



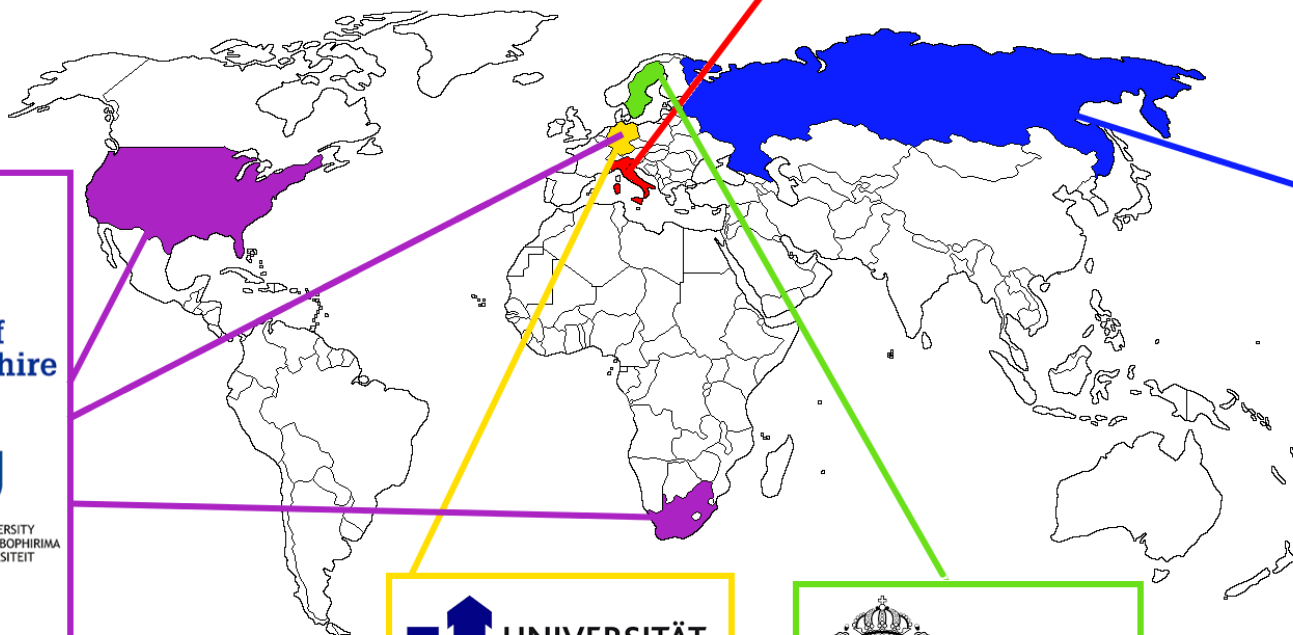
A full solar cycle of proton and helium measurements with the PAMELA experiment

ICRC 2021, 12-23 July 2021

Speaker: *Nadir Marcelli*

PAMELA collaboration

Naples Bari Florence Frascati Rome Trieste CNR, Florence



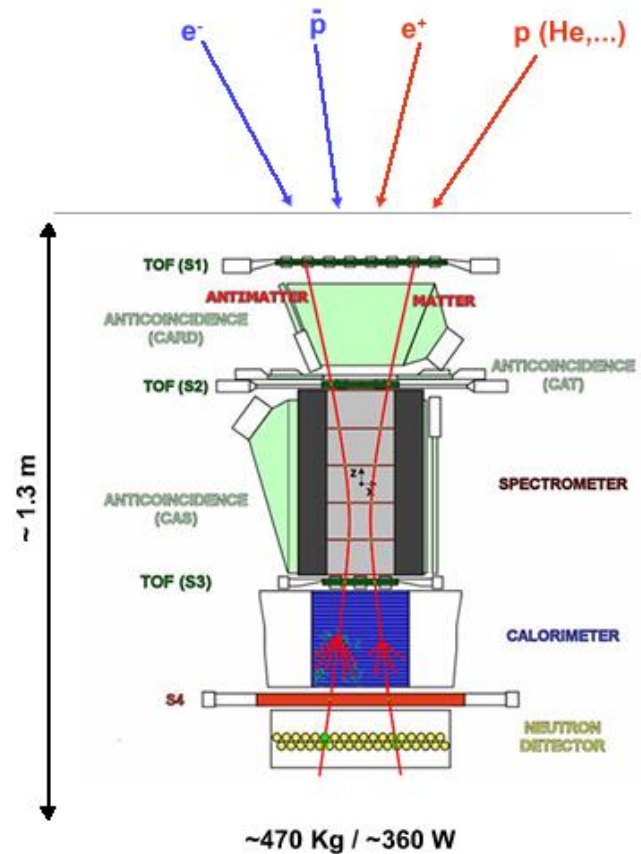
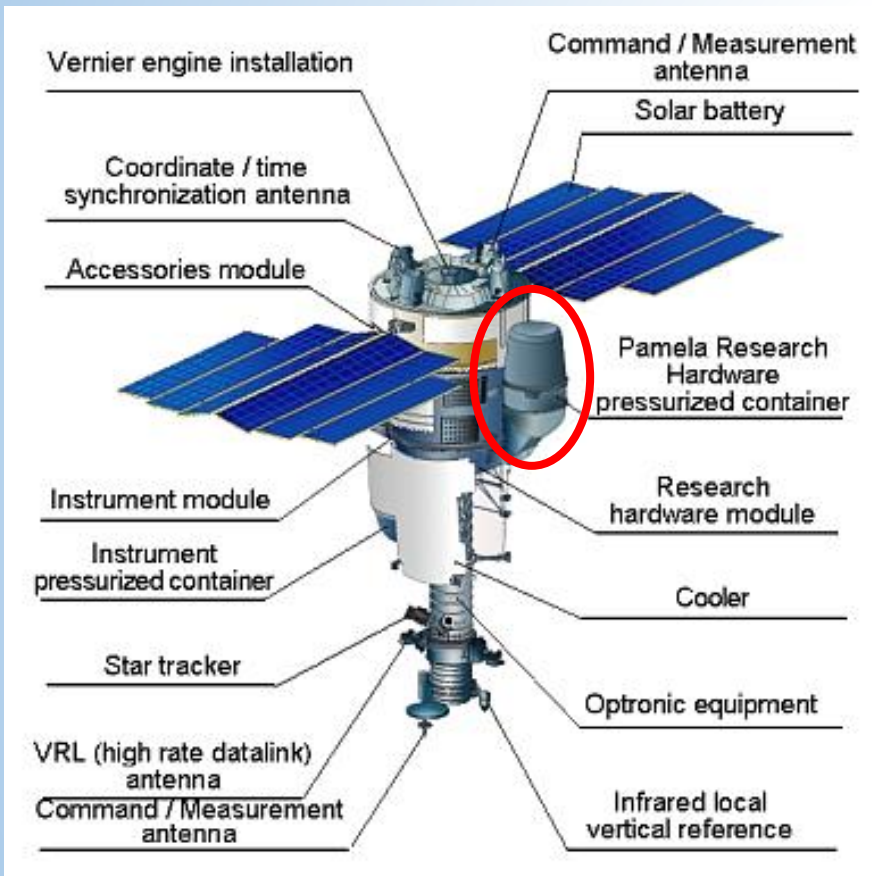
External collaboration

