



Building a Robust Sample of Fermi-LAT Blazars that Exhibit Periodic gamma-ray Emission

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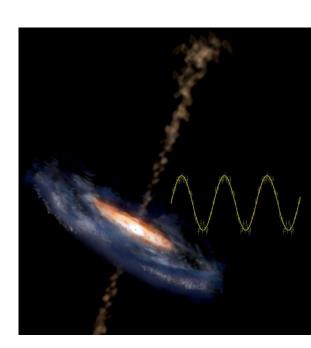
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Introduction

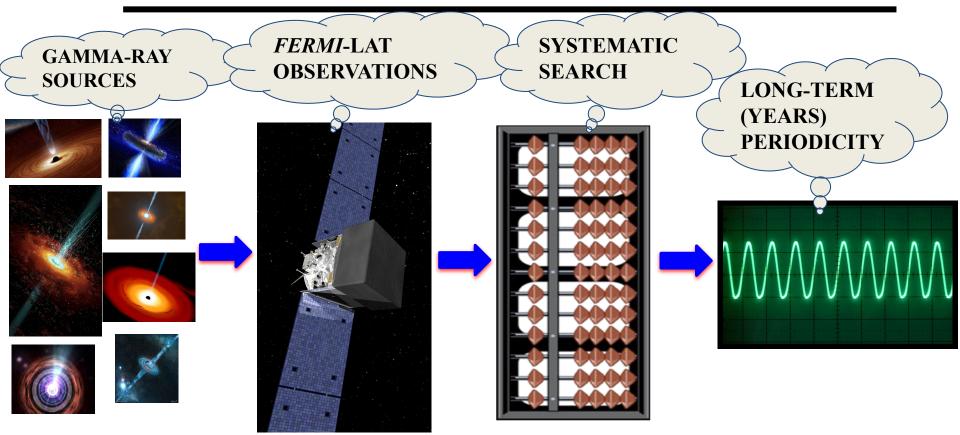


- About 10% of AGN → launch highly collimated, relativistic, jets (e.g., Sartori et al. 2019)
 - Pointed towards our line of sight → blazars
- Blazars: variability in the overall electromagnetic spectrum:
 - Different timescales:
 - Long-term variations → years or months
 - Short variations → days, hours or even minutes.
- Pattern → Periodicity
 - Provide information about its astrophysical nature



Methodology





Previous Results



3FGL+2FHL+3FHL blazars

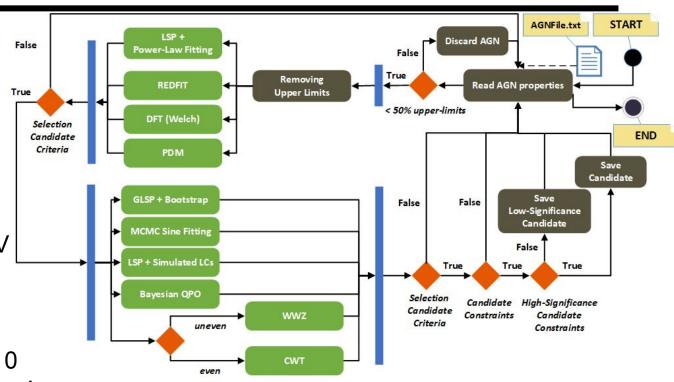
Telescope time: Aug. 2008-Sep. 2017

Data Reduction:
 Flux integrated ≥1 GeV
 28-days binning

Methods:

Periodicity detection: 10

Significance estimation: 4

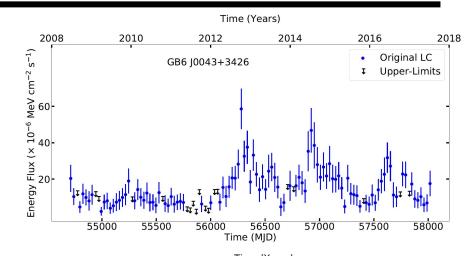


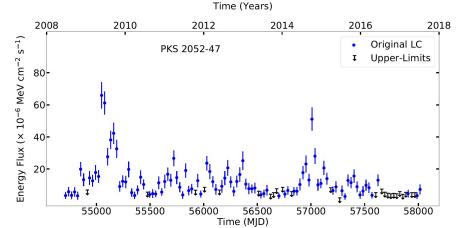
Peñil et al. 2020

Previous Results



- 11 High-significance candidates (4σ in 4 methods)
 - 9 New detections
- 13 Low-significance candidates (4σ in 3 methods)
 - 9 New detections
- 6 objects previously reported in the literature:
 - 5 with the same period
 - S5 0716+ 714





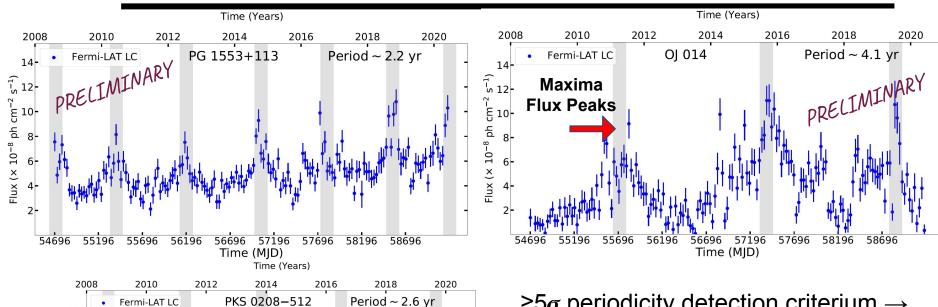
Gamma-Ray Data



- 24 periodicity candidates from previous work Peñil et al., 2020
- Light Curves
 - Telescope time: August 2008-December 2020
 - Extended with 3 extra years → total of 12 years
- Data reduction:
 - Flux integrated ≥100 MeV
 - 28-days binning

