Radio Simulations of Upgoing Extensive Air Showers **Observed from Low-Earth Orbit** Andrés Romero-Wolf* on behalf of the NuSpaceSim Collaboration Yosui Akaike¹, Luis Anchordoqui², Douglas Bergman³, Isaac Buckland³, Austin Cummings⁴, Johannes Eser⁵, Claire Guépin⁶, John F. Krizmanic^{7,8,9}, Simon Mackovjak¹⁰, Angela Olinto⁵, Thomas Paul², Sameer Patel¹¹, Alex Reustle^{9, 12}, Mary Hall Reno¹¹, Fred Sarazin¹³, Tonia Venters⁹, Lawrence Wiencke¹³, Stephanie

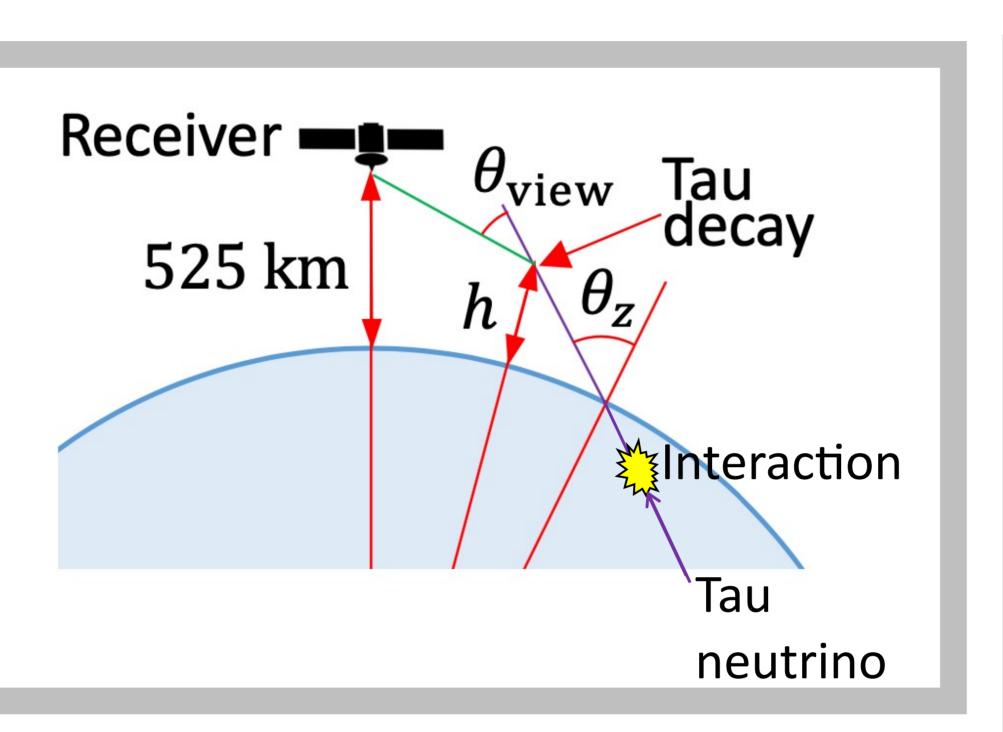
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Motivation

nuSpaceSim:

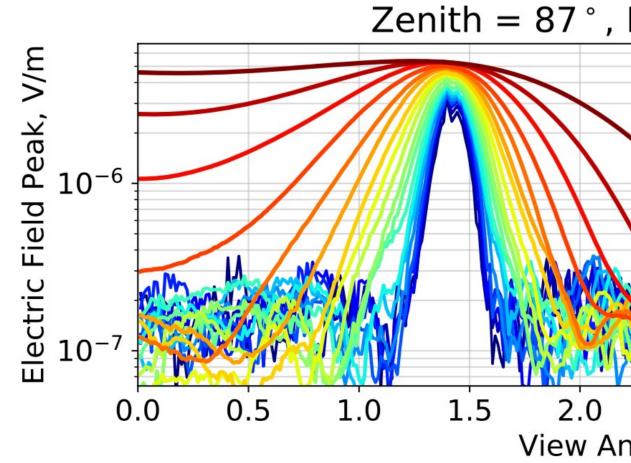
- Upgoing air showers sourced by tau neutrino interactions in Earth
- Detailed simulation development.
- NASA funded effort
- Includes optical and radio emission.

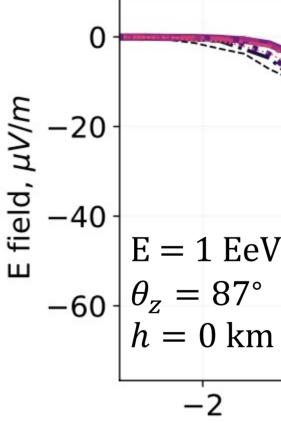
ZHAireS to model radio emission observed from low-Earth orbit.