NM STATE UNIVERSITY
NASA FLIGHT CENTER
University of New Hampshire
CIAU
Christian-Albrechts-Universität zu Kiel
NORTH-WEST UNIVERSITY
YUNIBESITI YA BOKONE-BOPHIRIMA
NOORDWES-UNIVERSITEIT

UNIVERSITÄT SIEGEN
Germany

KUNGL. TEKNISKA HÖGSKOLAN
Sweden

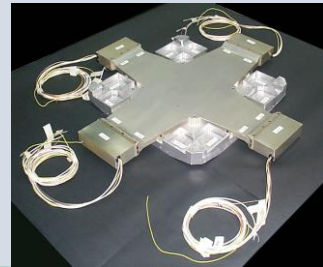
Ioffe Physical Technical Institute
Физический институт имени П.Н. Лебедева
ФИАН
МИАН
Moscow St. Petersburg



Time-Of-Flight

Plastic scintillators:

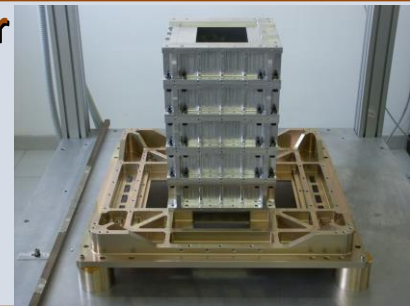
- Trigger
- Velocity
- Charge identification dE/dx



Magnetic spectrometer

Silicon strip + permanent magnet:

- Trajectory track
- Sing and absolute value of the charge



Calorimeter

Silicon Strip and tungsten:

- Electromagnetic shower energy
- Incident $e^{-/+}$ energy



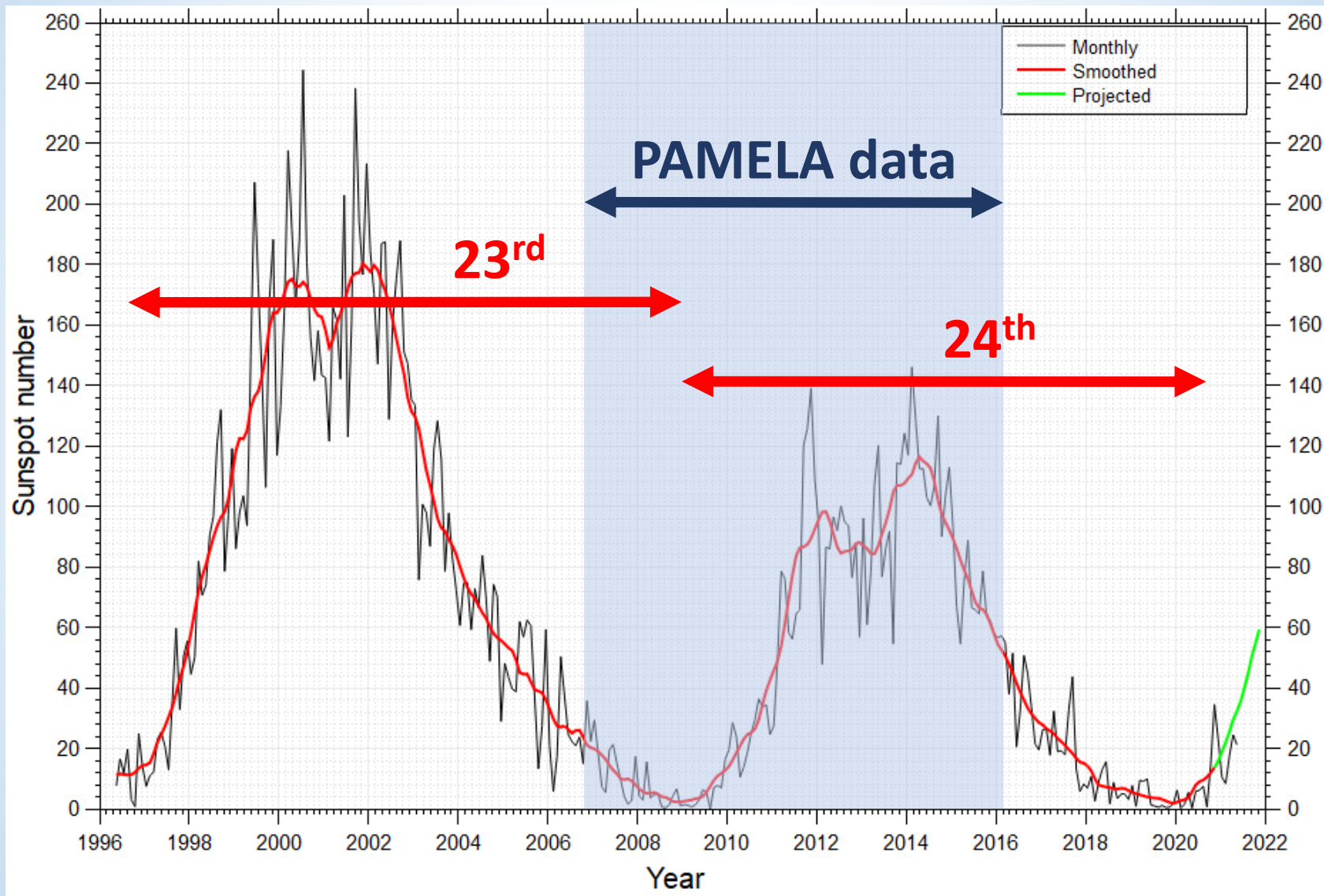
Neutron Detector

^3He gas cylinders:

