

# Resolving the origin of VHE γ-ray emission from the PeVatron candidate SNR G106.3+2.7 using MAGIC telescopes

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**T. Oka**, T. Saito, H. Kubo, M. Strzys for the MAGIC Collaboration

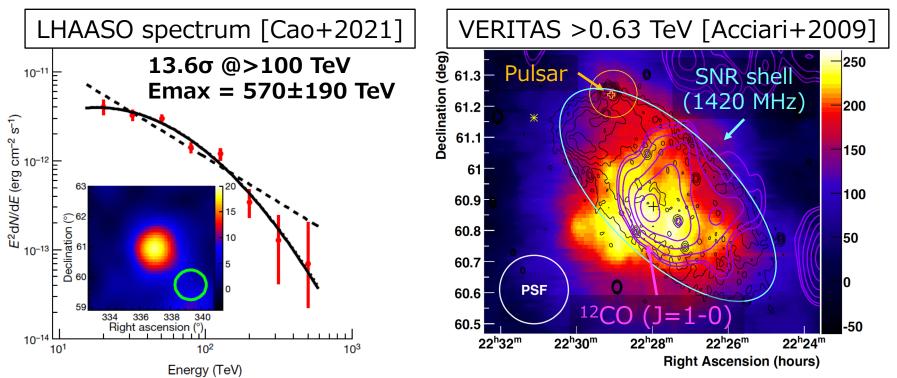
# Boomerang PWN/SNR G106.3+2.7

#### One of the PeVatron candidate

- HAWC, TibetASγ, and LHAASO detected 100 TeV γ-ray emission [Albert+ 2020; Amenomori+ 2021; Cao+2021]

#### ⊗ PWN & SNR complex (Boomerang PWN & SNR G106.3+2.7)

- age: 4-10 kyr [Halpern+2001, Kothes+2006]
- $\rightarrow$  The origin of the VHE emission is unclear. PWN or SNR? hadronic or leptonic?



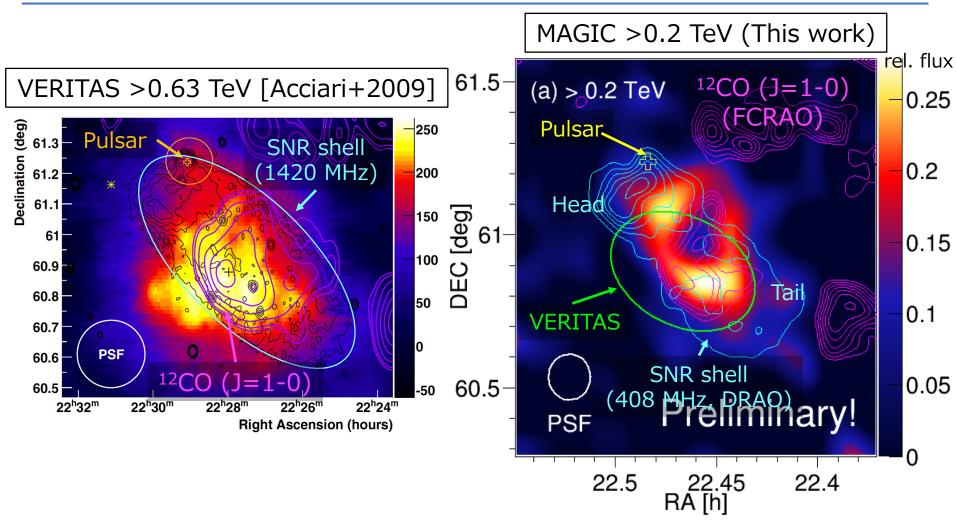
### MAGIC observations

The MAGIC observations were performed with the best angular resolution among the previous  $\gamma$  ray observations in the vicinity of SNR G106.3+2.7.



