## Expected performance of the ALTO particle detector array designed for 200 GeV - 50 TeV gamma-ray astronomy

Mohanraj Senniappan,<sup>a</sup> Yvonne Becherini,<sup>a</sup> Michael Punch,<sup>a,b</sup> Satyendra Thoudam,<sup>c</sup> Tomas Bylund,<sup>a</sup> Gašper Kukec Mezek<sup>a</sup> and Jean-Pierre Ernenwein<sup>d</sup>



<sup>a</sup> Department of Physics and Electrical Engineering, Linnaeus University, 35195 Växjö, Sweden
<sup>b</sup> Université de Paris, CNRS, Astroparticule et Cosmologie, F-75013 Paris, France
<sup>c</sup> Department of Physics, Khalifa University, PO Box 127788, Abu Dhabi, United Arab Emirates
<sup>d</sup> Aix Marseille Univ, CNRS/IN2P3, CPPM, Marseille, France





## Need for a continuous observation of the soft spectrum sources

- For ground-based particle detector arrays, extragalactic sources are difficult to detect because they have a soft spectrum due to extragalactic background light (EBL) absorption
- For instance, HAWC has detected two nearby sources: Mkn 421 and Mkn 501
- Our work focuses on the possibility of a wide field-of-view particle detector design and altitude optimised for **soft-spectrum** sources, opening the possibility for the detection, characterisation and monitoring of many more Active Galactic Nuclei and Gamma Ray Bursts

## Monte Carlo study results

- We have performed a full Monte Carlo analysis which includes CORSIKA, GEANT4, shower reconstruction, the SEMLA analysis
- We demonstrate here that, it is possible to detect AGNs and GRBs with the hardware choices we have made



VHE gamma-ray source	Approximate time to reach a 5- $\sigma$ detection
GRB 180720B *	~ 38 seconds
PKS 2155-304 flare	~ 24 minutes

\* extrapolated to the time of GRB alert

 In order to further enhance the performance during darkness, we also study the possibility of adding Cherenkov Light Collectors (CLiC)

See the contribution by **Gašper Kukec Mezek** at this conference

## Thank you for your attention