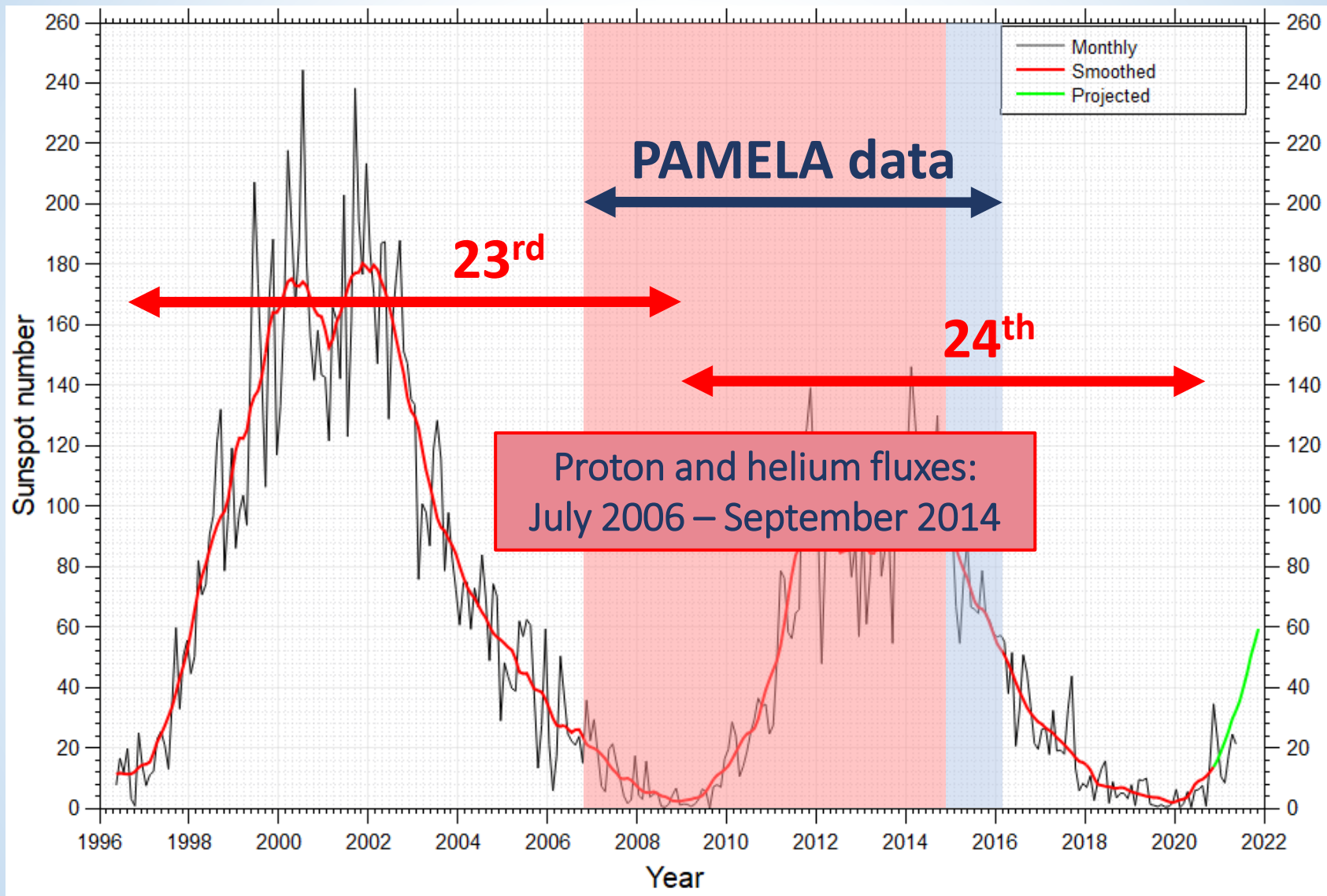
- Adrons and leptons discrimination



Solar activity

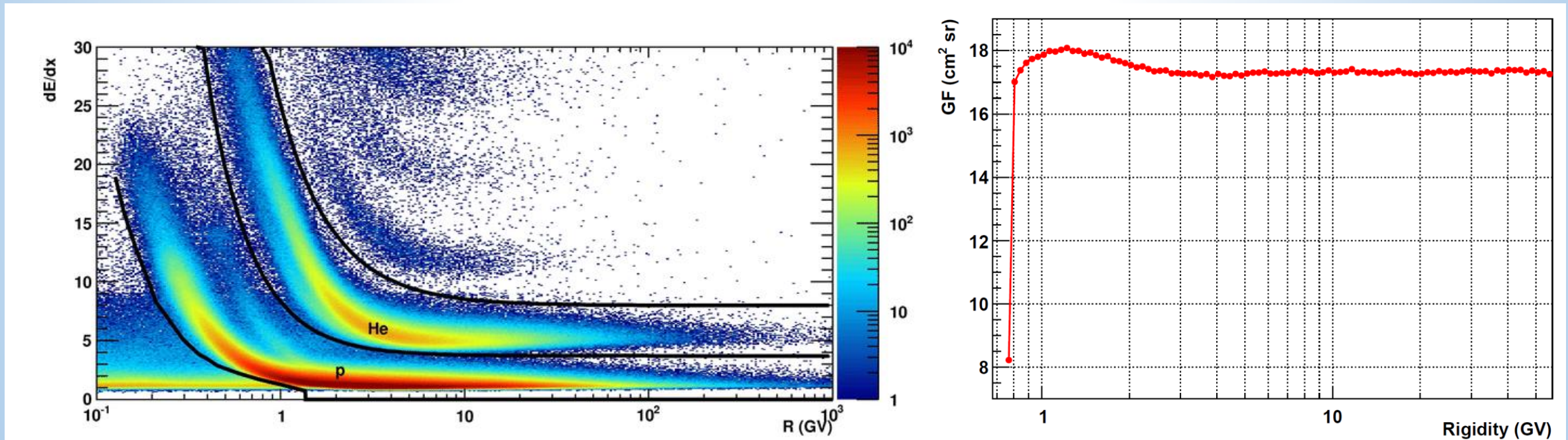


Solar activity



Flux computation

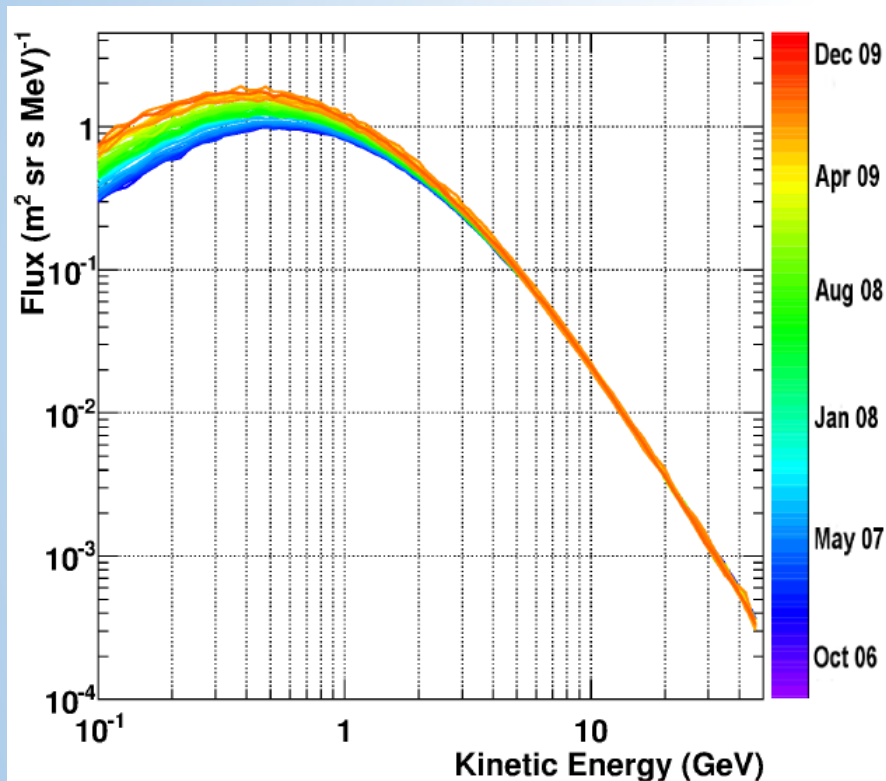
