Particle-in-Cell Simulations of Synchrotron Maser Emission and Associated Particle Acceleration in Relativistic Shocks

Abstract

We study synchrotron maser emission and associated particle acceleration at electron-ion relaivistic shocks by PIC simulations. Relativistic shocks emit intense electromagnetic waves from the shock front via the synchrotron maser instability. This synchrotron maser emission induces electrostatic plasma waves (i.e., wakefield) and transverse filamentary structures in the upstream region via the stimulated/induced Raman scattering and filamentation instability, respectively. Some lucky particles in the phase of the deceleration by the wakefield enter the filamentary structures. Then they are reflected by the wakefield and decoupled from the upstream bulk flow. These decoupled particles can feel the motional electric field and efficiently accelerated.

1. Introduction

- Relativistic shocks may generate ultra-high-energy cosmic rays (> 10^{18} eV) \rightarrow the detailed acceleration mechanism is still unknown ✓ Relativistic shocks can emit coherent
- electromagnetic waves via the synchrotron maser instability [1]
- ✓ 1D PIC simulations show that the intense coherent emission induces the electrostatic waves (i.e., wakefield) via the stimulated/induced Raman scattering and that the efficient particle acceleration occurs in the upstream [2,3]
- ✓ Our high-resolution 2D PIC simulations [4] shows that the same acceleration indeed work even in realistic multidimensional systems.



However.....

How do particles are accelerated? How much is the maximum energy?

2. Simulation Setting



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- transverse filamentary structures.
- \checkmark The nonthermal ions and electrons are generated in the upstream.

- Coupling through Wakefield in Relativistic Shocks, ApJL 883 (2019) L35.

$$\gamma_{max,e} \sim \gamma_1^3$$
, $\gamma_{max,i} \sim \left(1 + \frac{m_e}{m_i}\gamma_1\right)\gamma_1^2$

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