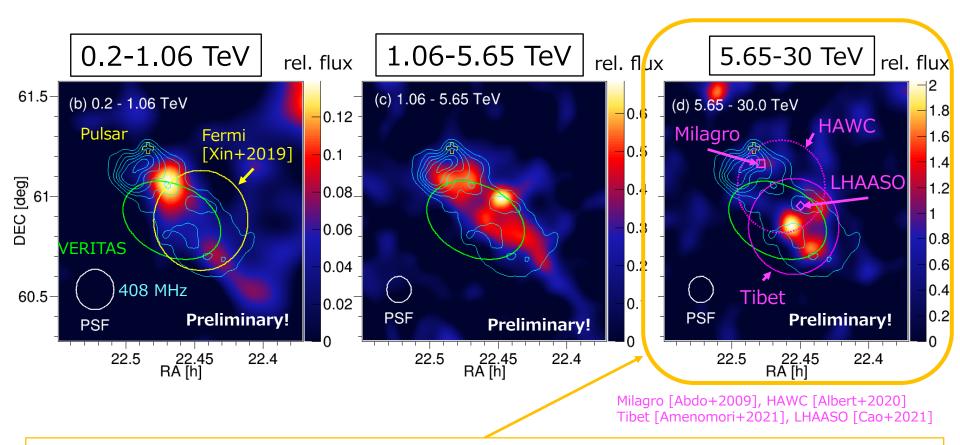
	VERITAS (pre-upgrade) [Acciari+2009]	MAGIC [This work]
Observation period	2008	2017 - 2019
Selected time	33.4 hr 🛁	6 122 hr
Analysis threshold	0.63 TeV	0.2 TeV
68%-containment radius of PSF	0.11° —	0.084° (>0.2 TeV) 0.072° (> 1 TeV)

### MAGIC Skymap



 γ-ray emission extends along with radio continuum emissions, named Head & Tail by Joncas & Higgs (1990).

### Energy-dependent morphology

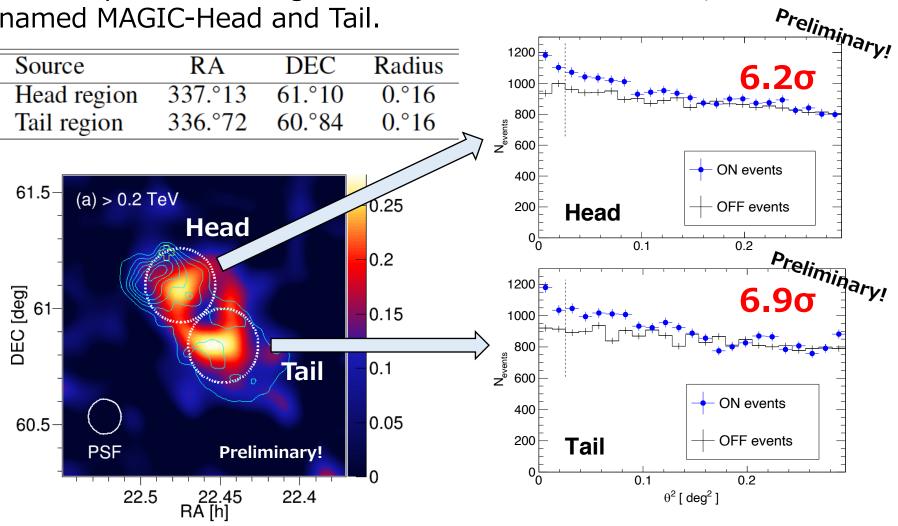


✓ In the highest energy bin,
 γ-ray emissions are offset from the pulsar position.