$$\phi(R_i) = \frac{N_{sel}(R_i)}{\Delta R_i} \cdot \frac{1}{GF(R_i) \cdot \varepsilon_{total}(R_i) \cdot T_{live}(R_i)}$$



Proton fluxes

Solar minimum

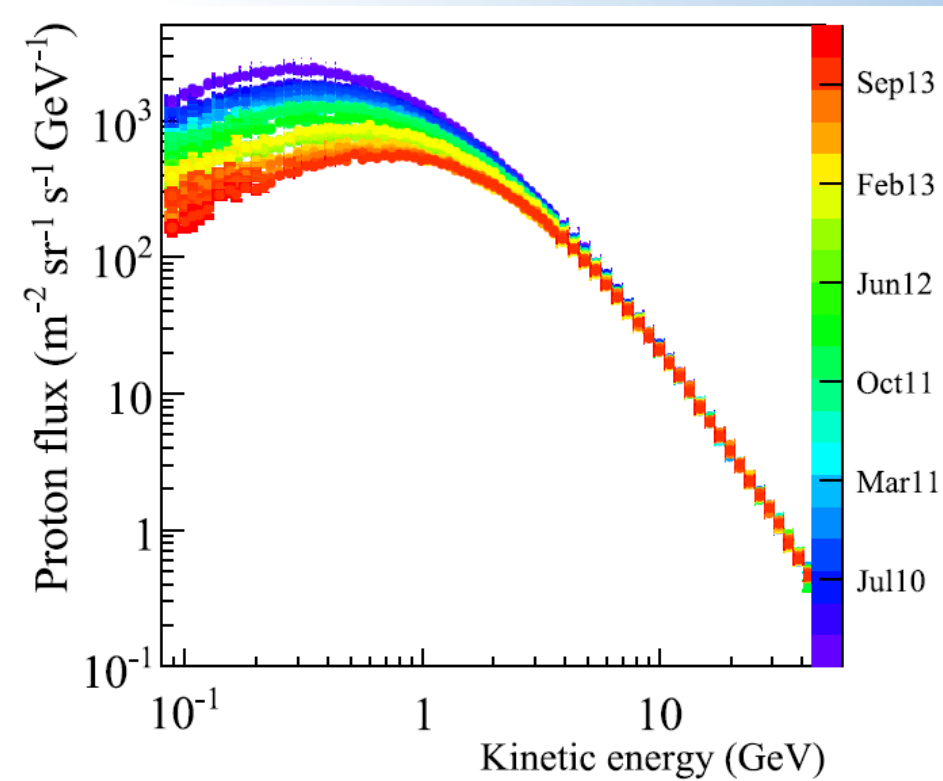
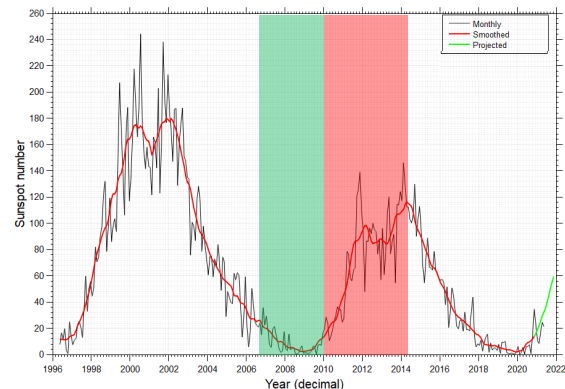
From July 2006 to December 2009



