

Limits on the flux of heavy compact objects from the the “Pi of the Sky” project

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Heavy compact objects

Heavy Compact Objects - hypothetical extremely heavy (and dense) objects of (usually) extremely small size, for example:

Strangelets (microscopic?) / **nuclearites** (macroscopic?) – hypothetical lumps of matter consisting of up, down and strange quarks, that could be more stable than proton/neutron based hadronic matter

- Could have been produced in very early universe
- Maybe are produced in astrophysical objects (neutron/strange stars)

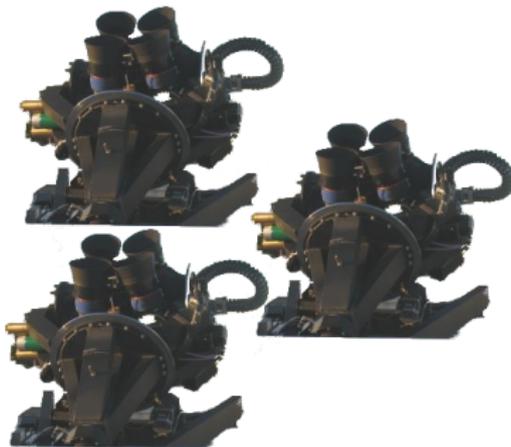
de Rujula & Glashow, Nature, vol. 312, 1984

- Could traverse through Solar System and Earth
- Could produce light tracks in the atmosphere due adiabatic compression of air (*brightness may vary with light emission mechanism by many orders of magnitude*)
- Assumptions: cosmological origin, isotropic flux, “galactic speeds“ of 250 km/s
- Can be detected in similar manner as meteors

The Pi of the Sky experiment

Closed GRB optical follow-up project consisting of 16+2 cameras monitoring the sky.

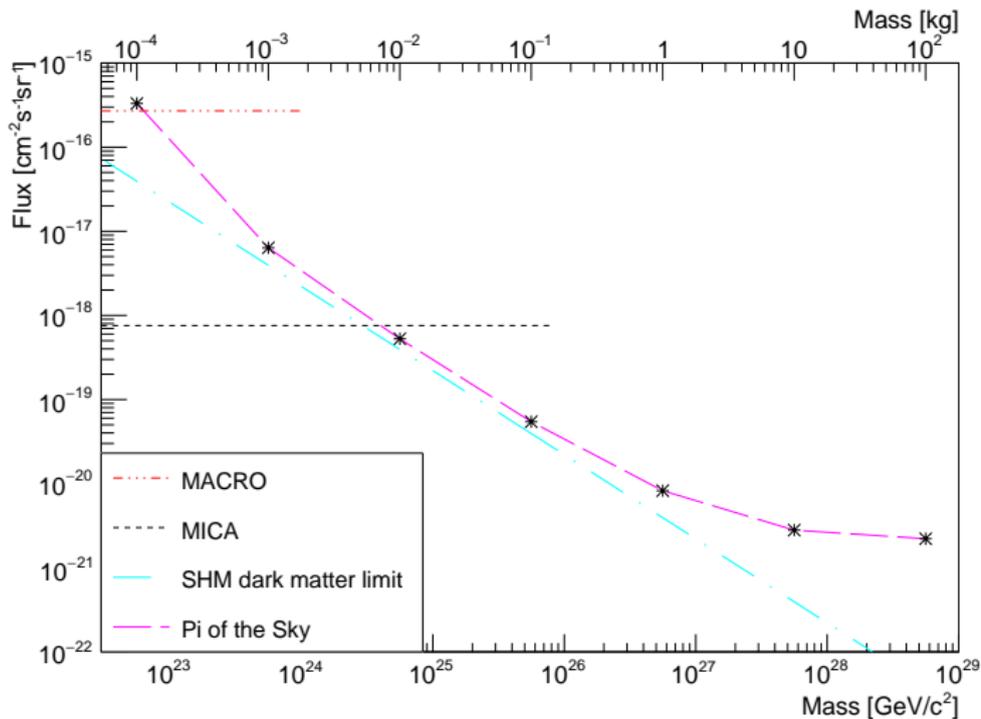
- FoV: 2 sr
- Temporal resolution: ~ 10 s
- ~ 2000 - 3000 frames/night/camera
- Large data stream: ~ 0.5 TB/night
- Detection of flashes in the real time
- Fully autonomous operation
- Range $12^m - 14^m$