Simulation Setup

- Apply to 525 km altitude observatory.
- AIRES particle shower
- ZHS algorithm with 0.3 ns sampling.
- Zenith angles: 50° 85° (in 5° steps), and 87°, 89°.
- Decay altitudes: 0-9 km in 1 km steps.
- In all cases $\vec{B} \perp$ shower axis.
- Time-domain pulse data processed to frequency domain.
- Produce frequency banded radio beam patterns.





Wissel⁴ in additional collaboration with Jaime Alvarez-Muñiz¹⁴, Washington Carvalho¹⁵, Harm Schoorlemmer¹⁶, and Enrique Zas¹⁴

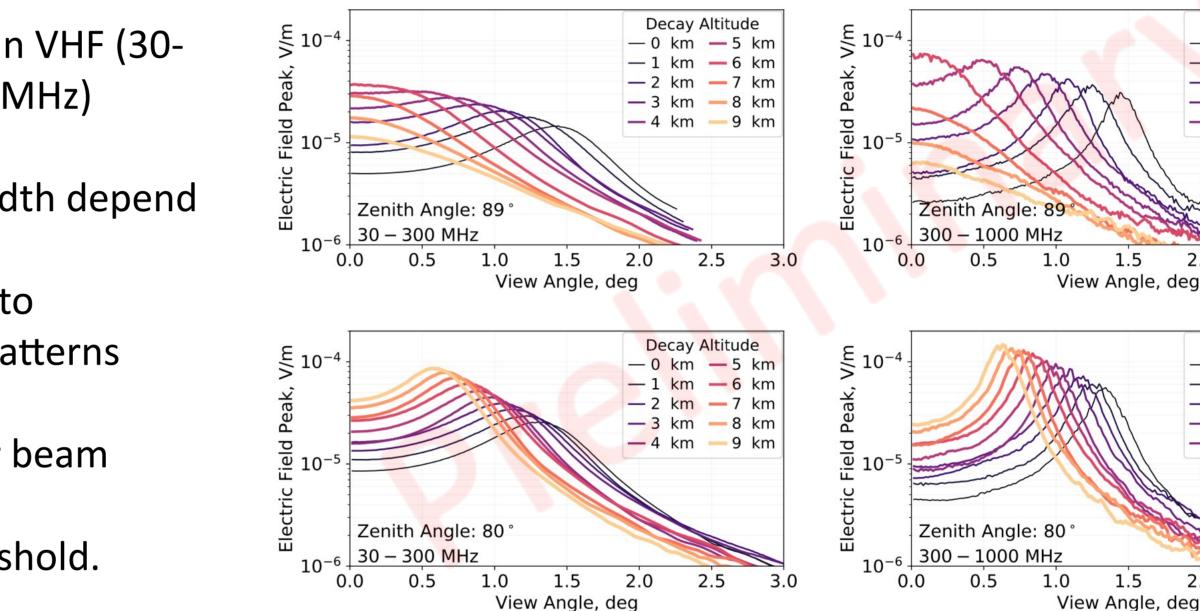
----- $\theta_{view} = 1.1^{\circ}$ $-\cdots$ $\theta_{view} = 1.2^{\circ}$ ••••• $\theta_{\text{view}} = 1.3^{\circ}$ $\theta_{\text{view}} = 1.4^{\circ}$ -- $\theta_{view} = 1.6$ $- - \theta_{view} = 1.7^{\circ}$ $\theta_{\rm view} = 1.8$ ° time, ns Zenith = 87° , Decay alt. 0 km 500 - 550 MHz 450 - 500 MHz ----- 950 - 1000 MHz 400 - 450 MHz 900 - 950 MHz 350 - 400 MH 850 - 900 MH 300 - 350 MHz 250 - 300 MHz - 200 - 250 MHz - 150 - 200 MHz 100 - 150 MH; 550 - 600 MHz 50 - 100 MHz 2.0 2.5 3.0 3.5 4.5 4.0 View Angle, deg