Adriani et al., 2013 - ApJ

Solar maximum

From January 2010 to February 2014

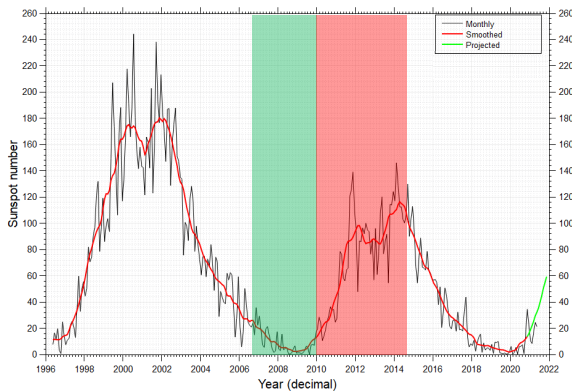
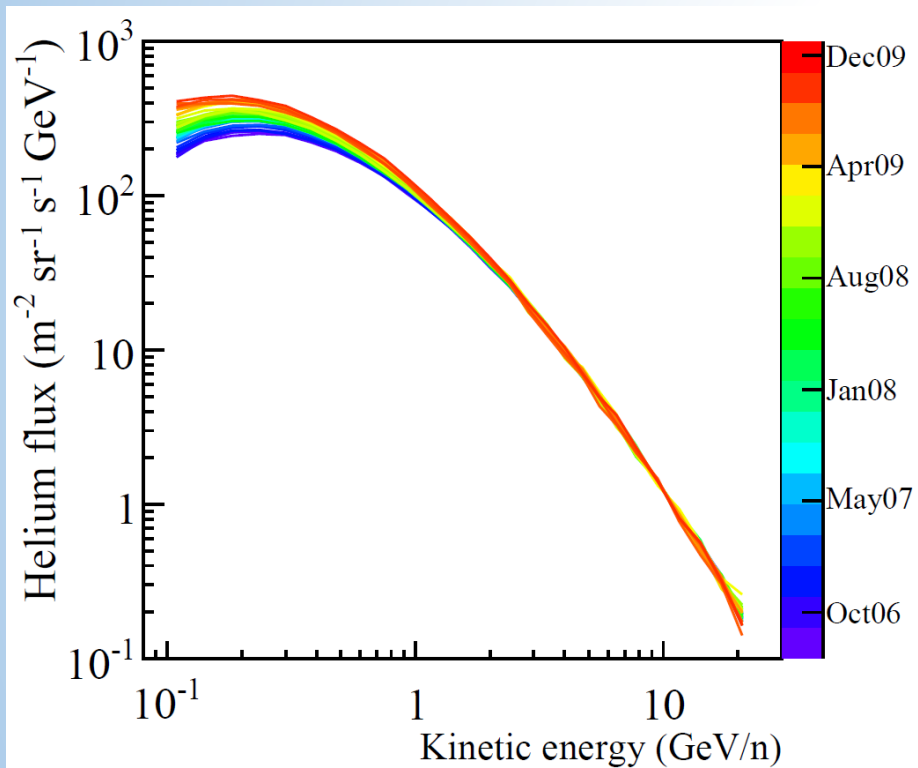


