





Performance of the ASTRI Mini-Array at the *Observatorio del Teide*

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OVERVIEW

- The ASTRI Mini-Array (ASTRI MA) is a next-generation ground-based gamma-ray observatory under construction at the Observatorio del Teide (Tenerife, Spain). It will be composed of nine small-sized (~4 m in diameter) and largefield-of-view (~10°) Cherenkov telescopes operating in the energy range from 1 to 100 TeV and beyond. Each telescope is characterized by a dual-mirror optical system and a silicon photo-multiplier camera.
- In order to assess the performance of the ASTRI MA at the Teide Observatory site detailed Monte Carlo simulations were generated (by means of the CORSIKA and sim telarray packages) and subsequently reduced with A-SciSoft, the scientific software package of the ASTRI Project.
- Thanks to its performance, the ASTRI MA will allow unprecedented observations of the galactic and extragalactic sky at the TeV and multi-TeV energies, in synergy with present- and next-generation gamma-ray observatories in the Northern Hemisphere.





PERFORMANCE





- Differential sensitivity better than present-generation IACTs above a few TeV.
- Differential sensitivity comparable to CTAO Northern Array (in construction phase configuration, dubbed Alpha) above a few tens of TeV.
- Differential sensitivity superseded by HAWC (507 days) and LHAASO (1 year) above a few tens of TeV (but at those energies substantially better angular/energy resolution).

→ ASTRI MA fully functional complement at the TeV energies to present- and nextgeneration gamma-ray observatories in the Northern Hemisphere.









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PERFORMANCE





- Angular resolution of a few arcmin above a few TeV \rightarrow important for morphological studies of sources detected by HAWC and LHAASO in the multi-TeV regime.
- Energy resolution of ~10% above a few TeV → important for spectral studies at TeV and multi-TeV regime.
 - Performance up to $3^{\circ}(5^{\circ})$ from the center of the field of view within a factor of $\sim 1.5(\sim 2)$ equal to the nominal on-axis performance \rightarrow important for simultaneous multi-target observations, large surveys of the sky, and possible serendipitous discoveries.







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