

array were performed.

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The next generation of the IceCube Neutrino Observatory, IceCube-Gen2, with its surface array, will be a unique detector to study cosmic rays in the energy region where extragalactic sources are expected to begin to dominate. Increased aperture, in particular for coincident events, will be beneficial not only for air shower studies but will also for vetoing atmospheric background for astrophysical neutrino events.

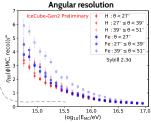
> A library of air showers was prepared and processed through the detector arrays using CORSIKA/CoREAS and Geant4 within the IceCube software.

Reconstruction algorithms were optimised for the scintillator array. The

signal and timing information is included in a likelihood minimization in

which the shower geometry, energy estimator, LDF and shower front

See F. Schröder (407)



scint trigge 3 antennas (E_{EM} reco) 5 antennas (E_{EN}+X_{max} reco) year 10, In the second se ^یا 10 Aperture: 9.839km² s ⊨ 10⁴ 9 10² 14.0 14.5 15.0 15.5 16.0 16.5 17.0 17.5 18.0 log₁₀(E/eV)

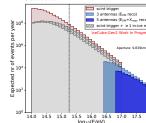
Sufficient number of events for scintillator. probing Galactic-to-extragalactic transition.

Simulations will be continued beyond 1 EeV





parameters are obtained. Scintillators provide sub-degree resolution above 10 PeV. Initial studies on the energy estimator for the scintillator Radio reconstruction is in preparation and will further improve energy See A.Coleman (317) Statistics



radio, and coincident in-ice detection for





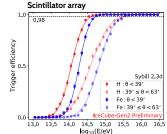


estimation and will deliver Xmax estimation.





Hybrid station with 8 scintillators and 3 antennas above each in-ice string.



Trigger efficiency for at least 5 hits. The threshold reaches sub-PeV for zenith angles up to 39° and increases with the increasing zenith angle.

References and further readings:

[1] IceCube-Gen2 Collaboration, A. Karle and M. Kowalski, PoS(ICRC2021) [2] IceCube-Gen2 Collaboration, F. Schröder, PoS(ICRC2021) 407 [3] IceCube Collaboration, A. Coleman, PoS(ICRC2021) 317 [4] IceCube Collaboration, A. Leszczyńska and M. Plum, PoS(ICRC2019) 332 (2020)

Radio antennas 1.0 0 - 18 18 - 2727 - 33 ົາຂ - 33 - 39 39 - 45 45 - 51 0.6 51-57 57 - 63 Q.4 × _{0.7} IceCube-Gen2 Preliminary 16.0 16,5 17.0 17,5 18 log₁₀(E/eV)

Condition requires 3 antennas with signal above the threshold and angular reconstruction within 5°. Radio detection is more efficient for inclined showers.



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1 km 1 km Surface array delivers infromation about cosmic-ray

10 CRI PoS (ICRC2021) 411

≈72 m

energy while the measurement of the *muonic core* in the ice improves the mass identification.

Summary & Prospects

- IceCube-Gen2, with its surface array, will be a unique detector to study cosmic rays in the PeV-EeV cosmic-ray energies.
- A hybrid air-shower array in combination with the in-ice detection of the muonic core has a potential to boost the CR energy and mass determination as well as to improve our