Martucci et al., 2018 - ApJ

Helium fluxes

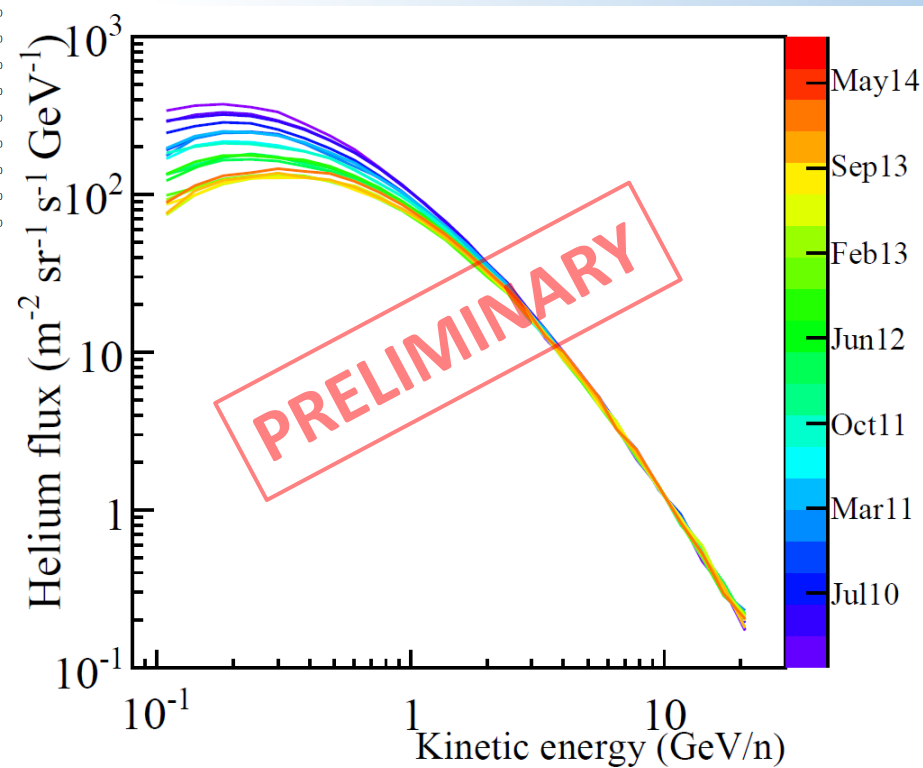
Solar minimum

From July 2006 to December 2009



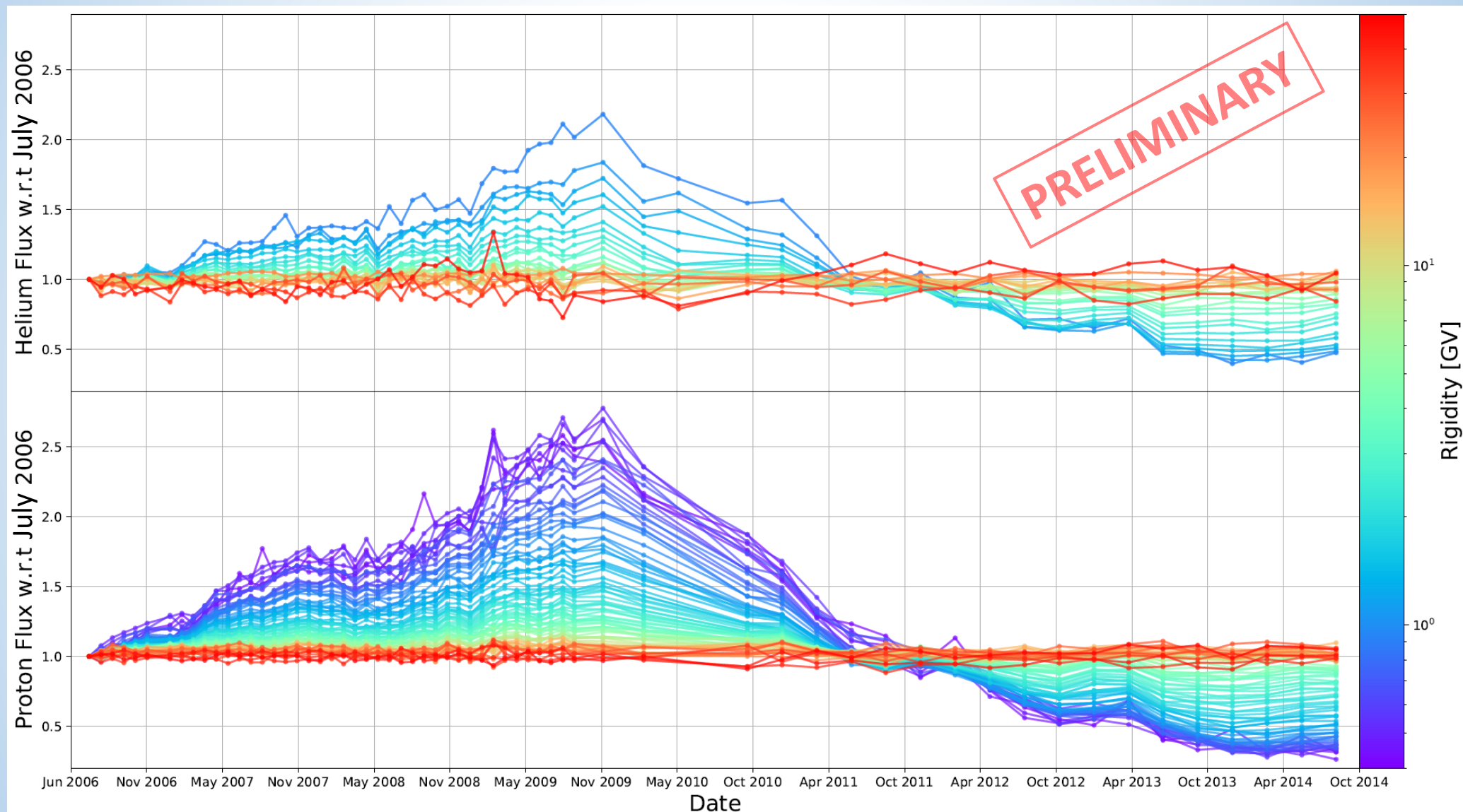
