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THE ASTROPARTICLE PHYSICS CONFERENCE



Berlin | Germany

Constraining LIV using the muon content of extensive air showers

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ICRC 2021

**For the first time, LIV effects have been
studied using the muon fluctuation of
extensive air showers measured at the
Pierre Auger Observatory!**



**New Limits on LIV parameters
have been derived**

How to break Lorentz Invariance

Modified dispersion relation

$$E^2 - p^2 = m^2 + f(\vec{p}, M_{Pl}; \eta) \longrightarrow E^2 - p^2 = m^2 + \sum_{n=0}^N \eta^{(n)} \frac{p^{n+2}}{M_{Pl}^n}$$

Where $\eta^{(n)}$ is a dimensionless constant and is called LIV parameter.
It depends on the secondary and the primary particle.

Leading order $n=1:$ $E^2 - p^2 = m^2 + \eta^{(1)} \frac{p^3}{M_{Pl}}$

Nuclei: $E_{A,Z}^2 - p_{A,Z}^2 = m_{A,Z}^2 + \eta_{A,Z}^{(1)} \frac{p_{A,Z}^3}{M_{Pl}}$
With $\eta_A = \eta/A^2$

We consider the right-hand side of the modified dispersion relation as a new mass:

$$m_{\text{LIV}}^2 = m^2 + \eta^{(n)} \frac{p^{n+2}}{M_{Pl}^n}$$

We can define the Lorentz factor as: $\gamma_{\text{LIV}} = \frac{E}{m_{\text{LIV}}}$ **In terms of the lifetime τ of particles:** $\tau = \gamma_{\text{LIV}} \tau_0$

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How to break Lorentz Invariance

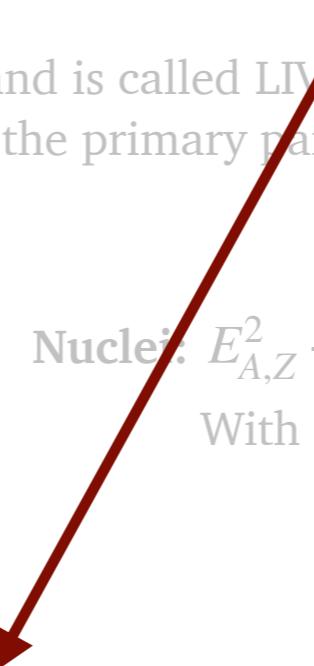
Modified dispersion relation

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We can define the Lorentz factor as: $\gamma_{\text{LIV}} = \frac{E}{m_{\text{LIV}}}$ In terms of the lifetime τ of particles: $\tau = \gamma_{\text{LIV}} \tau_0$

$\eta^{(n)}$ assumes both positive and negative values!

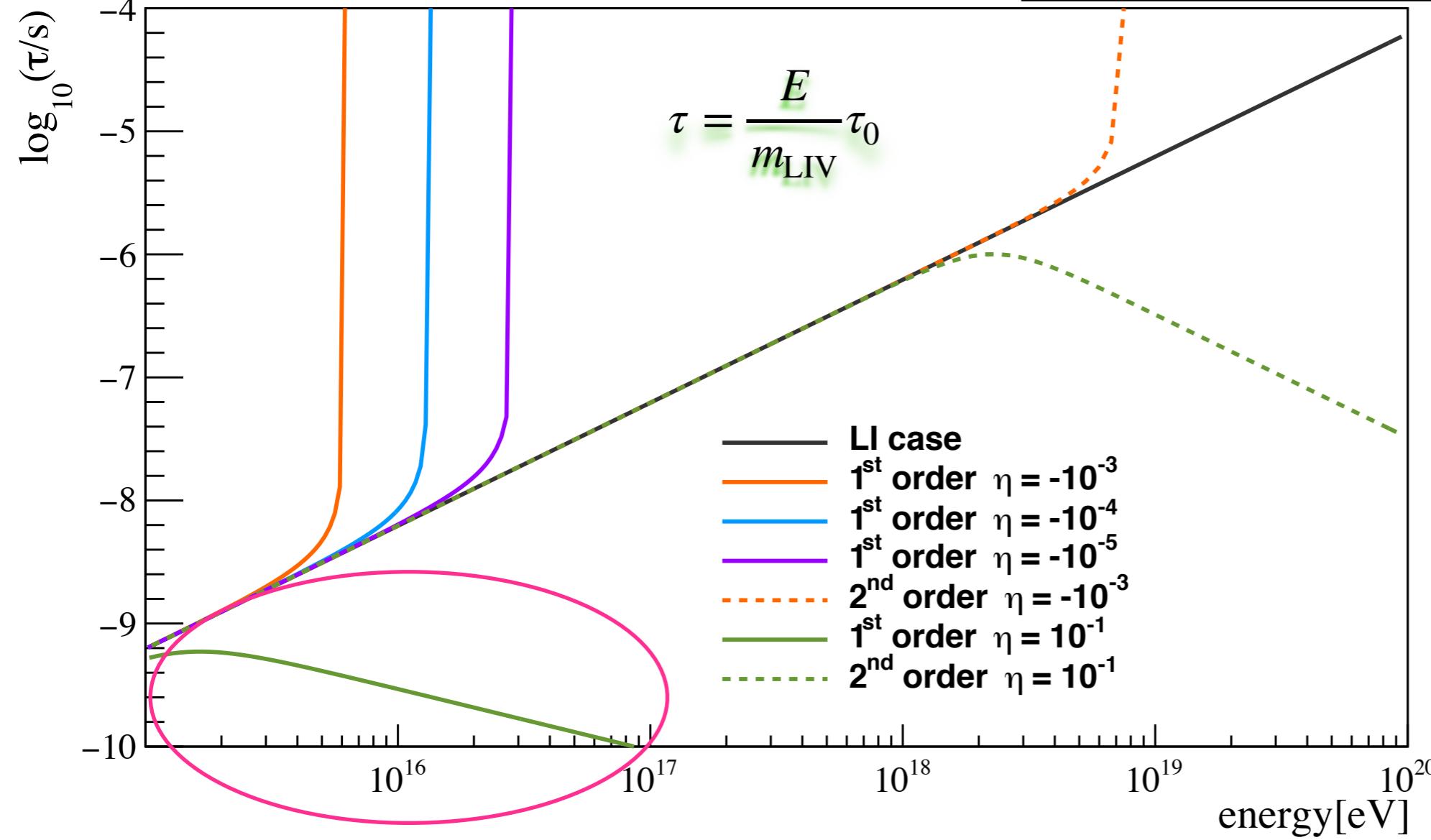
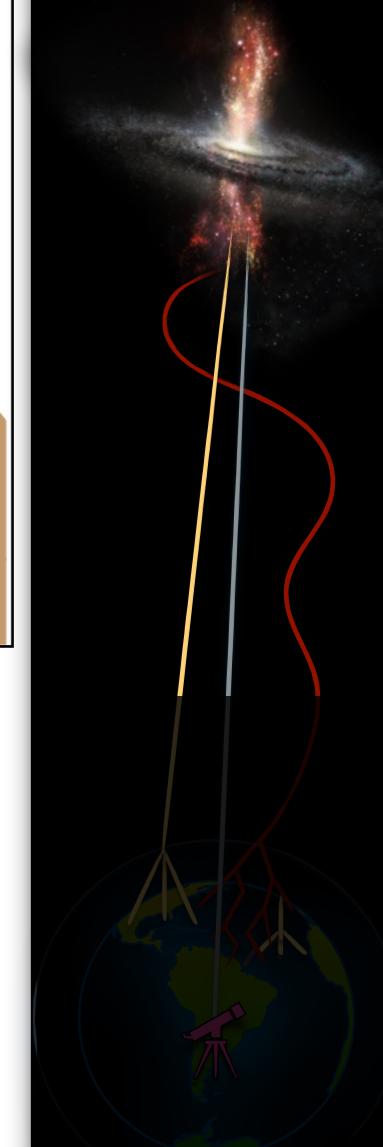
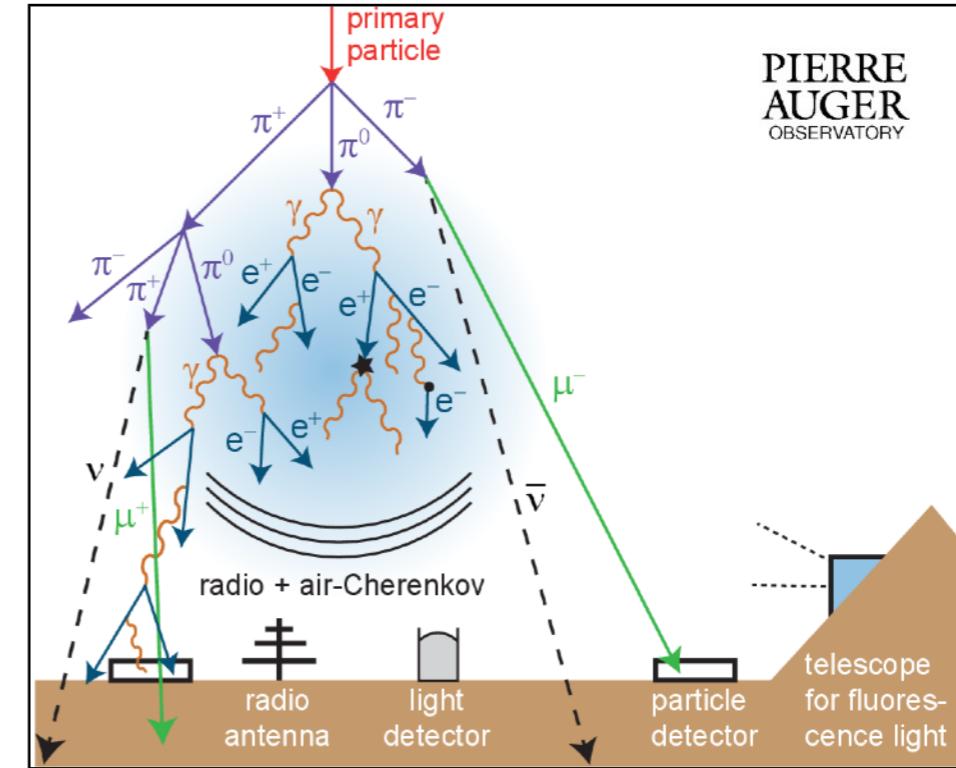
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Neutral pion decay

π^0 Decay: $\pi^0 \rightarrow \gamma\gamma$ $\tau_0 = 8.4 \cdot 10^{-17} \text{ s}$

If η_π assumes positive values!