Finished operations in 2016

90% CL limits on the flux of nuclearites

Assumptions like in the paper: isotropic flux from upper hemisphere, 250 km/s speed



L. W. Piotrowski et al., **Phys. Rev. Lett.** **125**, 091101 – Published 28.08.2020

The road to the flux limit

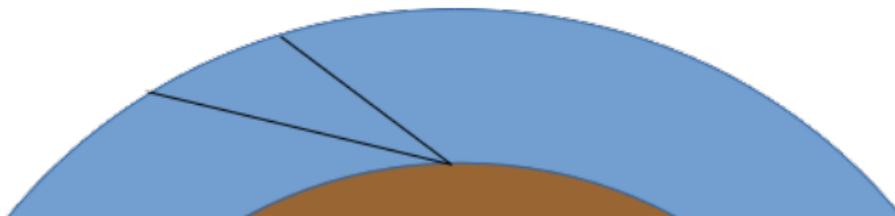
Assuming no detection, 90% CL isotropic flux limit approximation:

$$\Phi_{90CL}(m) = \frac{2.3}{S_E(m) \cdot t_e \cdot 2\pi}$$

where t_e is exposure time of a single frame,

$$S_E(m) = \sum_{\alpha=0^\circ}^{90^\circ} s_{FoV\alpha m} \cdot \epsilon_{\alpha m} \cdot n_\alpha \cdot \epsilon_S$$

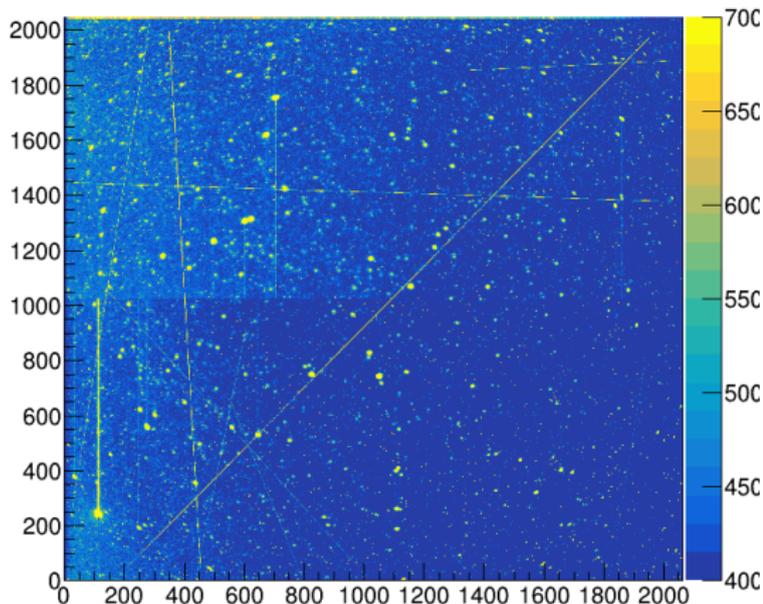
is the detector's effective area, $s_{FoV\alpha m}$ is FoV surface, $\epsilon_{\alpha m}$ is detection efficiency for given zenith angle and mass, n_α is the number of frames taken with given zenith angle, and ϵ_S is efficiency of separation from detected events (cuts to fulfil no-detection).