Solar maximum

From January 2010 to September 2014

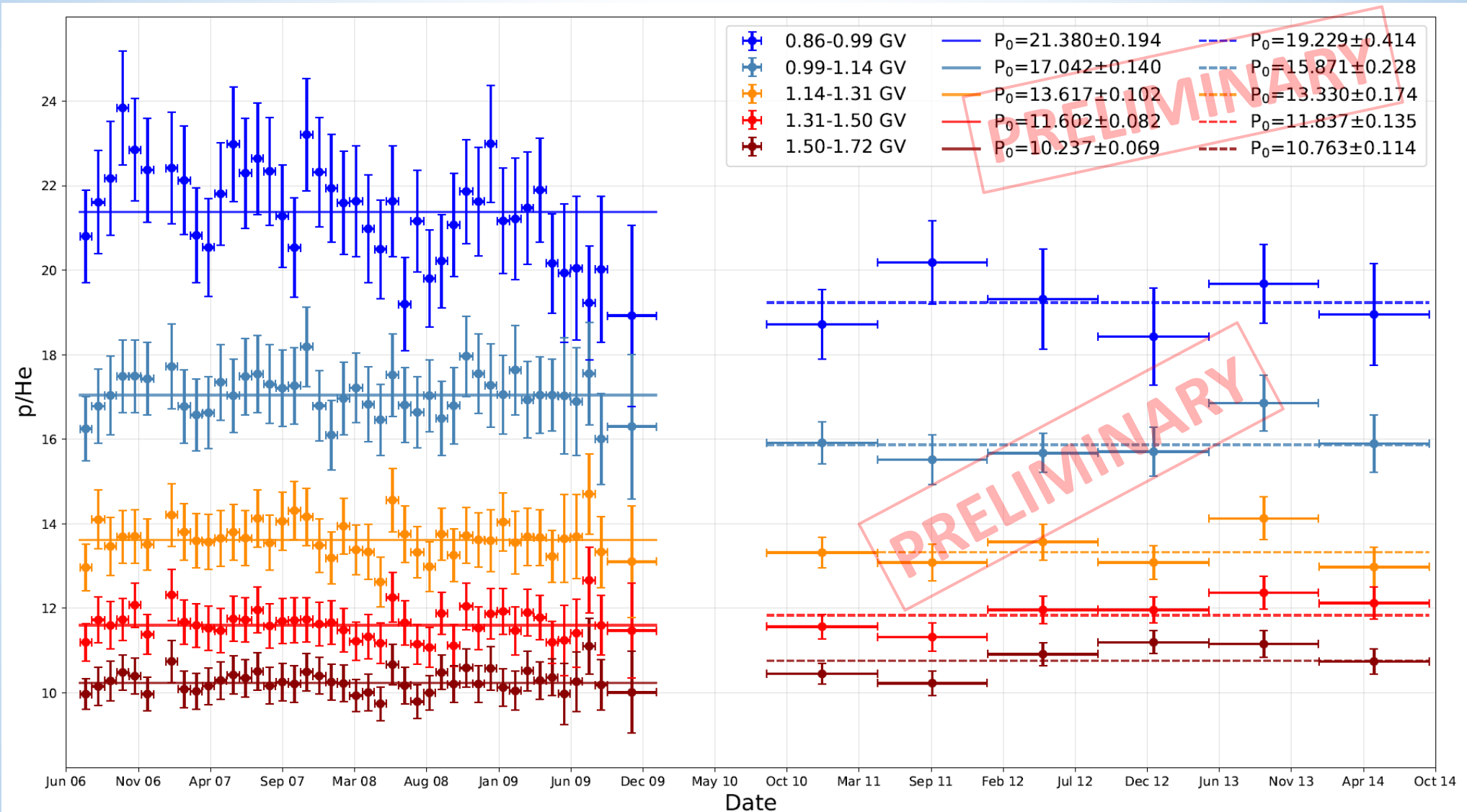


N. Marcelli et al., 2020 - ApJ

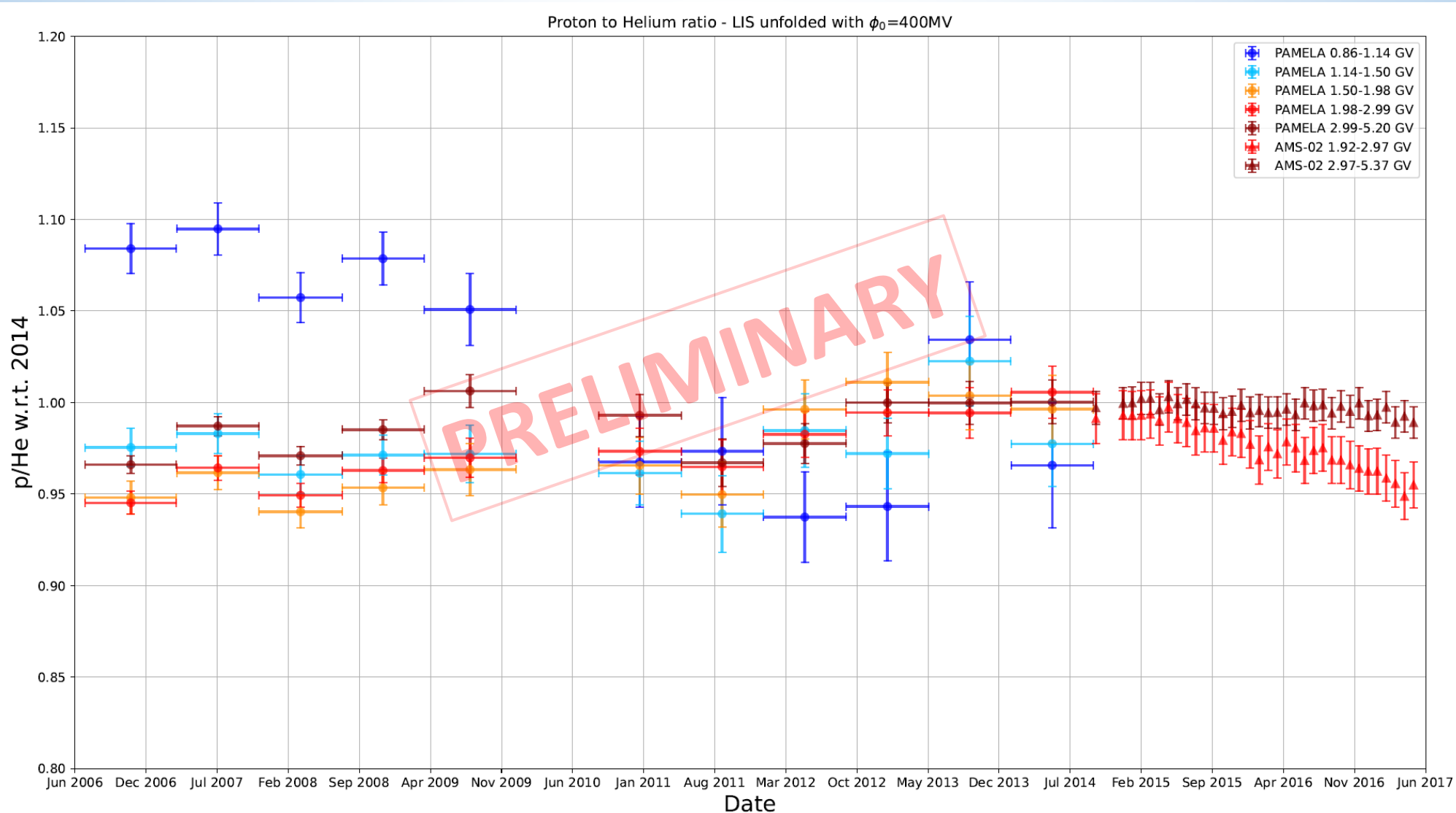