Negligible effects are observed!



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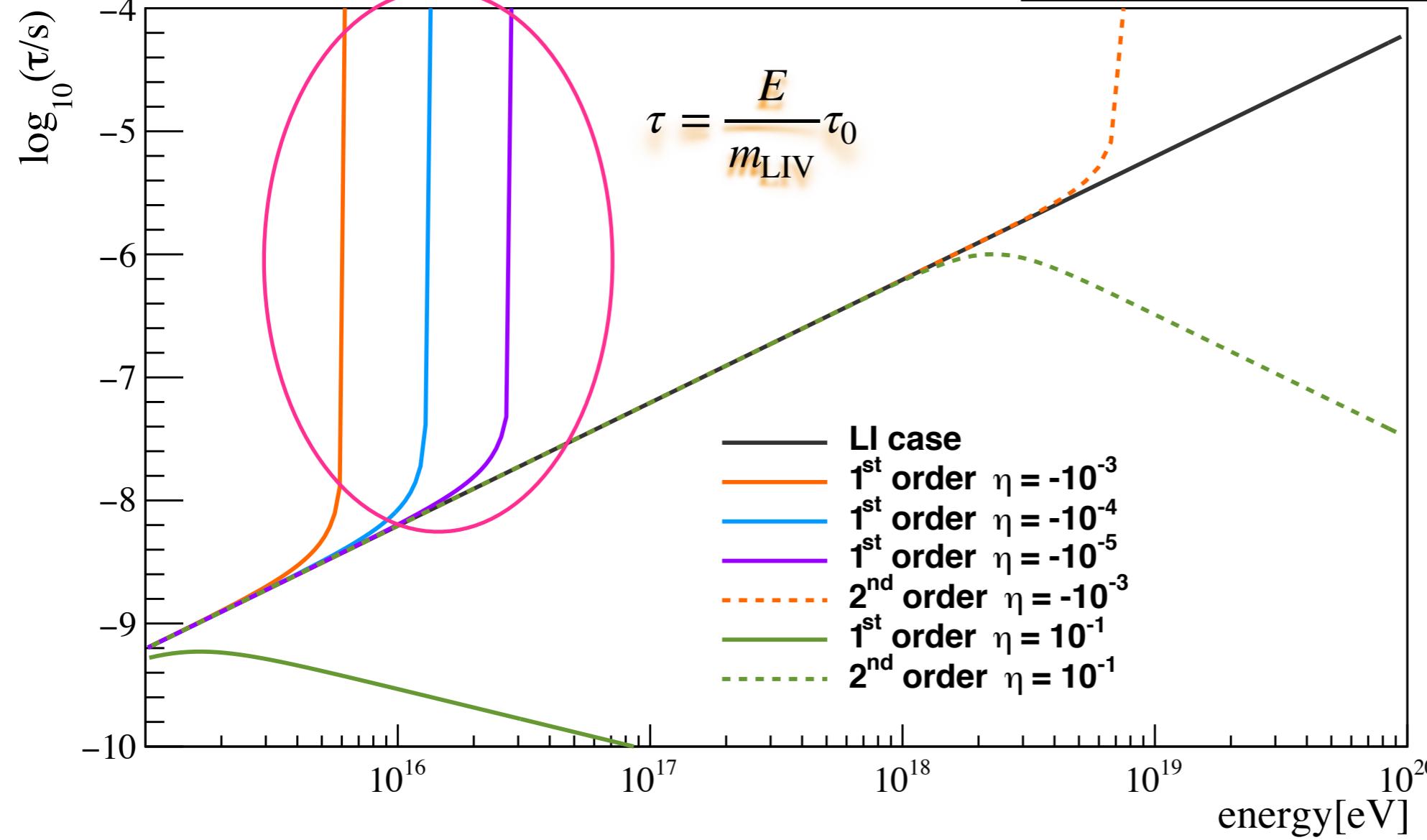
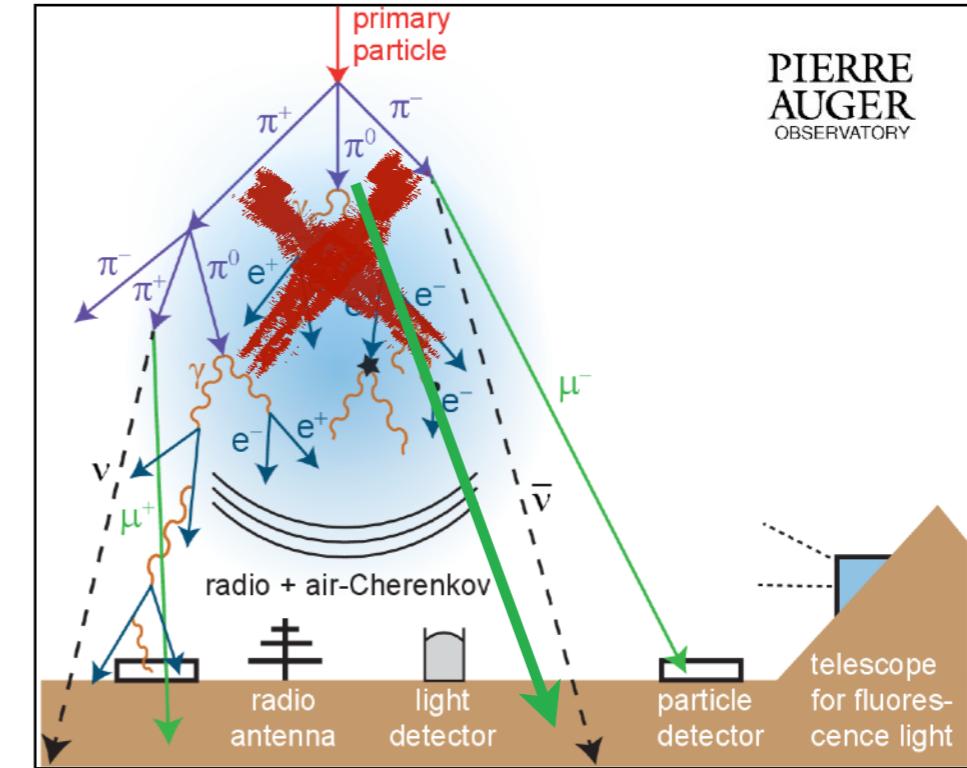
Neutral pion decay

π^0 Decay: $\pi^0 \rightarrow \gamma\gamma$ $\tau_0 = 8.4 \cdot 10^{-17} \text{ s}$

If η_π assumes negative values!

The decay is forbidden if:

$$m_\pi^2 + \eta_\pi^{(n)} \frac{p_\pi^{n+2}}{M_{Pl}^n} < 0.$$

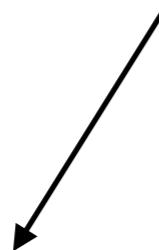


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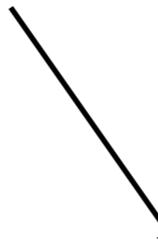
What to expect...

$$\eta < 0$$

- For increasing energy π^0 begins to interact;
- After the critical point (where $M_{\text{LIV}} = 0$) the decay $\pi^0 \rightarrow \gamma\gamma$ is forbidden;



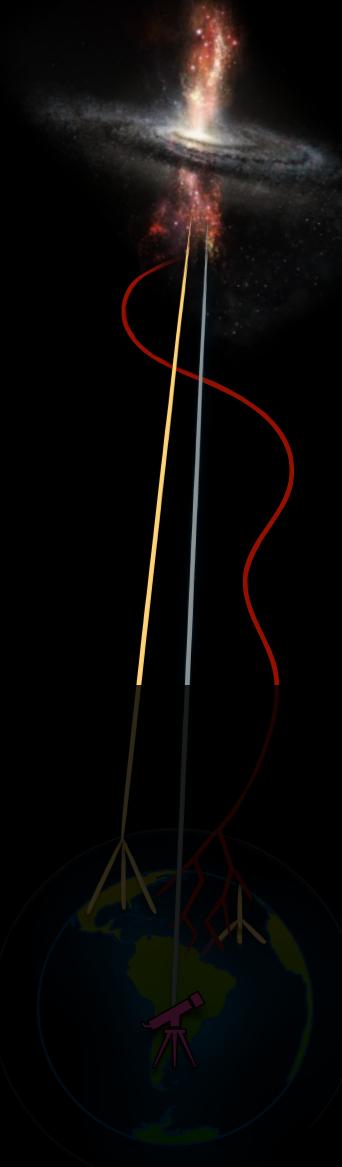
A growth of the hadronic cascade
Increase in the number of muons



Decrease in the
electromagnetic component

$$\eta > 0$$

- Negligible effects are produced



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First order: $\eta_\gamma^{(1)} > -1.2 \cdot 10^{-10}$

(R. G. Lang, H. Martínez-Huerta and V. De Souza 2018);

Second order: $-10^{-3} < \eta_\pi^2 < 10^{-1}$

(Maccione et al. 2009).