Detection efficiency

Efficiency of strangelets detection
– from simulation of strangelets
on Pi of the Sky sky photographs:

- Simulate trimmed (minimal burning altitude) FoV pyramid for range of masses and zenith angles – surface of walls of the pyramid as effective area of the detector
- Calculate track position and brightness on the frame: altitude, distance to the detector, atmospheric extinction



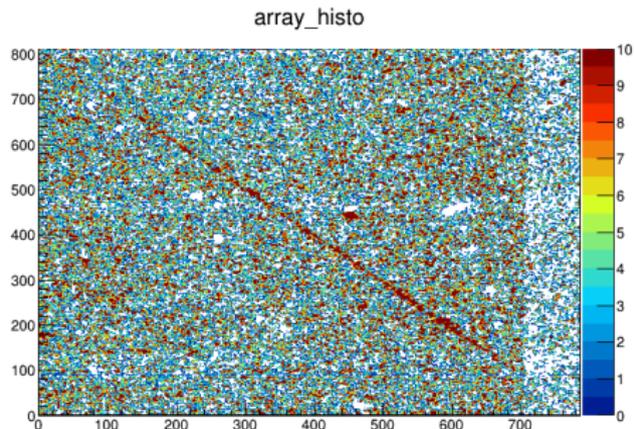
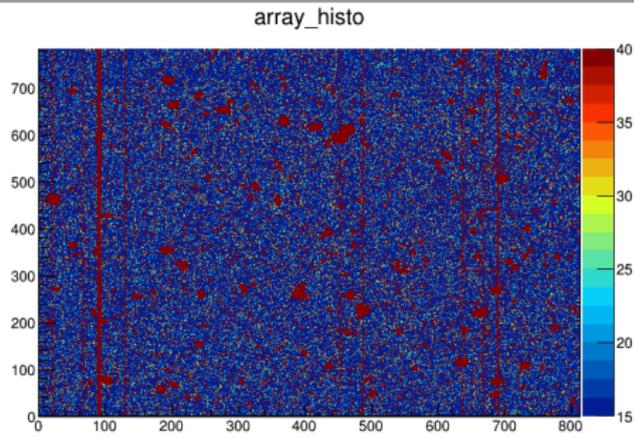
How well the strangelet tracks can be distinguished from other tracks.

- Find real events on the frame
- Find cuts that will eliminate all of them
- Find how much of simulated strangelets the cuts cut

Data and events statistics

Search for track in archival Pi of the Sky data:

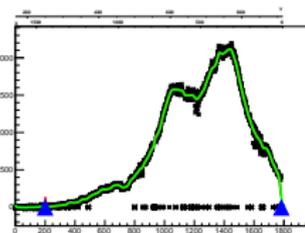
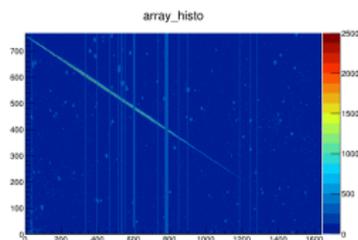
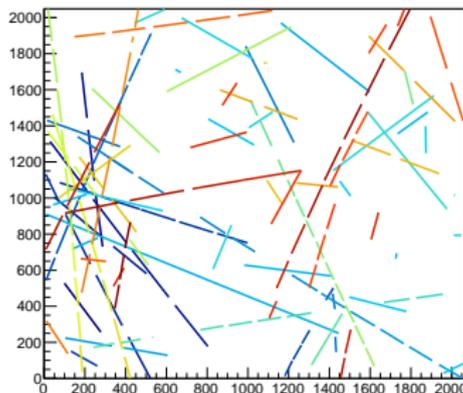
- Initially 337674 10 s frames, 34004 20*10 s summed frames
- 185258 10 s frames and 22237 20*10 s after frame quality cuts (clouds, tracking errors, horizon, dusk, down, etc.) = 1766.05 h eq. for single camera
- 35870 tracks in frames found by OpenCV HoughLinesP based algorithm, after merging and cuts



