LIGO



An Archival Search for Very-High-Energy Counterparts to Sub-Threshold Neutron-Star Merger Candidates ICRC 2021

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- LIGO and BNS Mergers
- EM Counterparts
 - Very-High-Energy
- Present
 - Sub-Threshold Searches in Archival VERITAS data
- Future
 - Potential to expand to:
 - Other IACTs
 - The Next-Generation: CTA



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First BNS Merger Detections - GW 170817



Adapted from <u>Abbott+ 2017</u>

http://maravelias.info/2017/12/a-short-talk-on-gw170817-at-uoc/





GRBs with IACTs

visibility

Long GRBs have been detected by current-generation IACTs: 201216C (MAGIC: t_{del}~ 57 s) 2. 180720B (HESS: t_{del}~ 10 h) 3. 190114C (MAGIC: t_{del}~ 50 s) 4. 190829A (HESS: t_{del}~ 4.3 h) H.E.S.S. visibility +180° -180 VTS/MAGIC

-90

3.

Credit: TeVCat

3 σ hint of a short GRB at VHEs:
160821B (MAGIC: t_{del} ~ 24 s)



Figure 3. Sky map showing the excess significance (standard deviation, pre-trial) as measured by MAGIC for events above ~0.8 TeV. The white cross marks the position of GRB 160821B according to Swift-XRT. The PSF corresponding to 68% containment is depicted as a white circle in the left lower corner, with radius 0.045 deg.

<u>Acciari+ 2020</u>

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Sub-threshold motivation



- LIGO had 103 sub-threshold binary neutron star (BNS) merger candidates in their first observing run (O1) (<u>Magee+ 2019</u>)
 - Given local BNS merger rate of $\mathcal{R} \approx 100 4000 \text{ Gpc}^{-3} \text{ yr}^{-1}$ at 90% confidence (nominal value of 1000 Gpc}^{-3} \text{ yr}^{-1} adopted) and estimated O1 sub-threshold search sensitivity $\langle \text{VT} \rangle = 6.7 \times 10^5 \text{ Mpc}^3 \text{ yr}$
 - Expect that $\mathcal{R} \times \langle VT \rangle = 0.67_{-0.60}^{+2.0}$ of these candidates are **real** gravitational wave signals
 - Despite contamination fraction being very high, motivates possibility to correlate real events with other messenger signals



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Coincidence Algorithm



- Identify VERITAS observations that are:
 - Spatially coincident with the 90% credible region of a sub-threshold BNS merger candidate
 - Temporally coincident with a pre-defined coincidence window around a LIGO subthreshold measurement.



Sub-threshold examples





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Results



Candidate Label	LIGO BNS Candidate Event ID	LIGO				VERITAS		
		$\overline{\text{FAR}}_{(\text{yr}^{-1})}$	S/N	$\begin{array}{c} \text{p-astro} \\ (10^{-3}) \end{array}$	$\frac{\text{Area}}{(\text{deg}^2)}$	t_{first}	t_{coinc}	Coverage Probability
C1	2015Oct12T02:40:22.39	142.27	8.42	3.82	2321	-0:11:17	0:18:53	0.22%
$C2^L$	2015Oct24T09:03:52.00	7.52	9.69	79.6	24218	1:33:08	1:11:08	0.06%
$\mathbf{C3}^H$	2015Nov17T06:34:02.07	7.52	8.84	181	24221	-0:08:02	2:37:43	0.18%
C4	2015 Dec 04 T 01:53:39.14	225.02	9.09	2.5	2909	0:16:20	1:00:00	0.19%
$C5^L$	2015 Dec 06 T06:50:38.17	77.45	7.72	6.64	24264	-0:09:02	2:10:18	0.15%
C6	2015 Dec 09 T 07:25:24.68	141.65	7.85	3.84	2606	1:36:25	0:15:00	0.03%
C7	2016 Jan 02 T 02:47:29.35	356.13	7.51	1.63	3487	1:44:55	0:30:00	0.18%

Adams+ 2021 (in press)

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Moving forward





Outlook



- Current-generation IACTs
 - ≥ 70 additional coincident VERITAS observations predicted in aLIGO O2/O3 (<u>Adams+ 2021, in press</u>)
 - Without a detection, enough coincident observations has potential to constrain VHE emission (need a LOT of observations - participation of H.E.S.S. and MAGIC would help bridge the gap)
 - **MoUs for near real-time alerts** would provide additional opportunities
 - Minimal operational burden
 - Can adjust pre-defined observing schedules to maximize overlap with candidates
- Future prospects
 - Factor of **5-to-16 improvement** in FoV area with CTA over VERITAS
 - CTA era also includes $O_4/O_5 \rightarrow$ **vastly expanded sensitive area**



Thank You