CONEX

shower simulation

Lorentz Invariant case & in presence of LIV

Shower Simulation Options:

Primary particles: H, He, N, Si, Fe;

Primary particle energy: 10^{14} - 10^{21} eV;

Zenith angle: $\theta = 70^\circ$;

21 energy bins of width $\Delta \log_{10}(E/eV) = 0.25$ ranging from 10^{14} to 10^{21} ;

Hadronic interaction model: EPOS LHC-LIV, QGSJETII-04.

in presence of LIV

LIV parameter η :

- **1st order:** $\eta = -10^{-1}, -10^{-3}, -10^{-4}, -10^{-5}, -10^{-6}, -5 \cdot 10^{-7}, -10^{-7}, -10^{-8}$

- **2nd order:** $\eta = -1, -10^{-1}, -10^{-2}, -10^{-3}$

A number of 5000 events has been simulated for each primary particle for definite energy intervals.

Modified particle velocity definition
$$\beta = \frac{P}{m\gamma_{LIV}}$$

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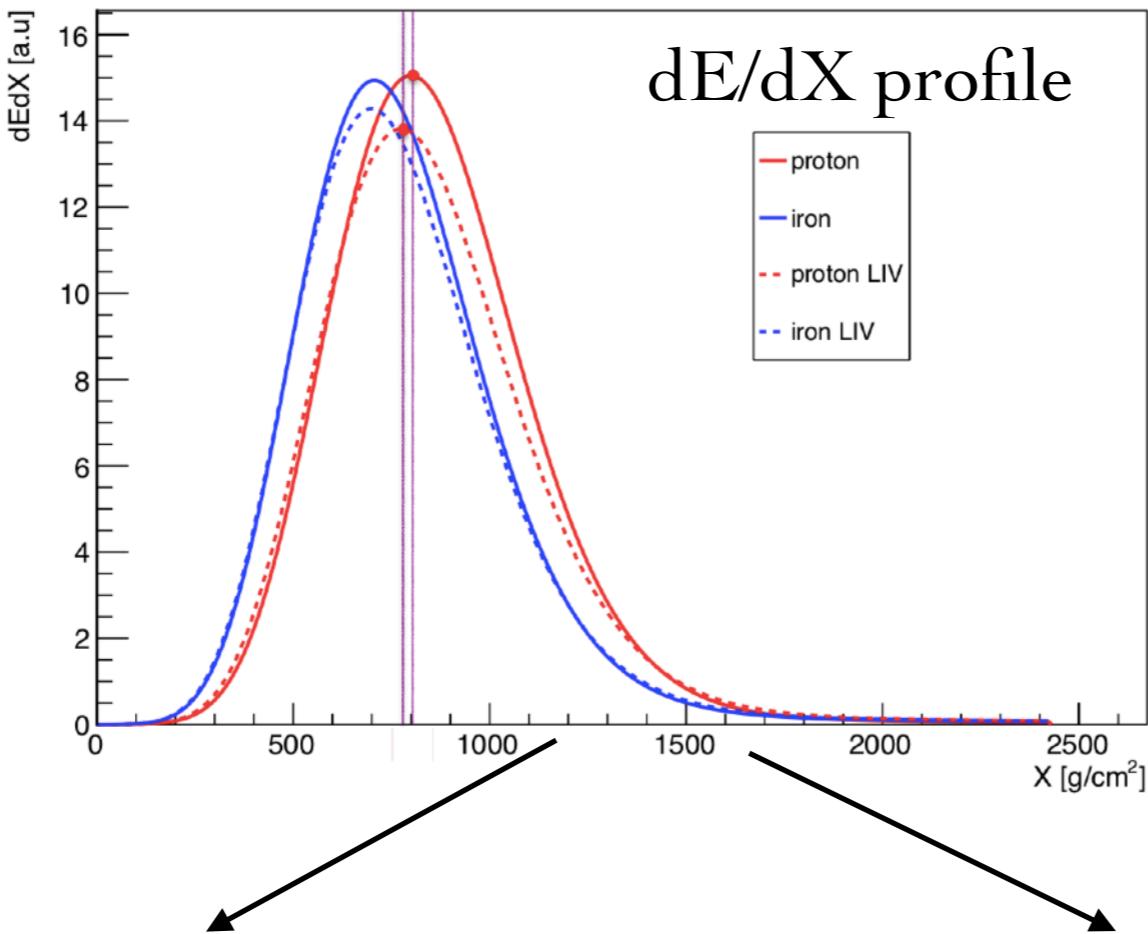
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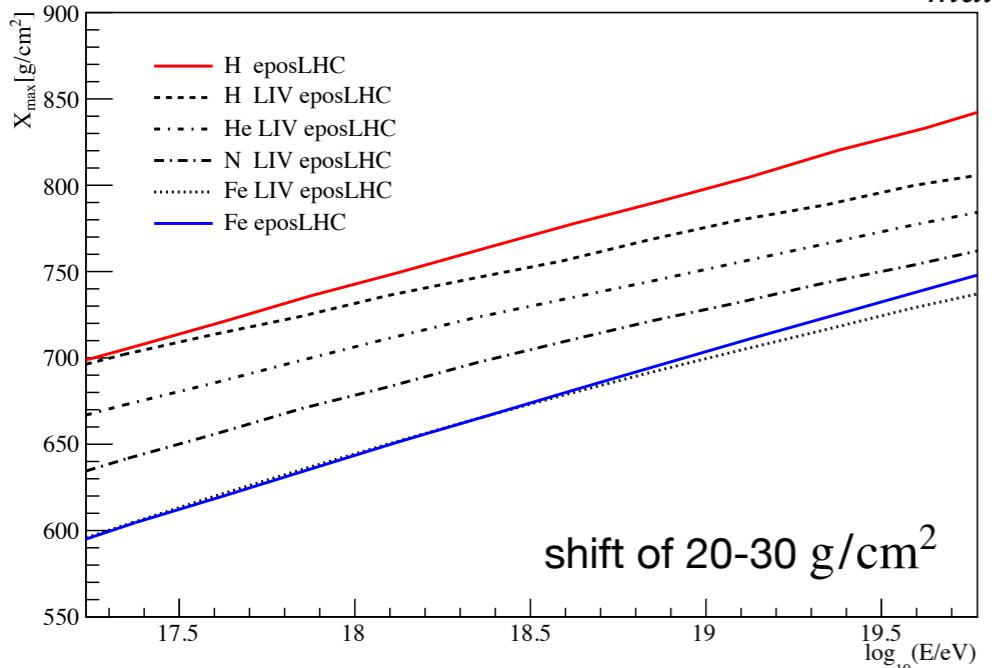
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LIV effects on air shower development

ORDER OF LIV $n=1$
EPOS-LHC
 $\eta = -10^{-3}$
at 10^{19} eV



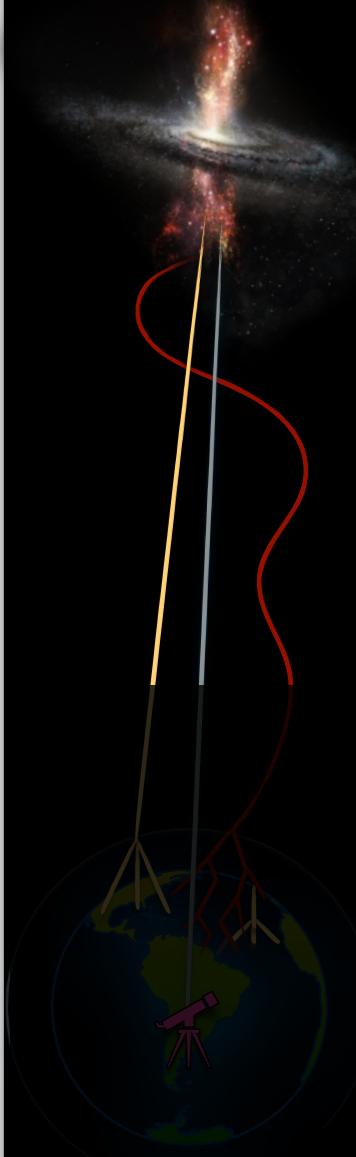
1. Shift of position of the maximum X_{max}



2. Reduction in the normalization

- Due to a change in the number of muons at the ground;
- The calorimetric energy deposited in the atmosphere in the presence of LIV is lower than the standard one.

