

The Physics Case

- Interactions of primary cosmic ray nuclei with ISM produces secondary cosmic ray nuclei.
- Light secondary nuclei include Li, Be and B.
- Cross sections values are a crucial input for propagation models.

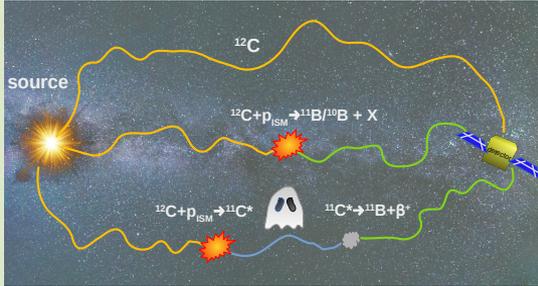
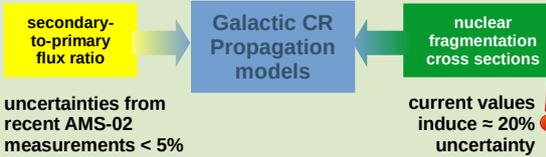


fig 1: CR propagation in the galaxy

Ghosts in Space

- ^{12}C (primary) + $p_{\text{ISM}} \rightarrow ^{11}\text{C}^* + n$ (secondary & short lived)
- $^{11}\text{C}^*$ decays to ^{11}B as $^{11}\text{C}^* \rightarrow ^{11}\text{B} + \beta^+$
- Thus infamously categorized as 'Ghost nucleus'
- Contribution to the total secondary B production cross section is significant and necessary!

Current Status [1]



Eminent need of precise laboratory measurements of nuclear fragmentation cross sections

NA61/SHINE to the Rescue

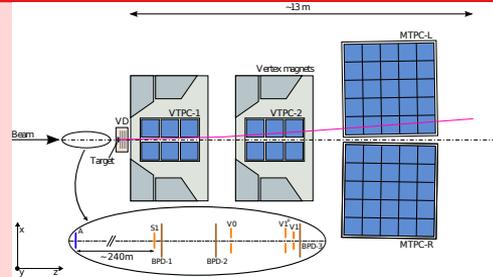
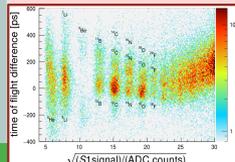


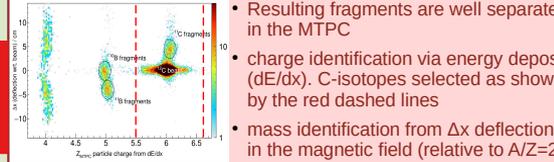
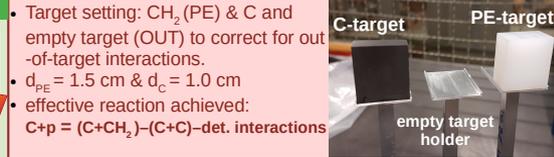
fig.2: NA61/SHINE setup

Experimental setup:

- Multipurpose experiment on the H2 beam line at CERN [2].
- Aim: to study hadron-nucleus and nucleus-nucleus interactions with fixed targets.
- Nuclear interactions can be studied up to 400 GeV/c beam momentum.
- Main detectors: 2 Vertex and 2 Main TPCs (VTPCs & MTPC) VTPCs inside superconducting magnets.



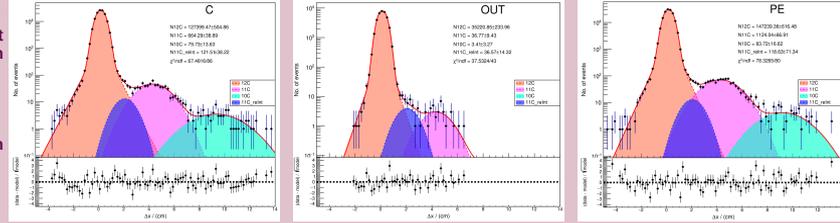
- Pilot run fragmentation data from 2018 [3]
- secondary ion beam composition at NA61/SHINE from the H2 beam line
- time of flight ($t_A - t_{S1}$) and dE/dx in S1 used for isotope tagging



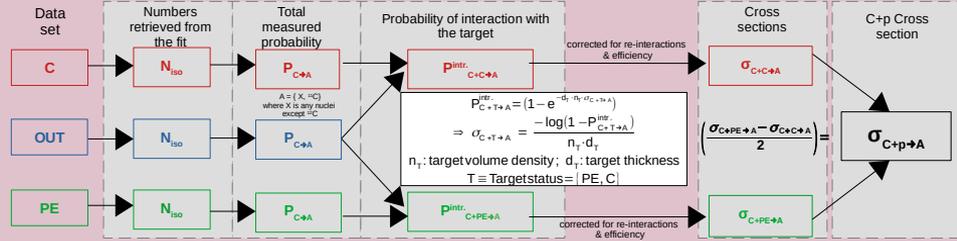
- Resulting fragments are well separated in the MTPC
- charge identification via energy deposit (dE/dx). C-isotopes selected as shown by the red dashed lines
- mass identification from Δx deflection in the magnetic field (relative to $A/Z=2$)

Analysis of Carbon Isotopes

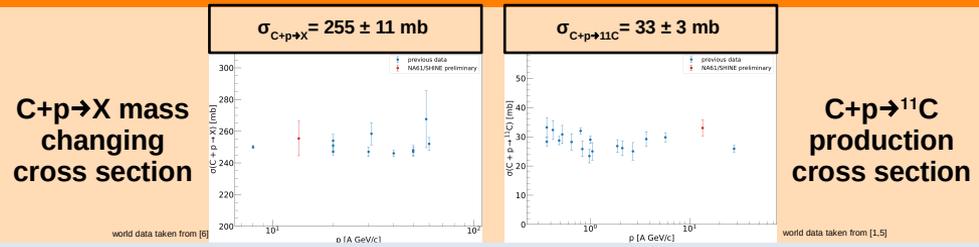
Determination of isotope yields: Fit of the distribution of deflections Δx in the magnetic field relative to the beam. (position measurement in MTPC)



Flow of the Analysis:



Results



C+p → X mass changing cross section

C+p → ^{11}C production cross section

Conclusion & Outlook

- Measurement of ^{11}B & ^{10}B [4] and ^{11}C (this work) with data from the 2018 pilot run on nuclear fragmentation with NA61/SHINE is in good agreement with previous data.
- Preliminary results prove the principle of performing fragmentation measurements at SPS energies.
- Dedicated data taking is planned for 2022: High statistics measurement of the production of light nuclear fragments (Li, Be, B) in light ion (C, N, O) fragmentation at the SPS [7].

References: [1] Y. Génolini, D. Maurin, I. V. Moskalenko, and M. Unger, Phys. Rev. C 98, 034611 (2018) [2] N. Abgrall et al., [NA61/SHINE Collab.] JINST 9 (2014) P06005 [3] A. Aduszkiewicz et al., [NA61/SHINE Collab.], 2017. CERN-SPSC-2017-035, SPSC-P-330-ADD-9. [4] M. Unger for the NA61/SHINE Collab. PoS ICRC 2019, arXiv:1909.07136v1 [5] C. Evoli, R. Aloisio, and P. Blasi Phys. Rev. D99 no. -10, (2019) 103023. [6] N. Abgrall et al., [NA61/SHINE Collab.] Eur. Phys. J. C 76 (2016) 84. [7] A. Aduszkiewicz et al., [NA61/SHINE Collab.], 2018. CERN-SPSC-2018-008. SPSC-P330-ADD-9.