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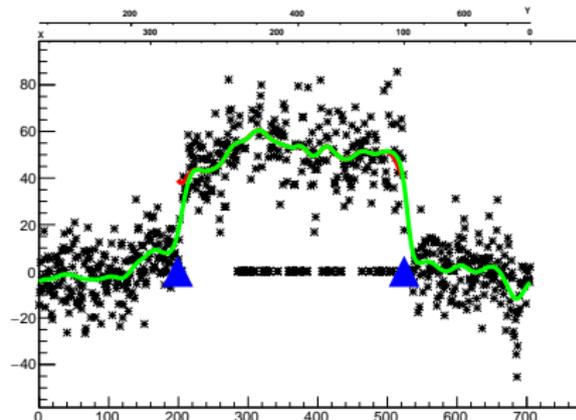
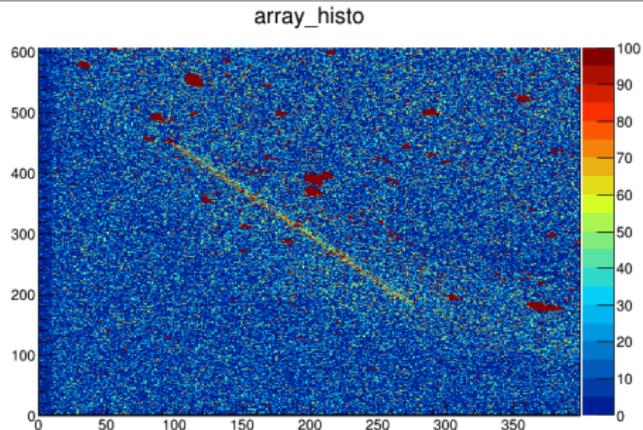
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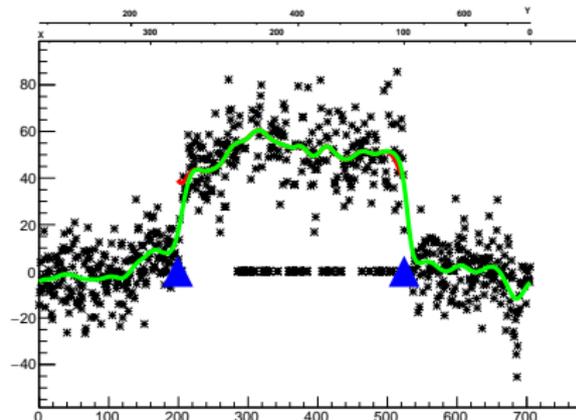
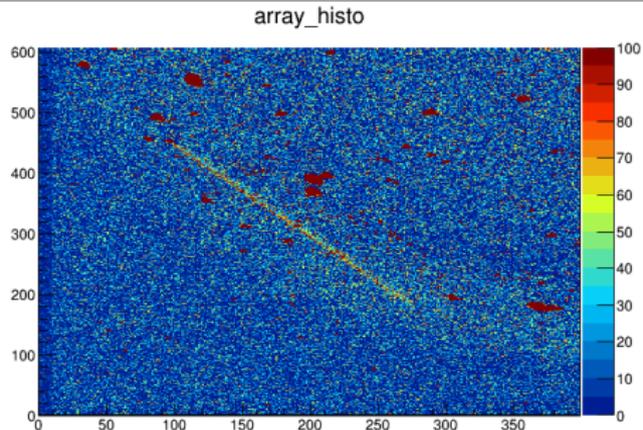
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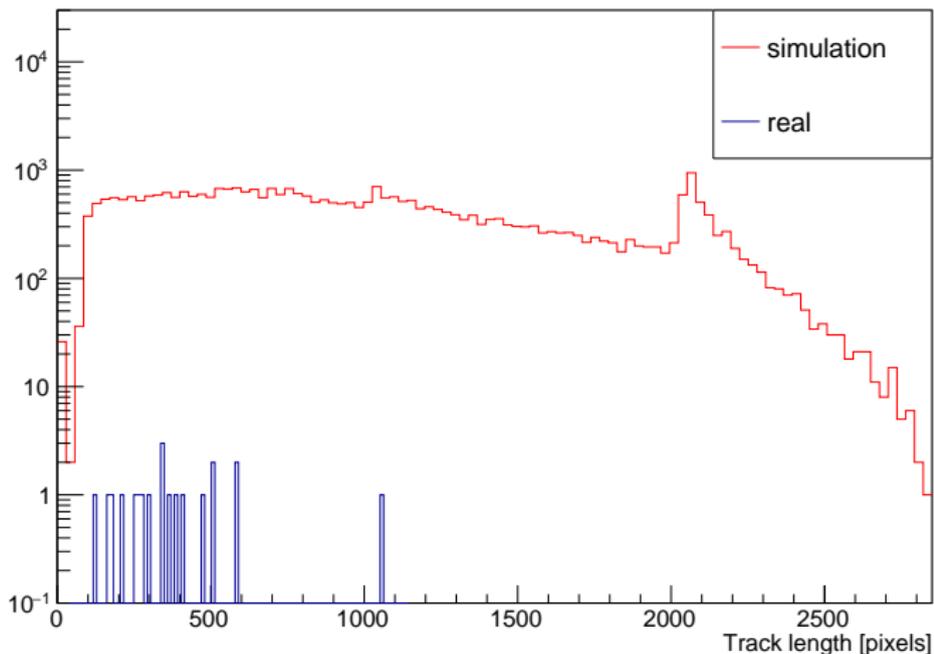
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- After cat. sat. cut: 20 candidates