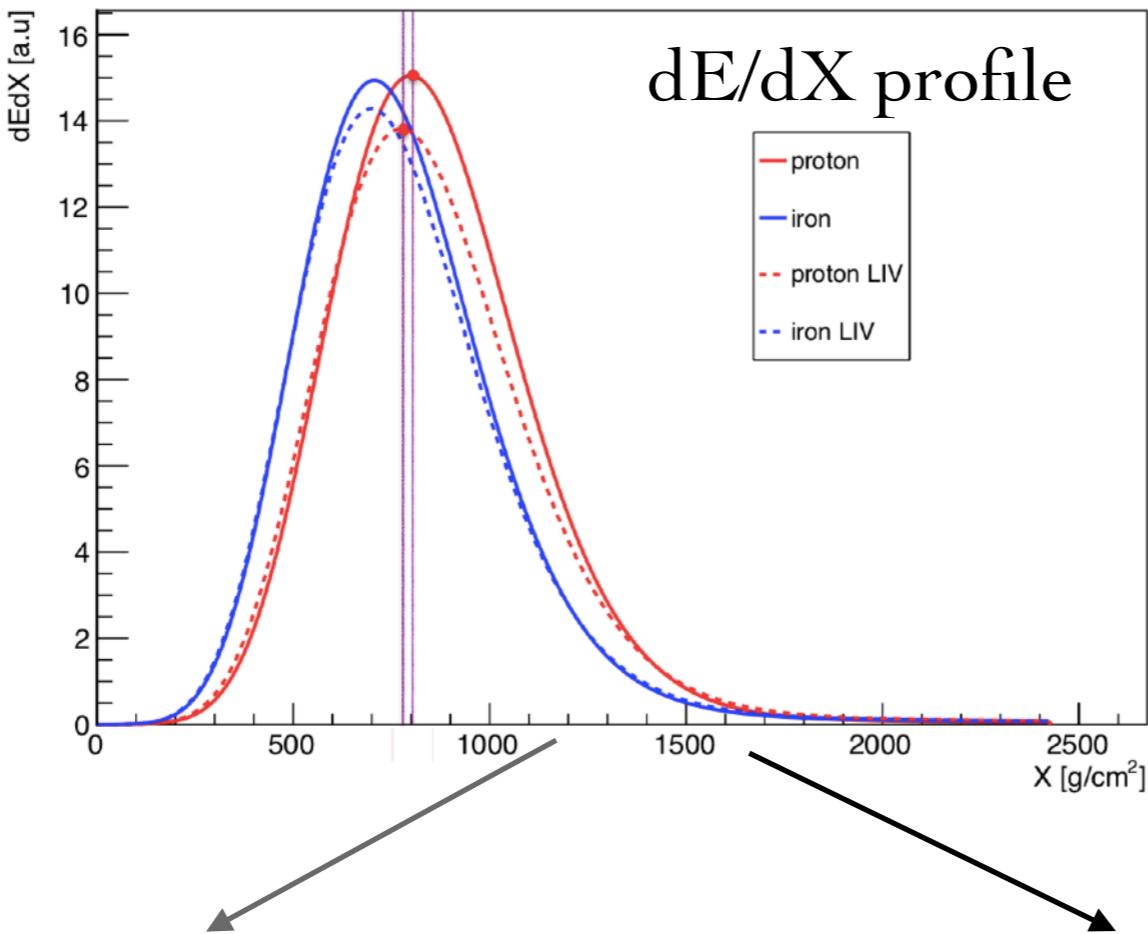
MUON CONTENT DISTRIBUTION



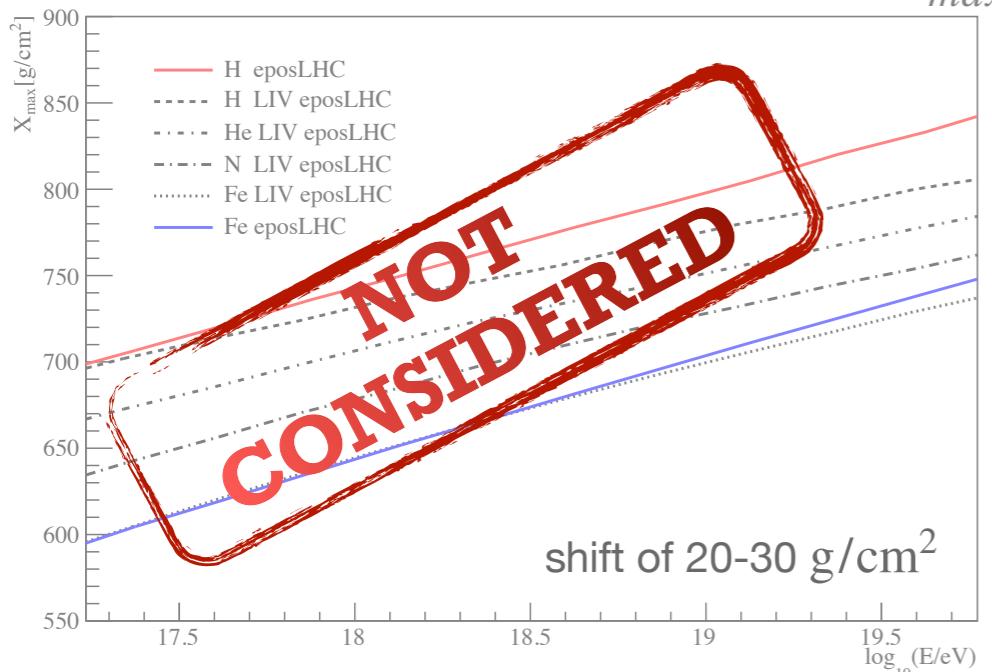
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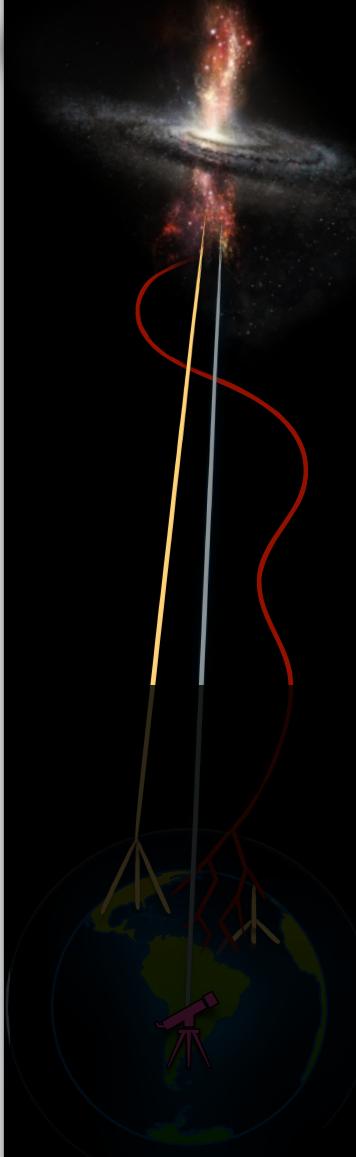
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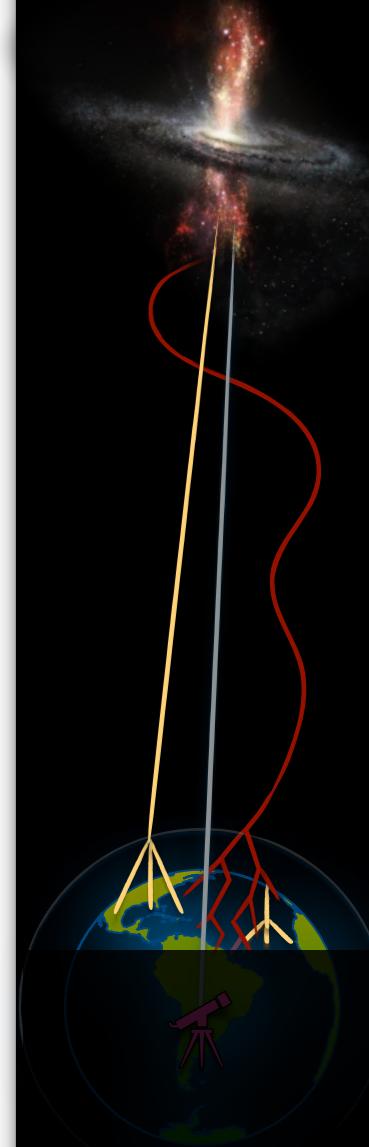
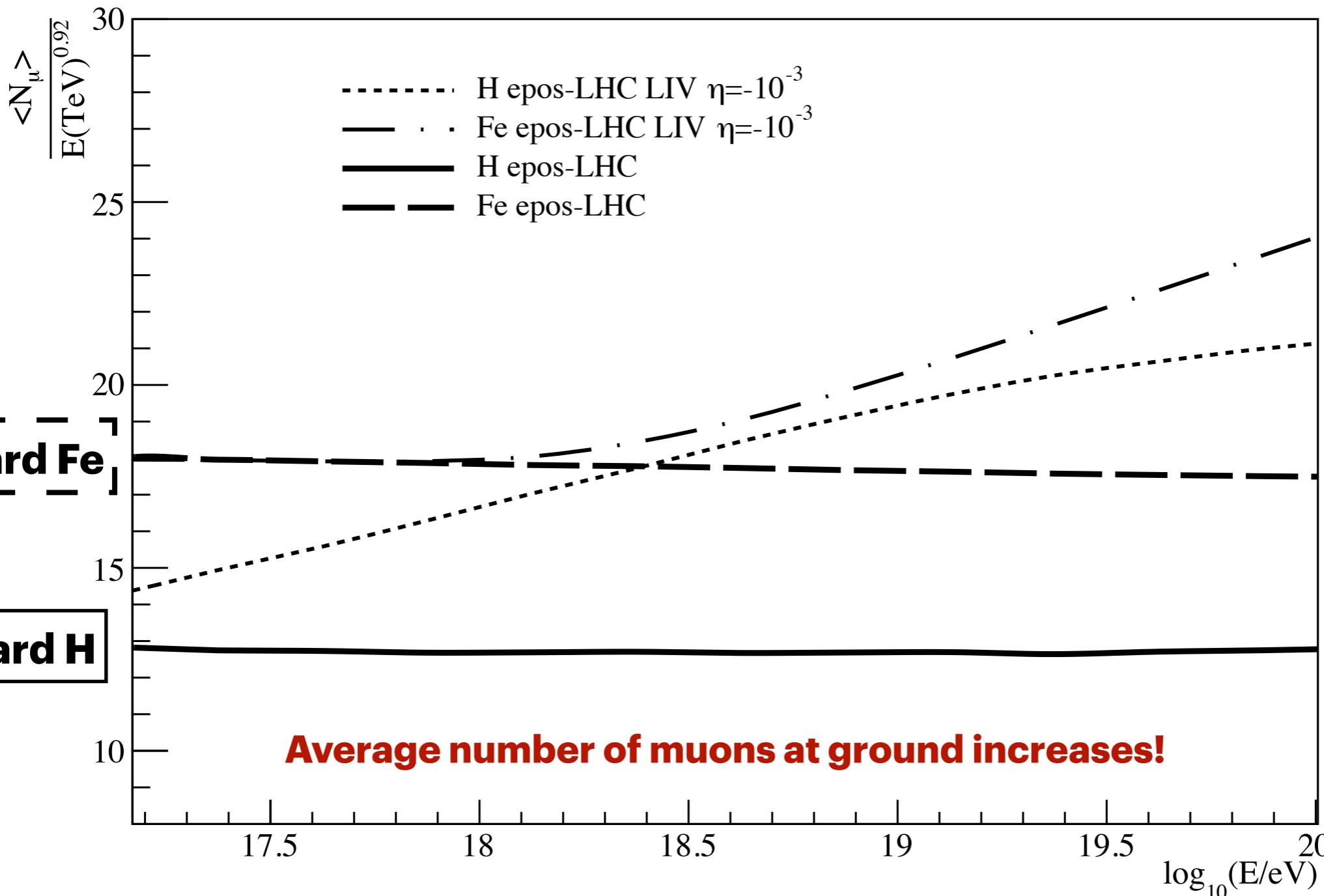