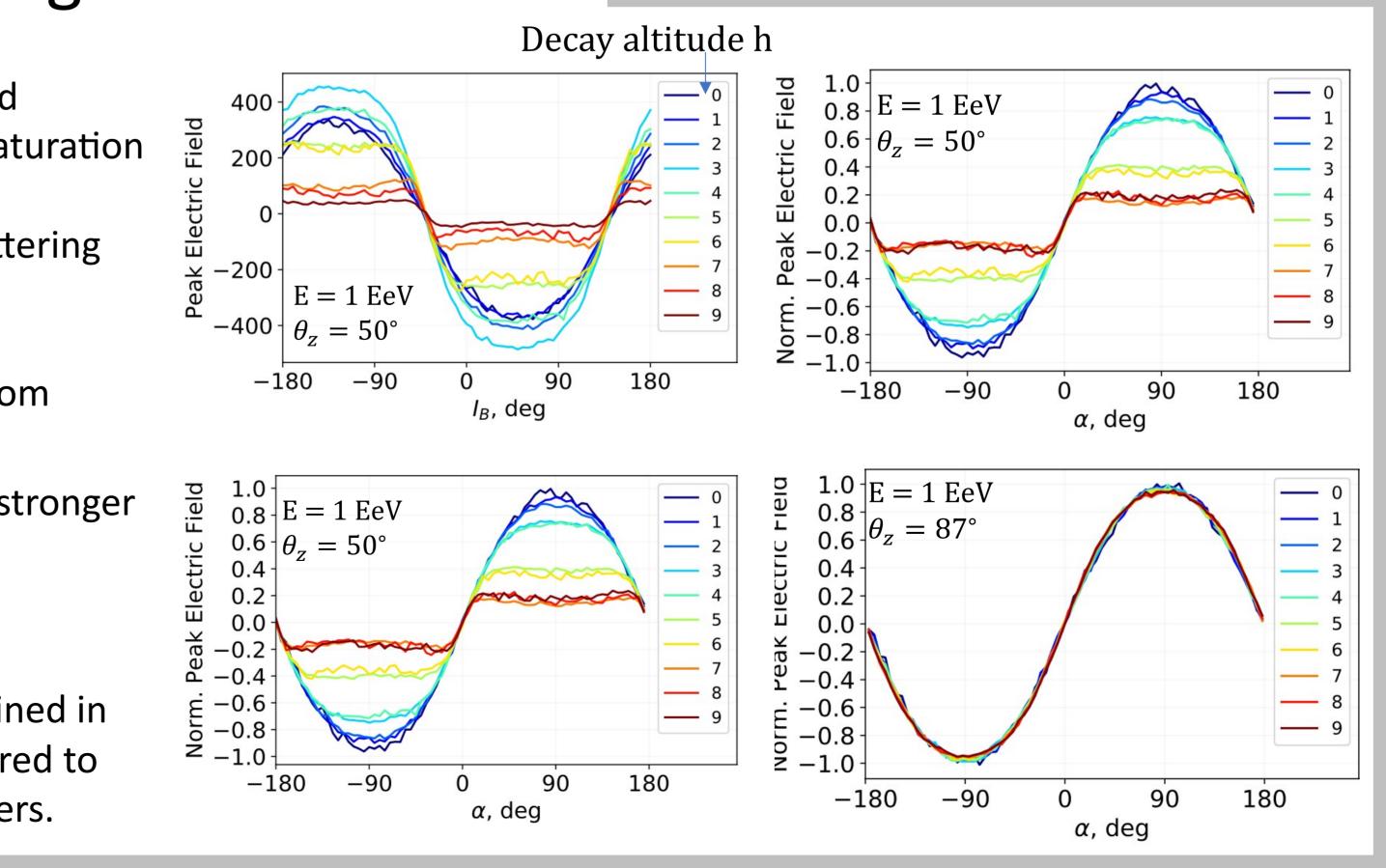
Results 1: Radio Emission

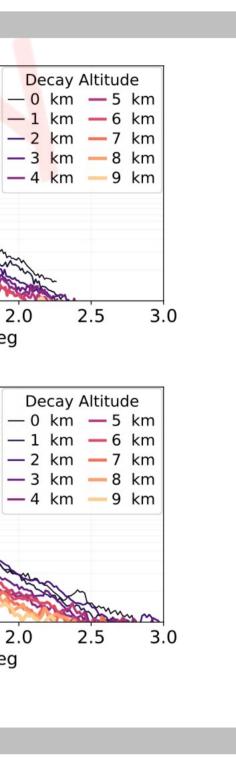
- Processed electric field peaks in VHF (30-300 MHz) and UHF (300-1000 MHz) bands.
- Cherenkov angle and beam width depend on band.
- VHF: weaker signal compared to background but wider beam patterns
 - Good for extreme energies
- UHF stronger signals narrower beam patterns
 - Good for lower energy threshold.

Results 2: Geomagnetic Saturation

- Variation of geomagnetic field inclination angle I_B reveals saturation effect at high altitudes.
- Potentially due to longer scattering interaction lengths in rarified atmosphere.
- Saturation effect deviates from $|\vec{E}| \propto \sin \alpha$.
- Saturation clipping effect is stronger at higher altitudes.
- Larger zenith angles restore $|\vec{E}| \propto \sin \alpha$.
- Shower development contained in less rarified medium compared to more steeply upgoing showers.







Conclusions

- 1. First results of radio emission from detailed track-level upgoing air shower simulations as observed form low-Earth altitude.
- 2. Results share similar features with simulations at lower altitudes for ANITA and **BEACON**.
- 3. As-of-yet unexplained behavior for nearly horizontal showers currently under investigation.
- 4. Main new finding is the saturation of geomagnetic radio emission for steep upgoing shower initiated at high altitude.