Separation efficiency

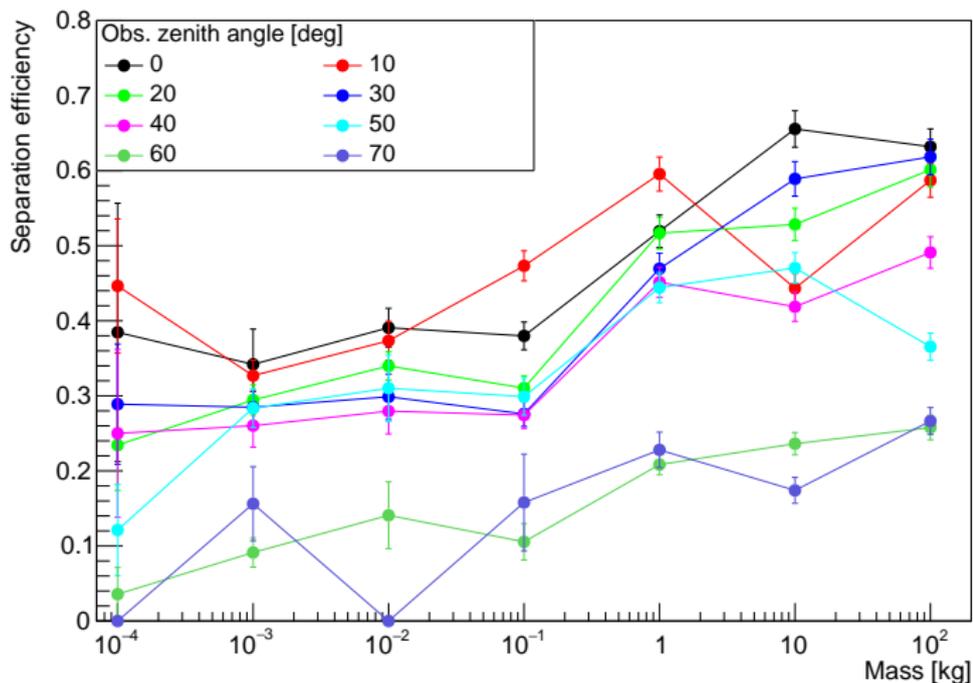
The easiest, very drastic (but controversial) cut:



Compatible with “maximum gap” method, but I could not find a good statistical method for this cut.