Results: 6 Blazars $\geq 5\sigma$ periodicity detection





F 80. PRELIMINARY

55696

57696

Time (MID)

58196

58696

120-

Flux (x 10⁻⁸ ph cm⁻² s⁻¹)

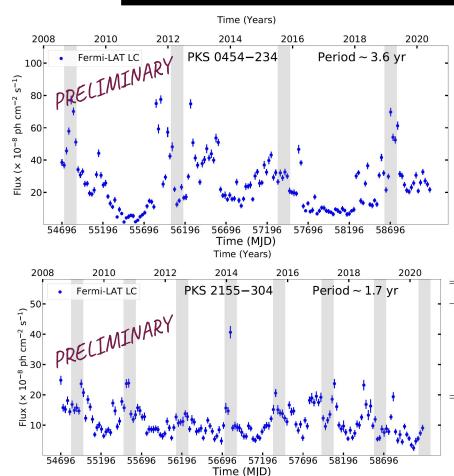
20-

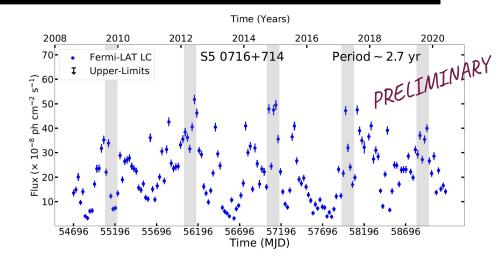
 $\geq 5\sigma$ periodicity detection criterium \rightarrow 5σ in average

Peñil et al. 2020: $>4\sigma$, $>3.5\sigma$, $>3\sigma$ respectively

Results: 6 Blazars $\geq 5\sigma$ periodicity detection







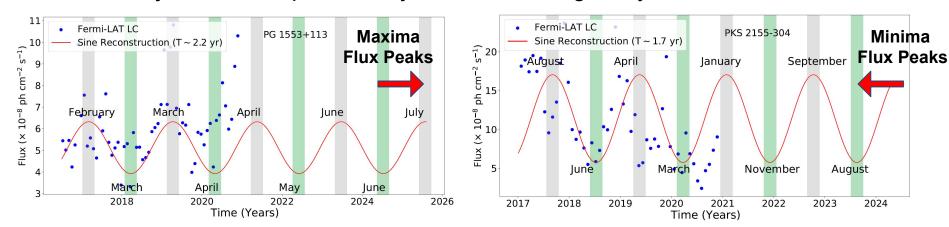
| 4FGL Source Name | RAJ2000 | $\mathrm{DecJ}2000$ | Type | Redshift | Association Name | Period (yr) |
|------------------|-----------|---------------------|------|----------|------------------|-------------|
| J0210.7-5101 | 32.68952 | -51.01695 | fsrq | 1.003 | PKS 0208-512 | 2.6 |
| J0457.0 - 2324 | 74.26096 | -23.41384 | fsrq | 1.003 | PKS $0454-234$ | 3.6 |
| J0721.9 + 7120 | 110.48882 | 71.34127 | bll | 0.127 | S5 0716+714 | 2.7 |
| J0811.3 + 0146 | 122.86418 | 1.77344 | bll | 1.148 | OJ 014 | 4.1 |
| J1555.7+1111 | 238.93169 | 11.18768 | bll | 0.36 | PG 1553+113 | 2.2 |
| J2158.8 - 3013 | 329.71409 | -30.22556 | bll | 0.116 | PKS 2155-304 | 1.7 |

Peñil et al. 2020: $>2.5\sigma$, $>2.5\sigma$, $>3\sigma$ respectively

Predicting the Future



- Predict the future:
 - 4 objects have a peak/valley emission during this year



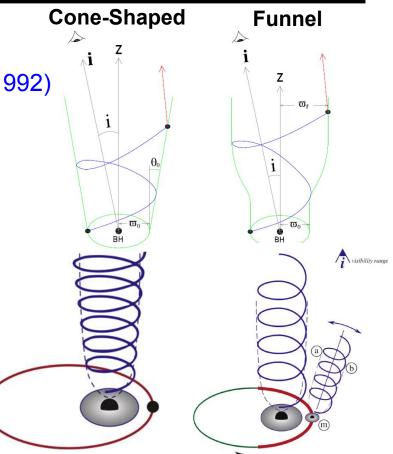
Strategy to optimize Cherenkov telescopes observations

Physical Interpretation



• Single SMBH:

- Lighthouse effects in jets
 (e.g. Camenzind and Krockenberger 1992)
- Density inhomogeneities in jets (Mohan and Mangalam, 2015)
- Binary SMBH
 - PG 1553+113:
 - Perturbed Jet(Cavaliere, A., et al., 2017)
 - Double jets (Tavani, M., et al., 2018)
- Increasing redshift → binary SMBHs (Volonteri, M., et al., 2009)



Conclusions



- We find 6 blazars with periodicity detected with $\geq 5\sigma$
 - High redshift → binary system

- Estimating the future behavior of the periodicity candidates:
 - 4 blazars to be evaluated over this year
- Any questions: <u>ppenil@clemson.edu</u>