 $\checkmark$  The position is in good agreement with other experiments

# Two region analysis (> 0.2 TeV)

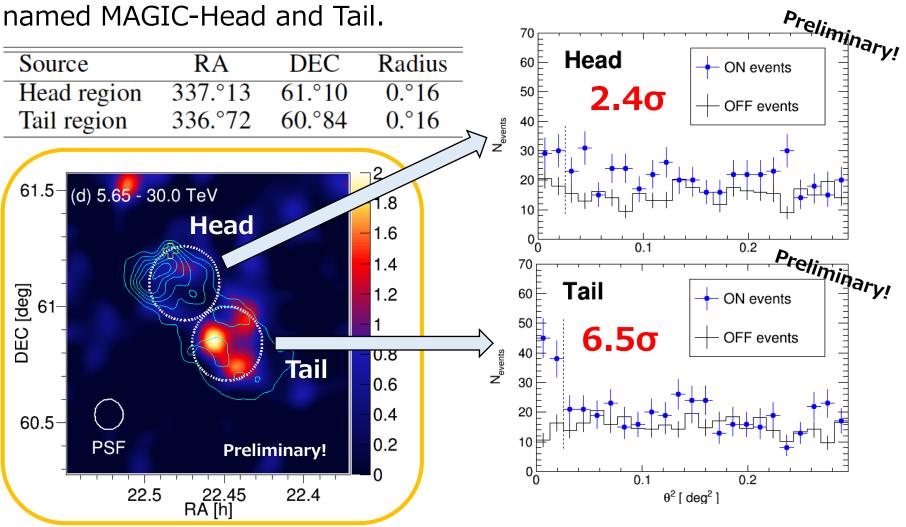
We analyze the two regions defined in the table below, named MAGIC-Head and Tail.



✓ Significant detection at both regions.

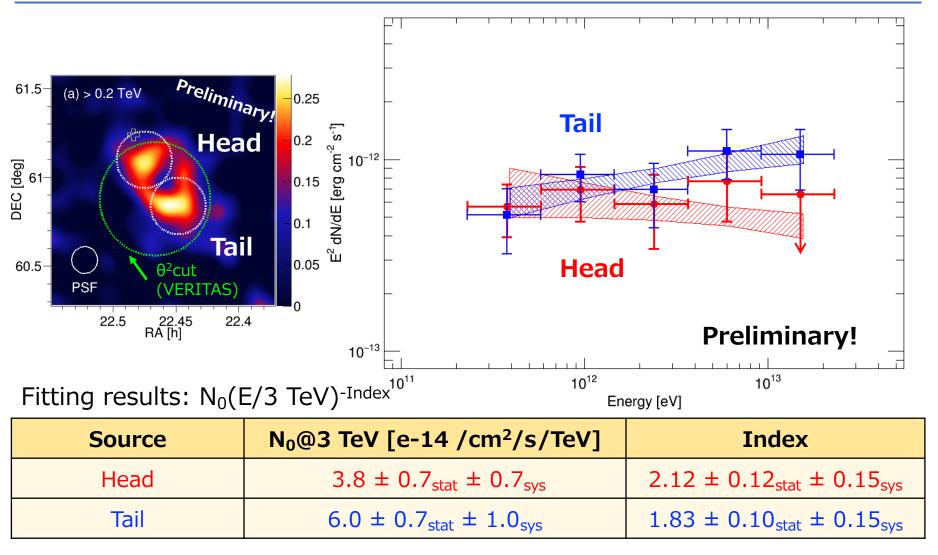
# Two region analysis (5.65-30 TeV)

We analyze the two regions defined in the table below, named MAGIC-Head and Tail.



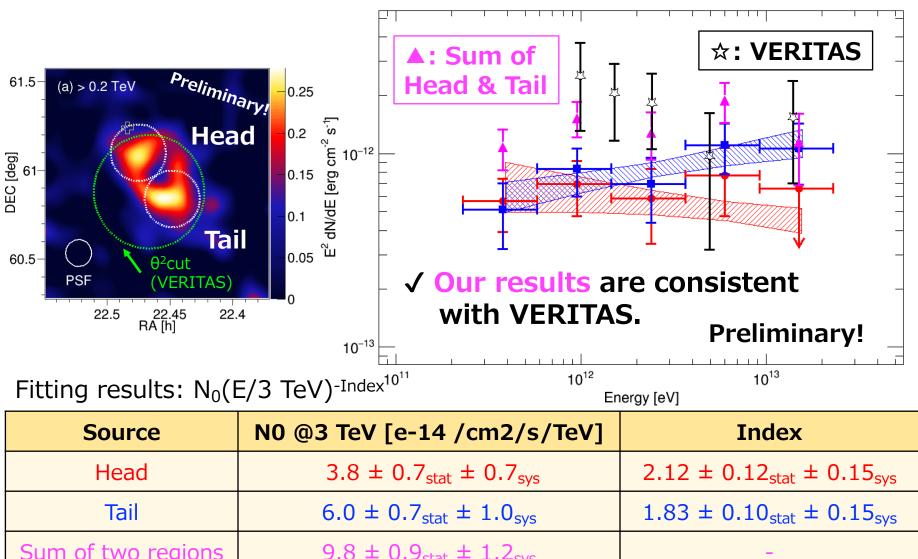
✓ Tail emission is significant, while Head is not significant