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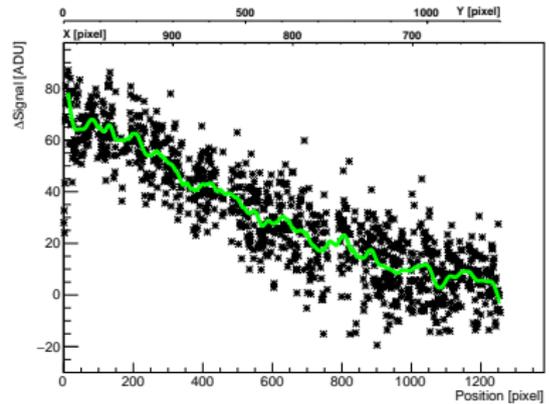
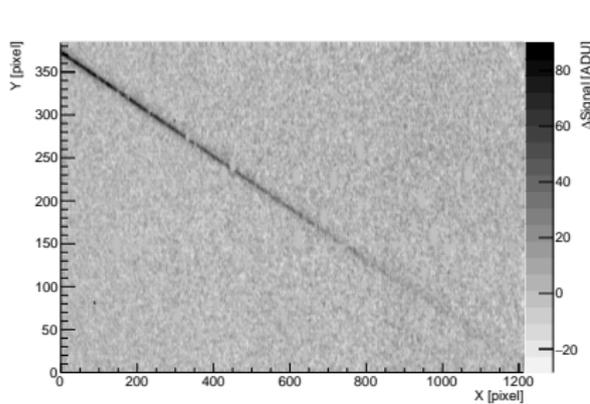
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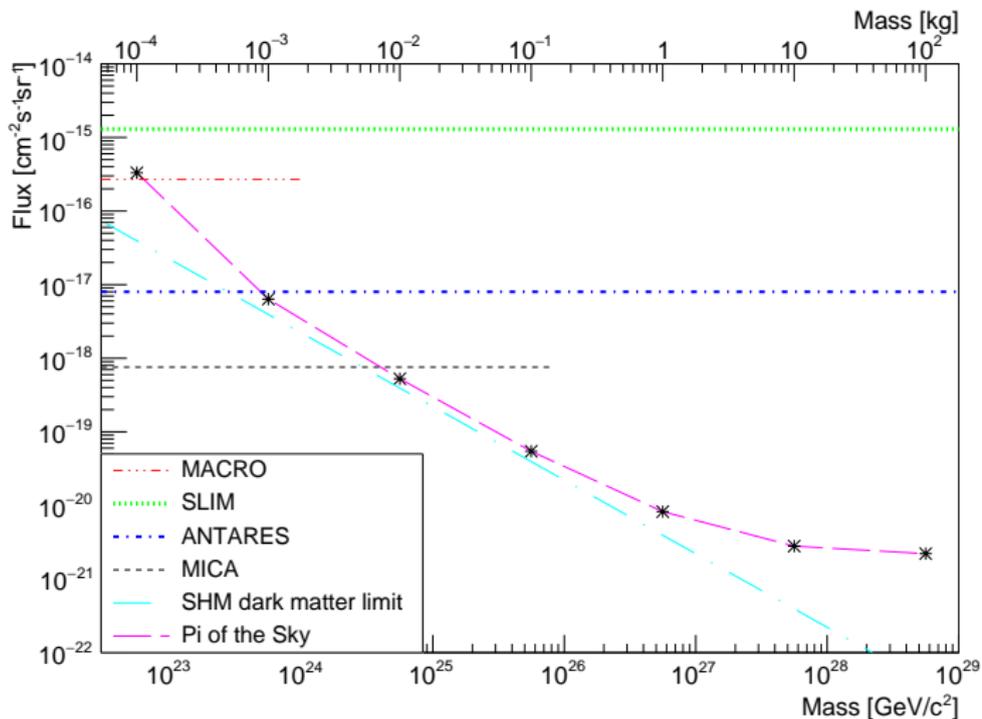
Longest candidate