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Average number of muons

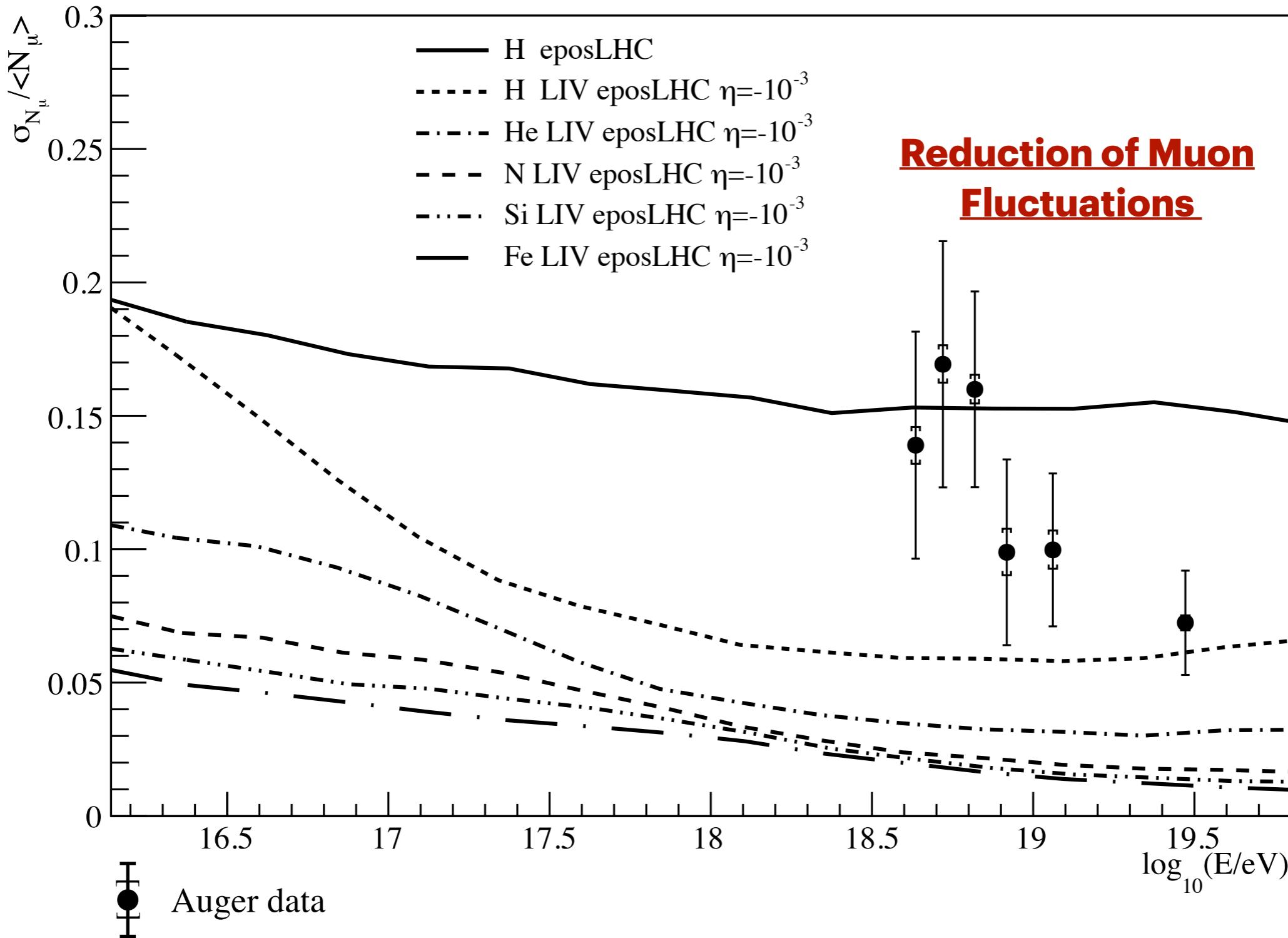


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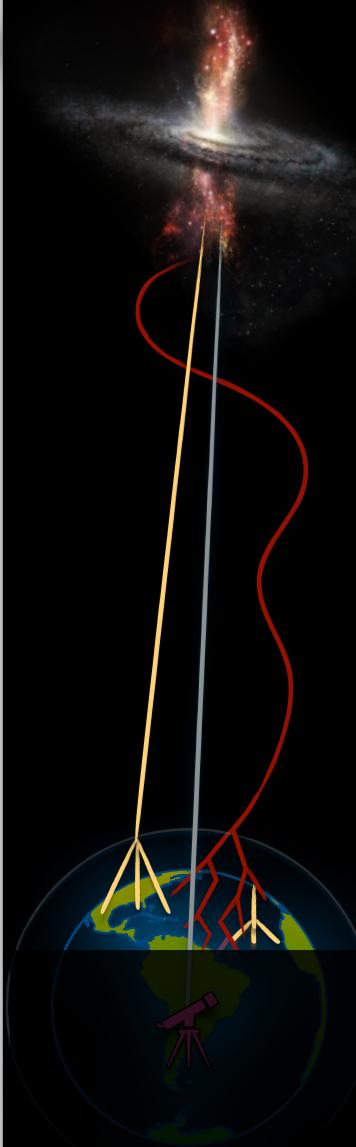
Relative Fluctuations

