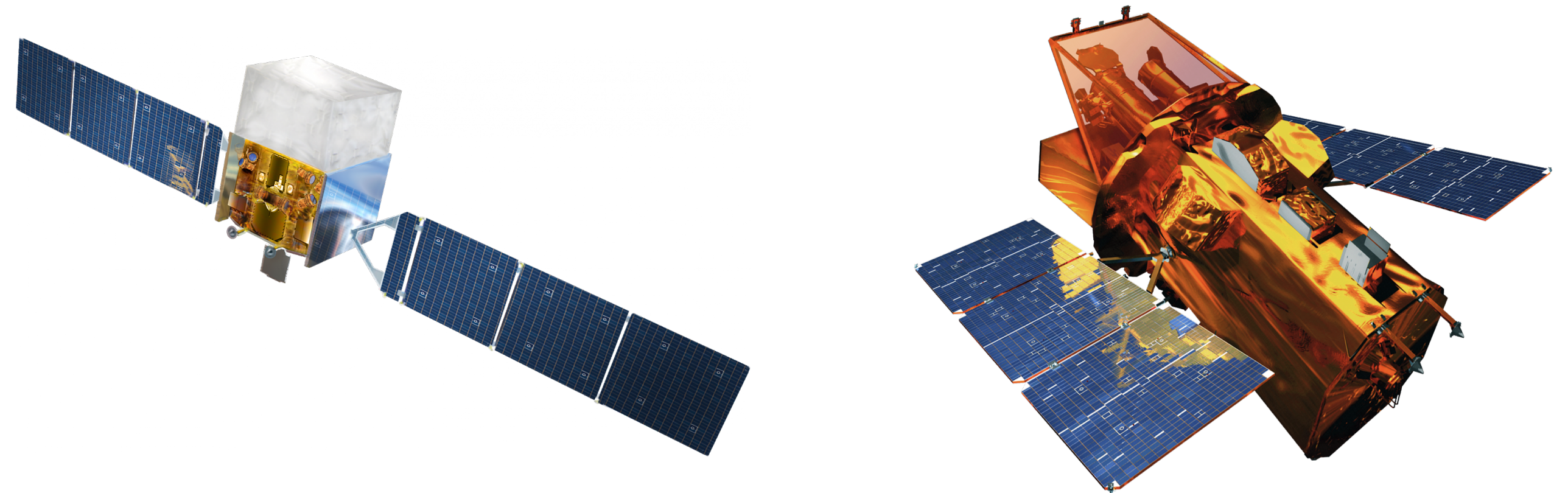
Optical spectrum of 1ES 0647+250 presented by Paiano et al. (2017) for constraining the redshift lower limit of the source.

MAGIC and MWL data set

- MAGIC data from 2009 to 2020
- MWL data from several facilities from HE γ -rays to radio



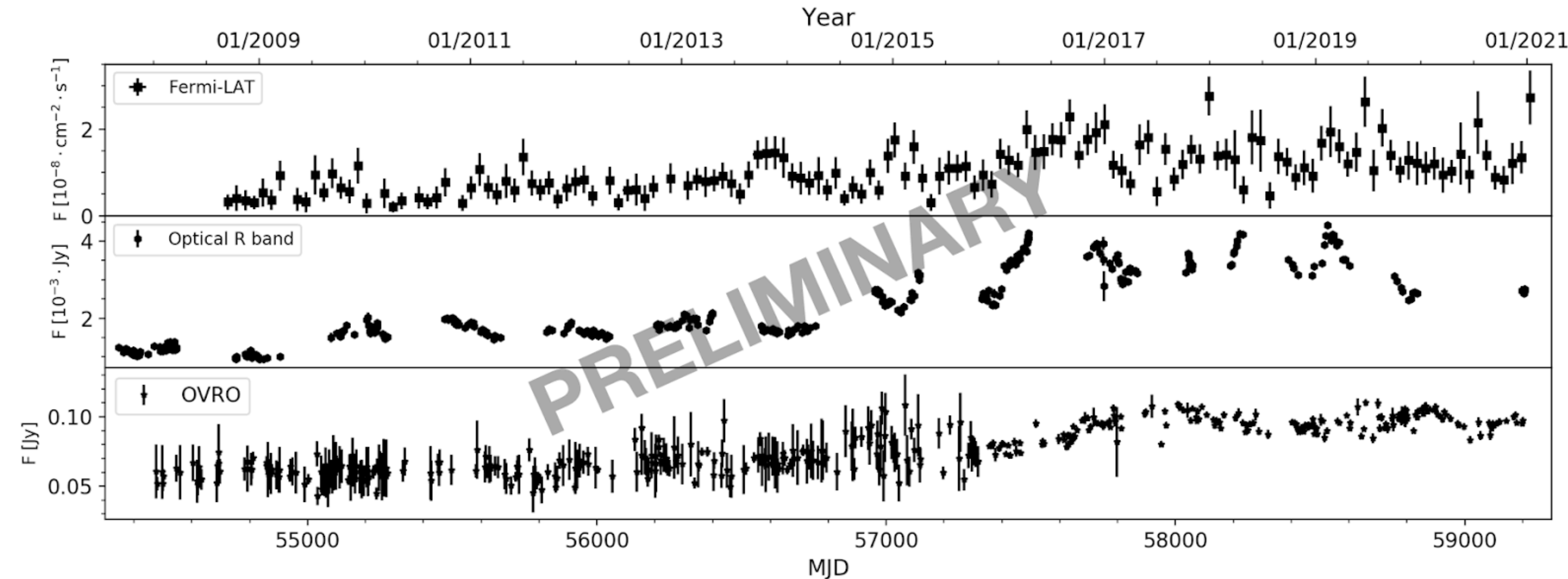
MAGIC Telescopes at the Roque de los Muchachos Observatory, La Palma, Spain. (Credit: Daniel López/IAC).



Ground- and space-based facilities and instruments used for this work. *Top*: optical telescopes (Tuorla-KVA, Las Cumbres, PIRATE and Liverpool Telescope) and radio telescope (OVRO) used. *Bottom left*: Fermi satellite. *Bottom right*: Swift satellite (Credit: NASA).

Results

MWL light curves
11-year LCs for long-term analysis



HE γ -rays (*Fermi*-LAT), optical R-band (KVA, Las Cumbres, PIRATE and LT) and 15 GHz radio (OVRO) light curves.

Spectral characterization

	$F(E > 100 \text{ GeV})$ [% Crab Units]	Spectral parameters	
		E_0 [GeV]	Spec. index
2009-2011	2.0 ± 0.5	190	$\alpha = -3.1 \pm 0.4$
2014	3.4 ± 1.6	100	$\alpha = -3.3 \pm 0.7$
2019	8.0 ± 1.8	100	$\alpha = -3.7 \pm 0.6$
2020	15.0 ± 1.0	100	$\alpha = -3.2 \pm 0.2$ $\beta = -1.9 \pm 0.7$

Integral flux above 100 GeV and spectral parameters of 1ES 0647+250 for the different detected periods.

Redshift estimation:

Prandini et al. (2010) empirical method + upper limit with joint MAGIC + *Fermi*-LAT spectrum

Estimated redshift, z_{est}	z_{UL} [95% C.L.]
0.45 ± 0.05	0.81

Results of the redshift estimation. z_{est} is the estimated value through the empirical relation from Prandini et al. (2010) and z_{UL} the 95% C.L. upper limit with the joint HE and VHE spectra.