### Spectra of the Two Regions



✓ The spectra of two regions are very similar but at higher energies of ~10 TeV, Head has lower flux

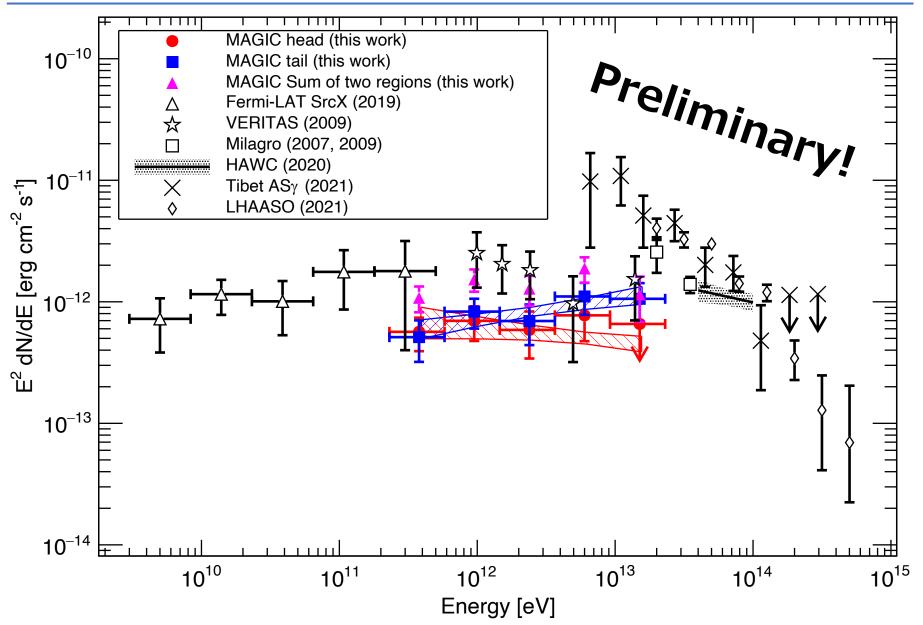
### Spectra of the Two Regions



um of two regions
 
$$9.8 \pm 0.9_{stat} \pm 1.2_{sys}$$
 -

 VERITAS [2009]
  $11.5 \pm 2.7_{stat} \pm 3.5_{sys}$ 
 $2.3 \pm 0.33_{stat} \pm 0.30_{sys}$ 