Ratio of the fluctuations to the average number of muons



Reduction of Muon Fluctuations

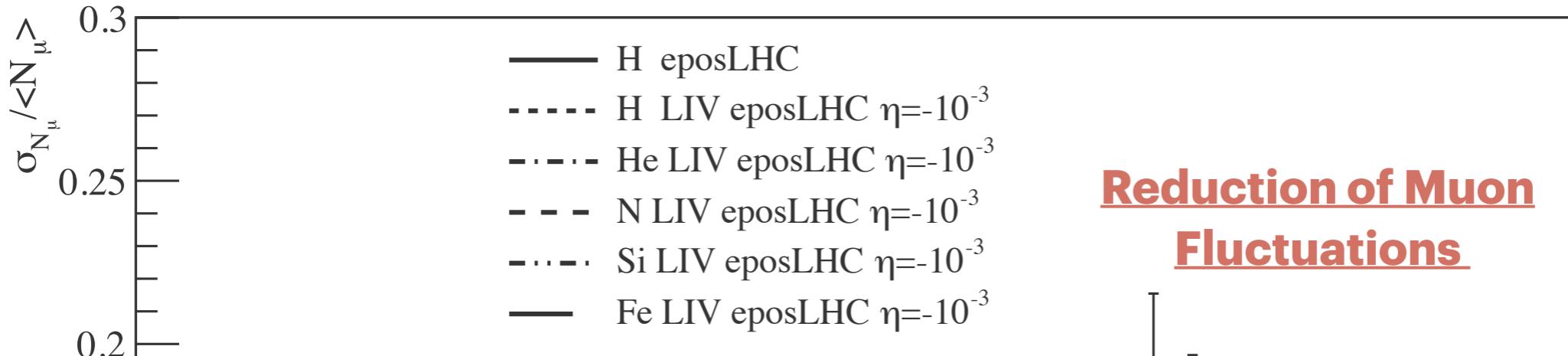
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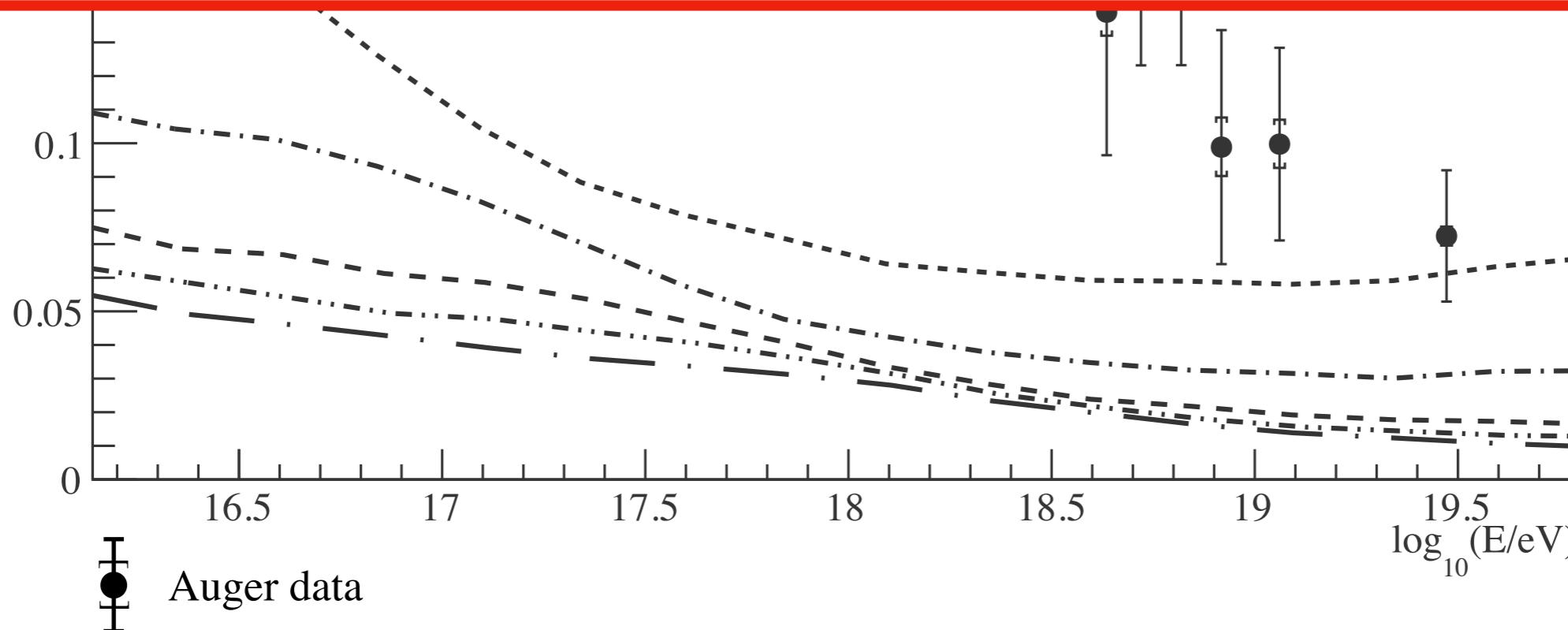
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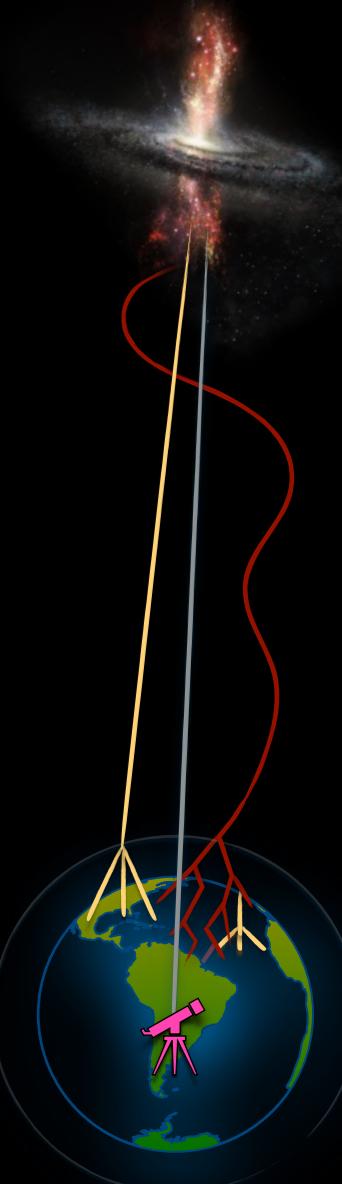
Considering the dependence of the decrease of the relative fluctuations on the different LIV strengths, a new bound for the LIV parameter can be obtained



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Looking for... the most conservative relative Fluctuations

1. **Effects of the different composition scenarios:** The mixture of the two components p and Fe gives the maximum value of relative fluctuations.
2. Define $\frac{\sigma_\mu}{\langle N_\mu \rangle} = \frac{\sqrt{\sigma^2(N_\mu)_{\text{mix}}(\alpha; \eta)}}{\langle N_\mu \rangle_{\text{mix}}(\alpha; \eta)}$ *
 1 – α is the fraction of proton
 α is the fraction of iron
3. Parametrize as **function of η and energy** $\langle N_\mu \rangle_p$, $\langle N_\mu \rangle_{Fe}$, $\sigma(N_\mu)_p$ and $\sigma(N_\mu)_{Fe}$



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*

$$\langle N_\mu \rangle_{\text{mix}}(\alpha; \eta) = (1 - \alpha)\langle N_\mu \rangle_p + \alpha\langle N_\mu \rangle_{Fe}$$
$$\sigma^2(N_\mu)_{\text{mix}}(\alpha; \eta) = (1 - \alpha)\sigma^2(N_\mu)_p + \alpha\sigma^2(N_\mu)_{Fe} + (\alpha(1 - \alpha)(\langle N_\mu \rangle_p - \langle N_\mu \rangle_{Fe})^2)$$

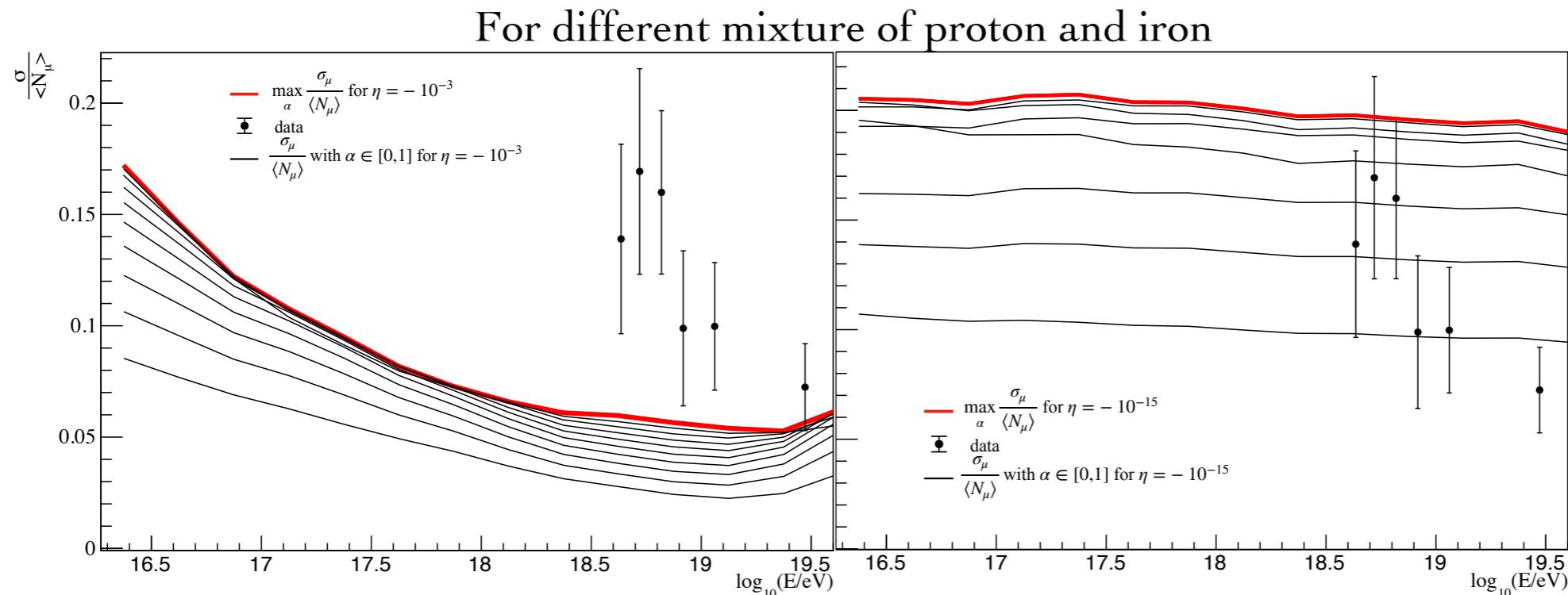
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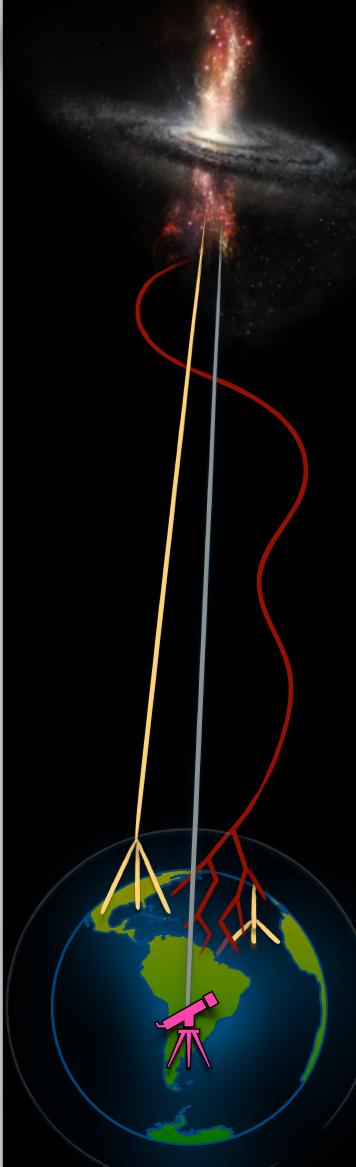
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The maximum wrt α curve is always above the curves given by any other α combinations



Only if the fluctuations stand below the data the $\max_{\alpha} \frac{\sigma_\mu}{\langle N_\mu \rangle}$ is the most conservative LIV model