Probably a satellite – flat start and flat top

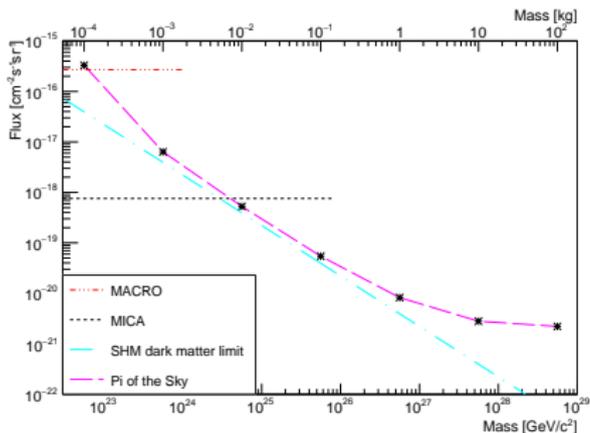


Separation efficiency

Removes all “candidates”, leaves above 50% for heavy masses



The assumptions of the original paper

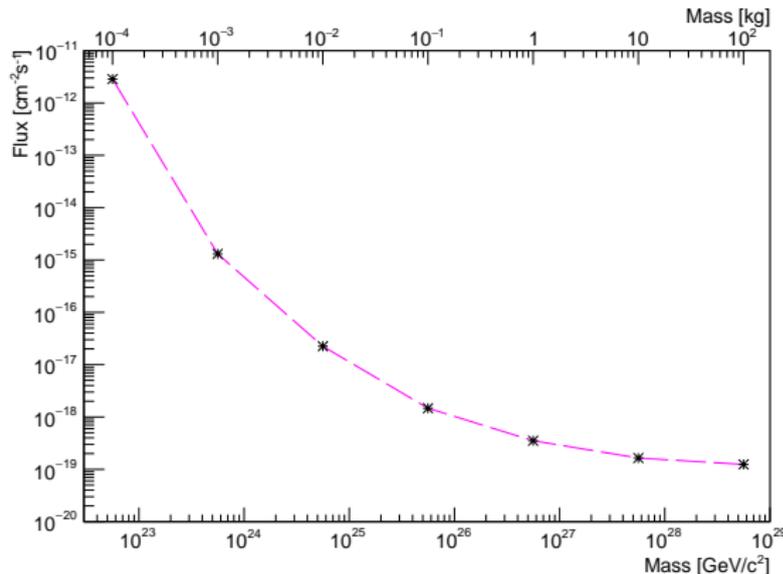


- Why isotropic shower of cosmological strangelets of speed 250 km/s?
 - “chaotic velocities characteristic of the Sun’s galactic rotation”
 - Due to the movement of Sun in the galaxy, direction would be highly unisotropic
 - Non uniform, non-gaussian (?) speed distribution modulated by Earth’s movement, etc., somewhat similar to dark matter
- Nuclearites have a mass spectrum!
- Maybe not cosmological at all?
 - Extragalactic – most likely isotropic (short GRBs, etc.)
 - Gal. neutron stars, strange stars mergers/explosions, etc. – v ? distribution?

Directional flux

Perhaps more plausible ideas for cosmological strangelets:

- A “sea” of immobile strangelets
- A “dipol” of strangelets having all possible Keplerian orbits around the centre of the galaxy convolved with Sun’s/Earth’s movement
- Some more complicated speed/direction distribution, like WIMPs isothermal sphere



A “sea” of immobile strangelets – a starting point for specific sources distribution

Result from Pi of the Sky:

- The best (only?) limits on downward isotropic flux of nuclearites with 250 km/s speed in 100 g – 100 kg range
- (The only?) limits on the “sea of nuclearites” directional flux in 0.1 g – 100 kg

Plans:

- Publish the limits in the cross-section vs mass phase space
- Analyse alternative light emission mechanisms
- Analyse some of the specific scenarios
- Repeat the analysis for Mini-EUSO, DIMS, etc.