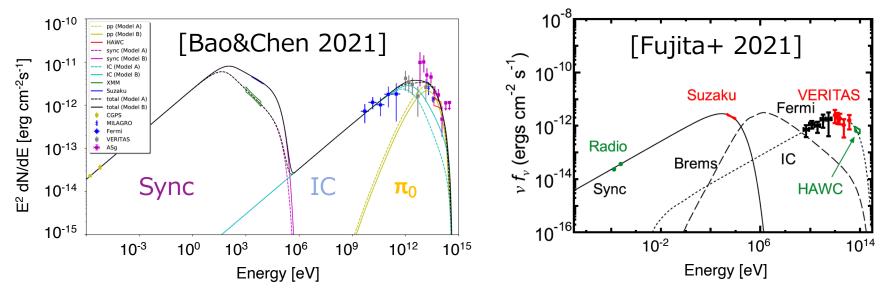
### Comparison with other experiments



### Interpretation of the VHE emission

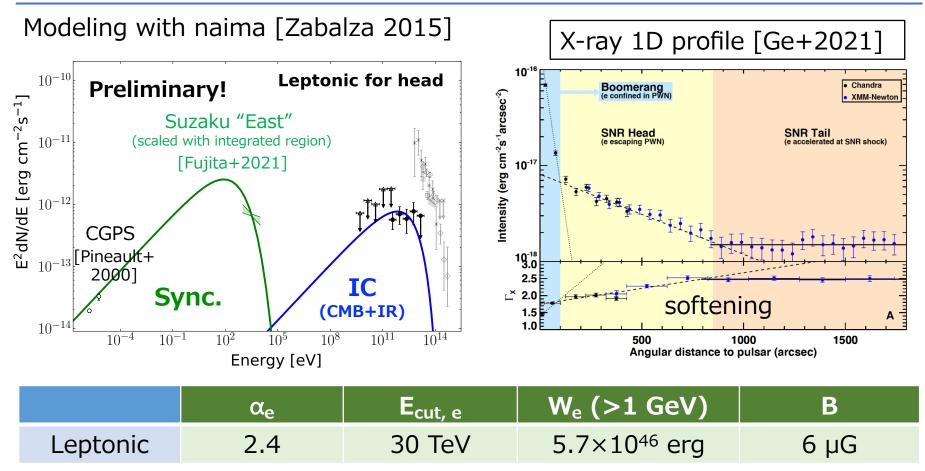
#### Case 1: the origin of Head and Tail are same

- ✓ The extension along with the shell suggests that the gamma-ray emissions originate in the SNR.
- ✓ Various models have been considered for Case 1.
  - Hadronic models can reproduce well the observed spectra >10 TeV.
  - However, there is still a leptonic scenario which can explain MWL SED.



<u>Case 2: the origin of Head and Tail are different</u>  $\rightarrow$  Discuss in the following pages.

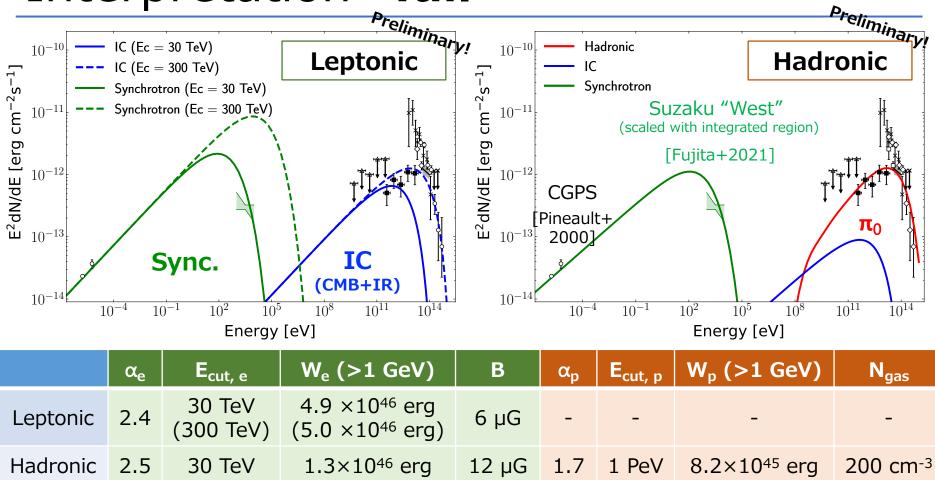
# Interpretation -Head-



✓ Leptonic emission can reproduce the MWL spectrum in the head region.

✓ X-ray results for the head region (Ge+2021, Fujita+2021) suggest the synchrotron emission originates in the Boomerang PWN.

# Interpretation -Tail-



Leptonic: Electrons of SNR-tail can reproduce the MAGIC spectrum but if assume gamma-ray > 10 TeV is only from tail, in tension.

Hadronic: Protons escaped from SNR in the past can explain PeV energies and the hard index at the middle-aged SNR (4-10 kyr)

### Summary

✓ SNR G106.3+2.7/Boomerang PWN is a PeVatron candidate.

- ✓ MAGIC has detected extended gamma-ray emission spatially coincided with the SNR radio morphology.
- ✓ At higher energies (5.65-30 TeV), MAGIC-tail emission is significant, while the head is not significant.
- ✓ MAGIC spectra at both regions could be reproduced with IC emission. Head: electrons escaped from PWN? Tail: SNR? Clouds?

Once we assume the spectra measured by air shower experiments, suggested to be a PeV proton acceleration in SNR.

#### Future prospective

For precise discussion at the break energies of Head & Tail, further VHE observations with better angular resolution are required to resolve the head and tail emissions.