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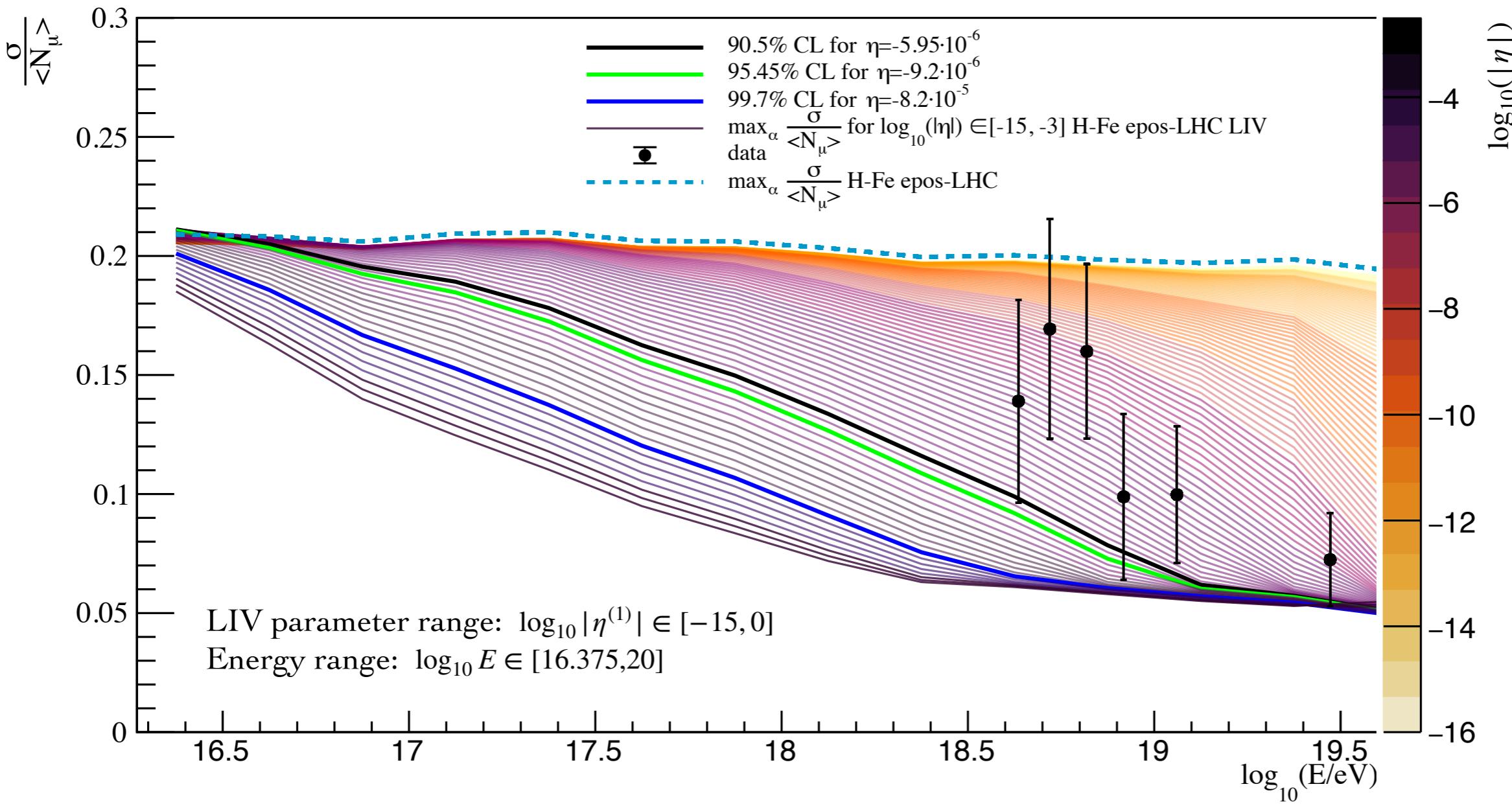
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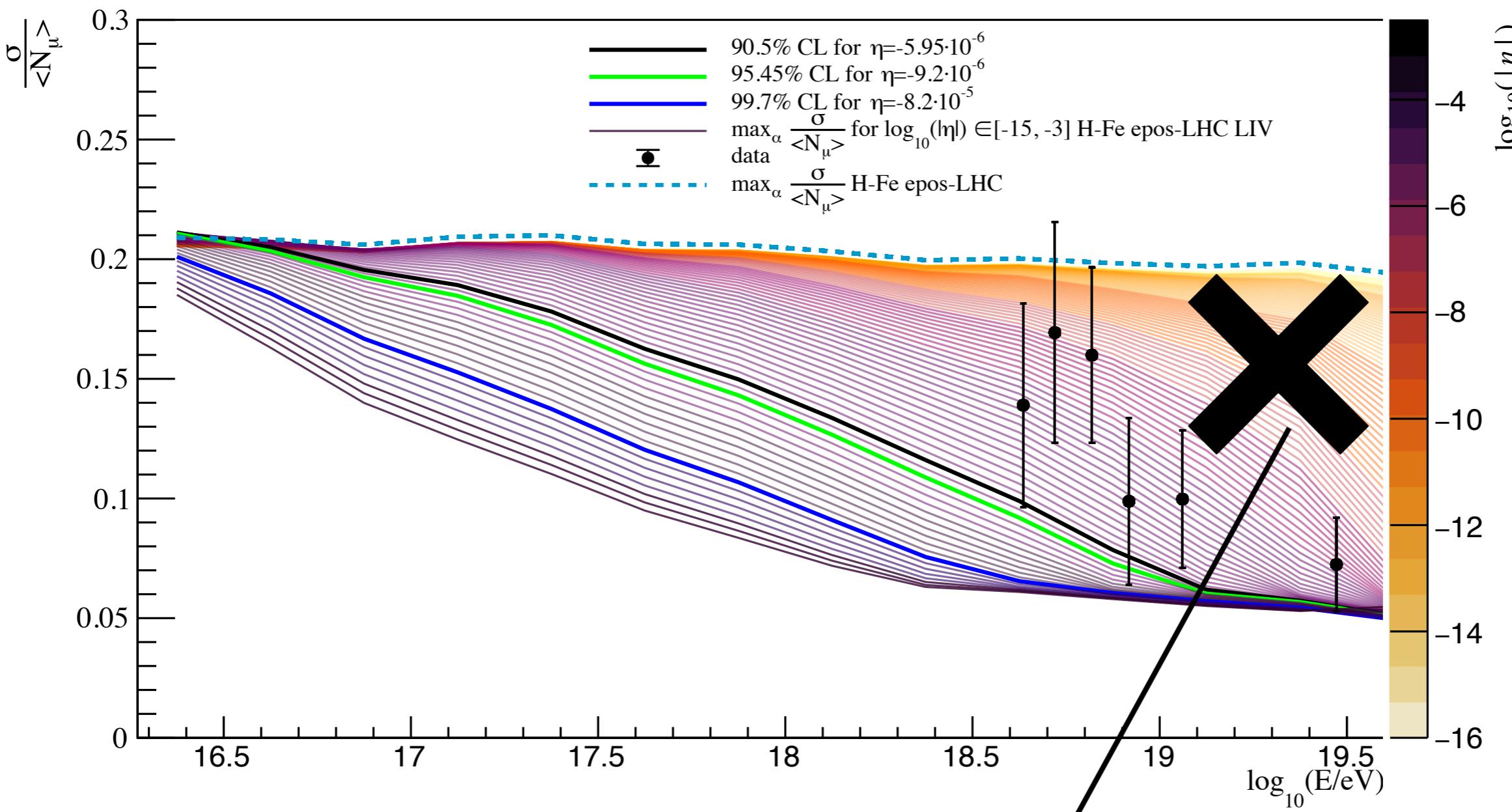
Maximum Mixed Relative Fluctuations

for any LIV parameter value we can calculate the most conservative LIV relative fluctuations as a function of the energy without repeating any shower simulation

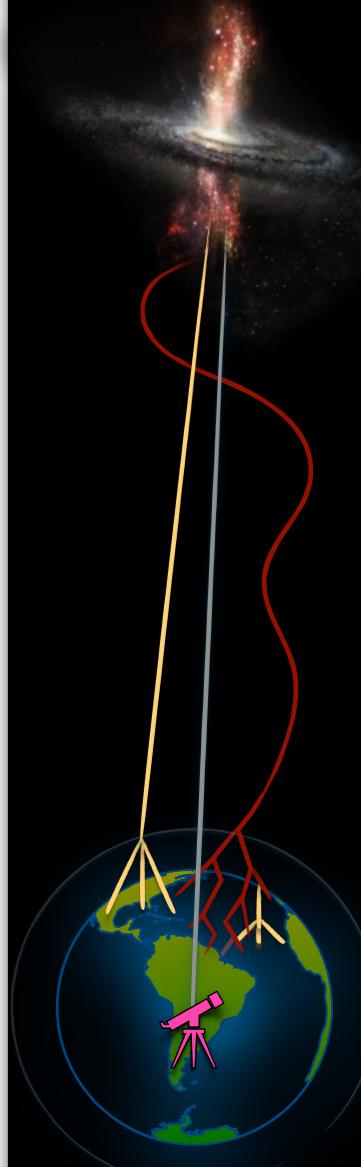
We found $\max_{\alpha} \frac{\sigma_{\mu}}{\langle N_{\mu} \rangle} = \frac{\sqrt{RMS^2(N_{\mu})(\alpha)}}{\langle N_{\mu} \rangle(\alpha)}$ wrt α



Most conservative Mixed Relative Fluctuations

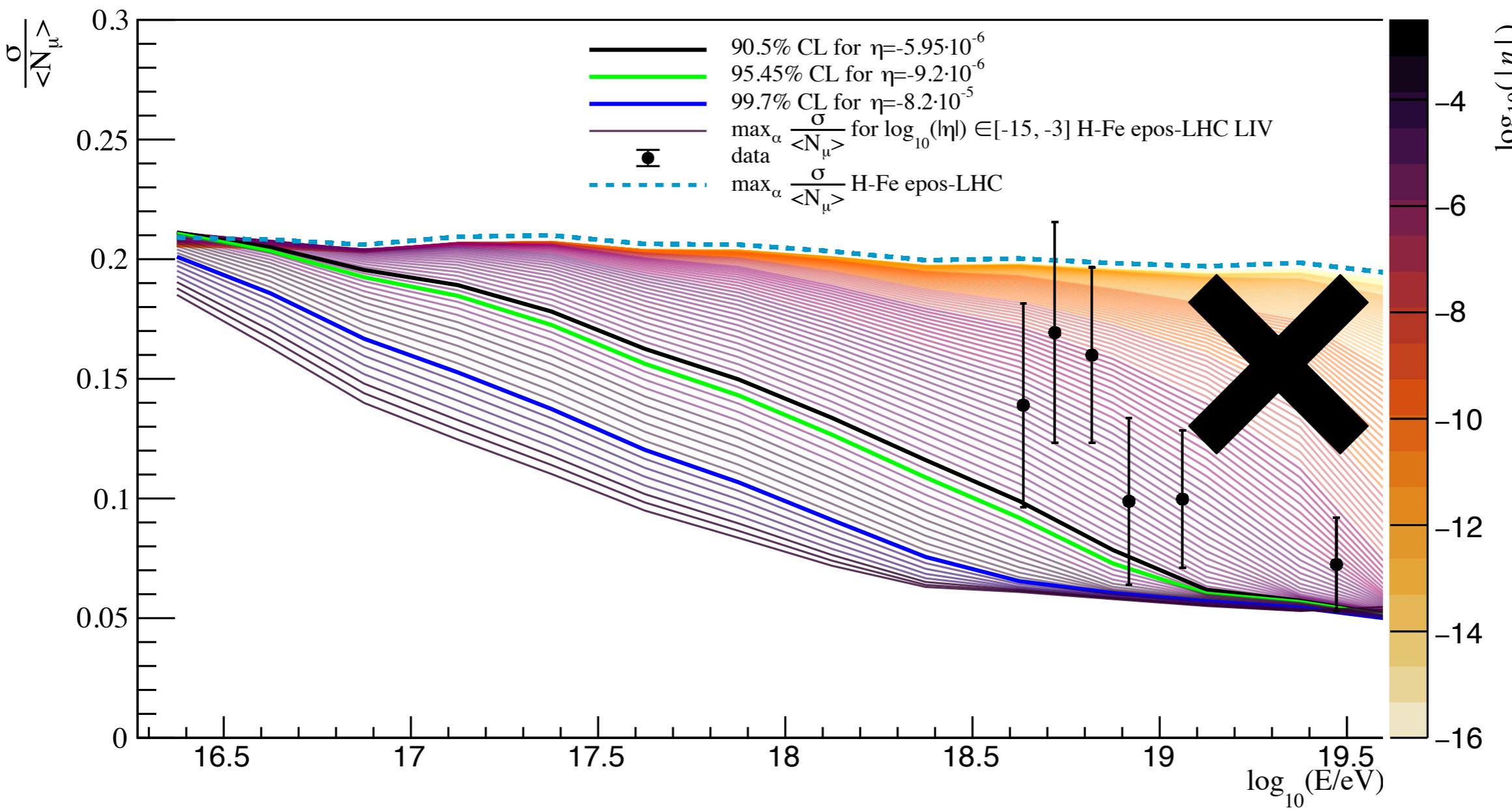


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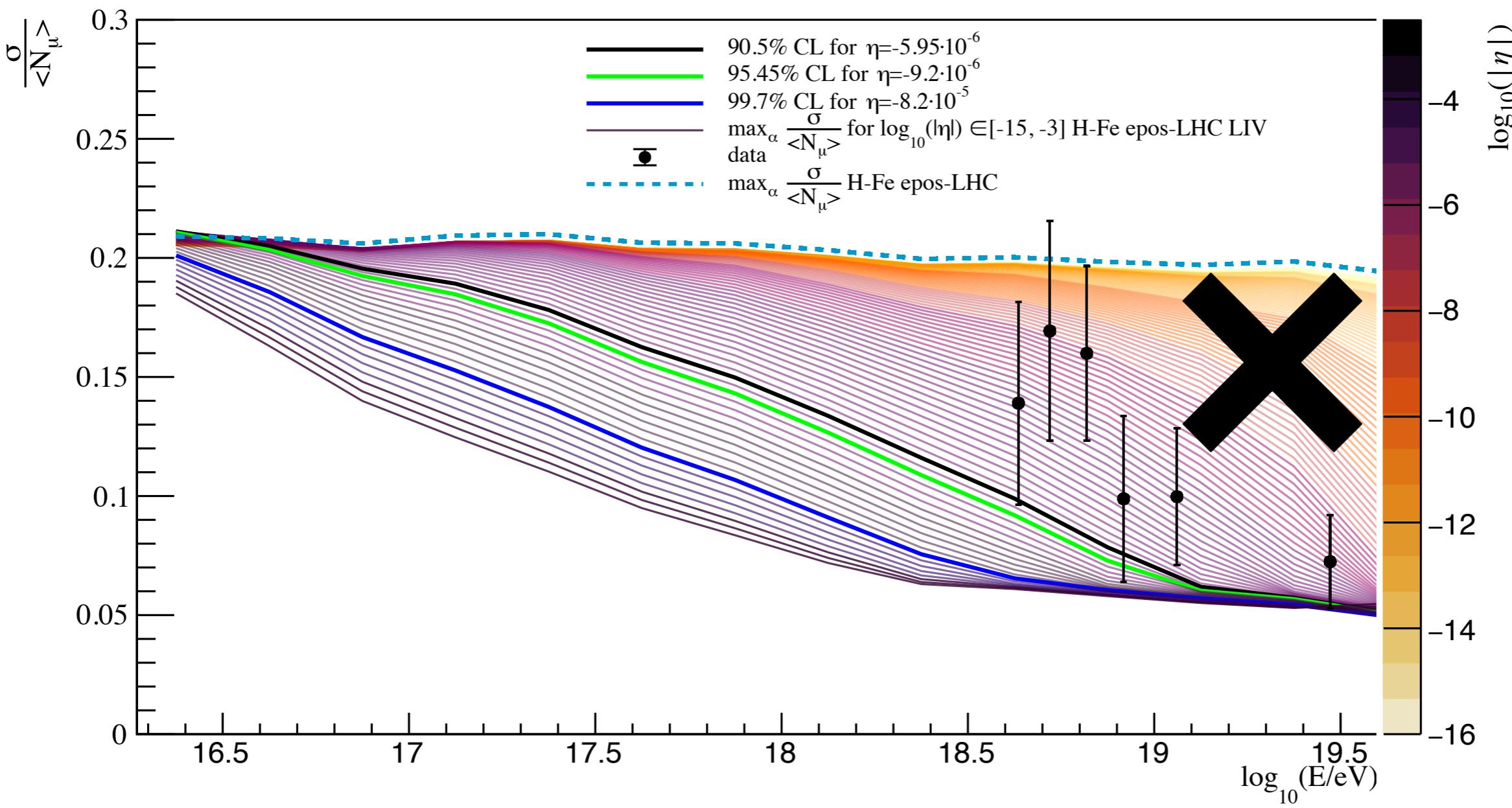
χ^2 is calculated as a function of η considering all the data points



Continuous confidence levels to exclude LIV models

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New bound!
 $\eta^{(1)} > -5.95 \cdot 10^{-6}$ with (90.5% C.L.)

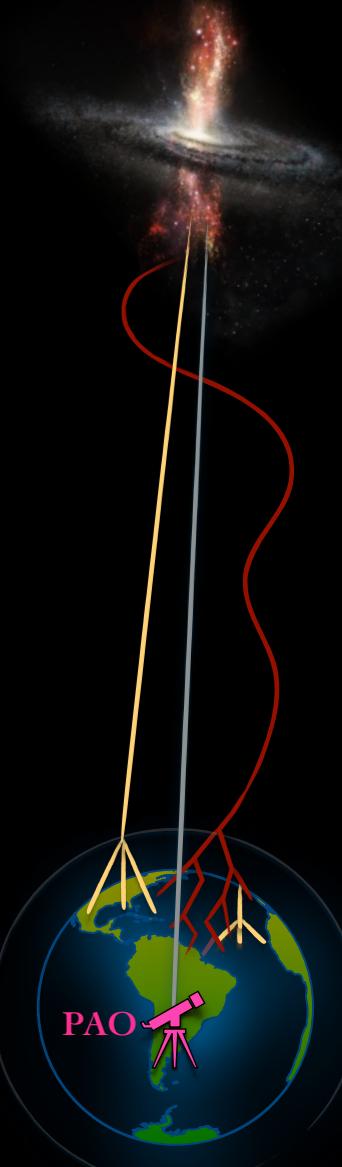
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Summary

- ◆ For the first time LIV effects have been studied considering muon fluctuations;
- ◆ Using the parameterization we obtain the muon fluctuation as a function of energy without shower simulations;
- ◆ We found α that corresponds to the most conservative H-Fe mixed case;
- ◆ Using the parameterization we obtained a new bound for LIV parameter values;

Future prospects

- ◆ A similar approach using the minimum of the relative fluctuation with respect to α could lead to the definition of a negative upper bound of the LIV parameter;
- ◆ We will use the overall procedure for LIV at 2nd order;
- ◆ Limits on η parameter could be found through a combined analysis considering simultaneously muon content distribution and the mass composition derived from the X_{\max} measurements;
- ◆ Future works will involve other hadronic interaction models as QGSJETII-04 and SiBYLL 2.3c
but we expect EPOS-LHC LIV model represents the most conservative case!



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Thank you for your attention!