Proton and helium fluxes



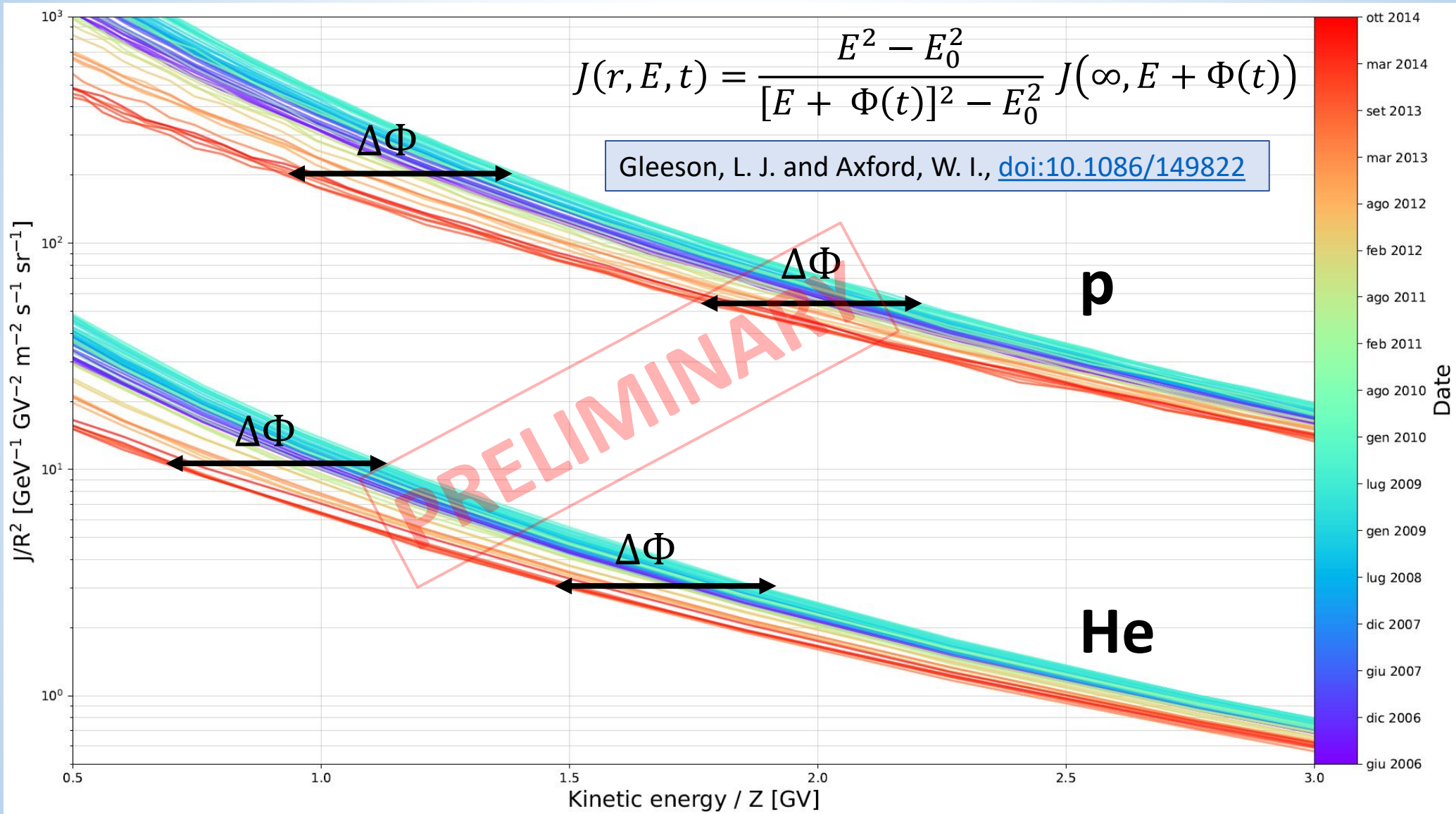
Proton over helium ratio



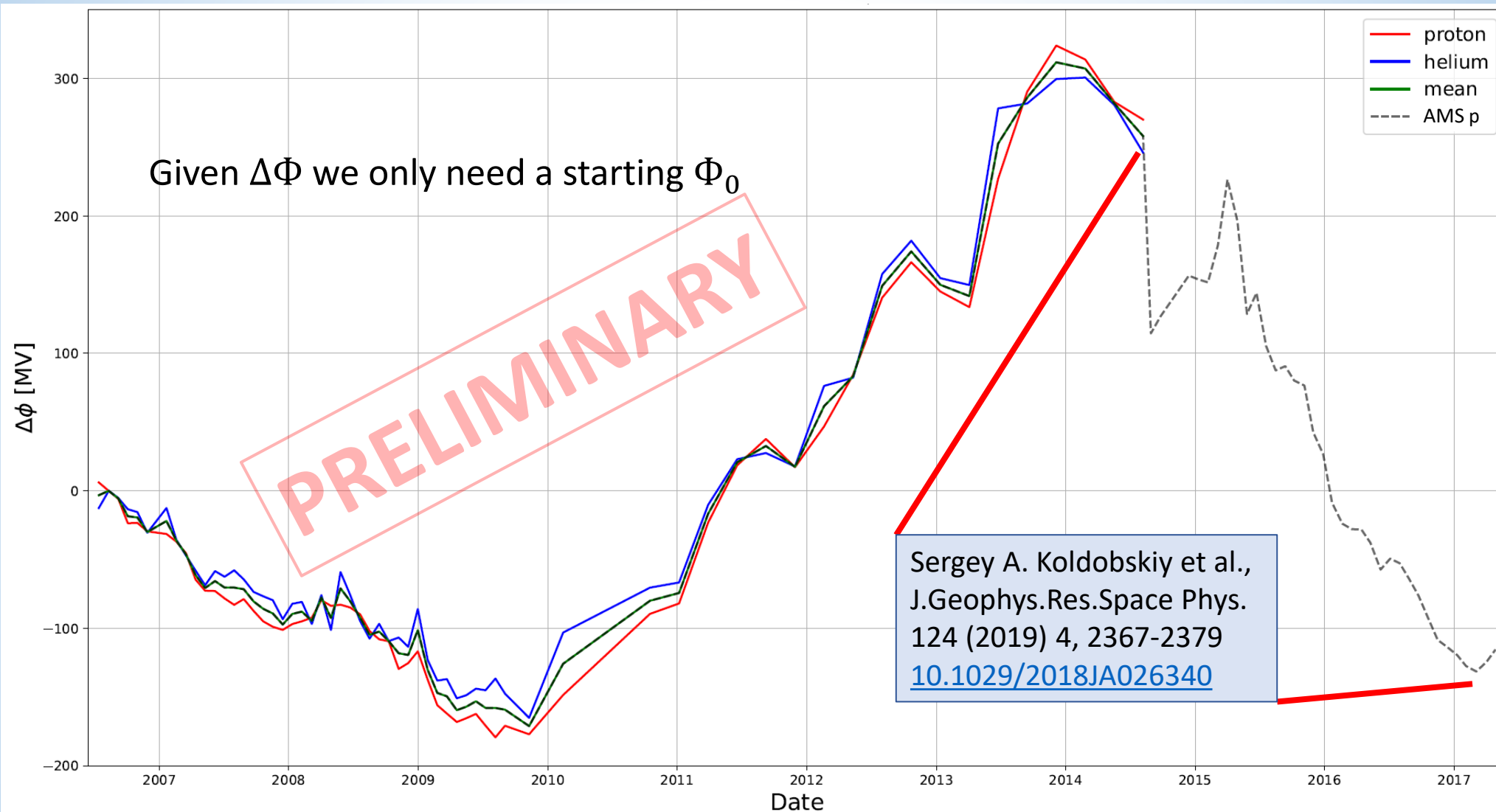
Proton over helium ratio



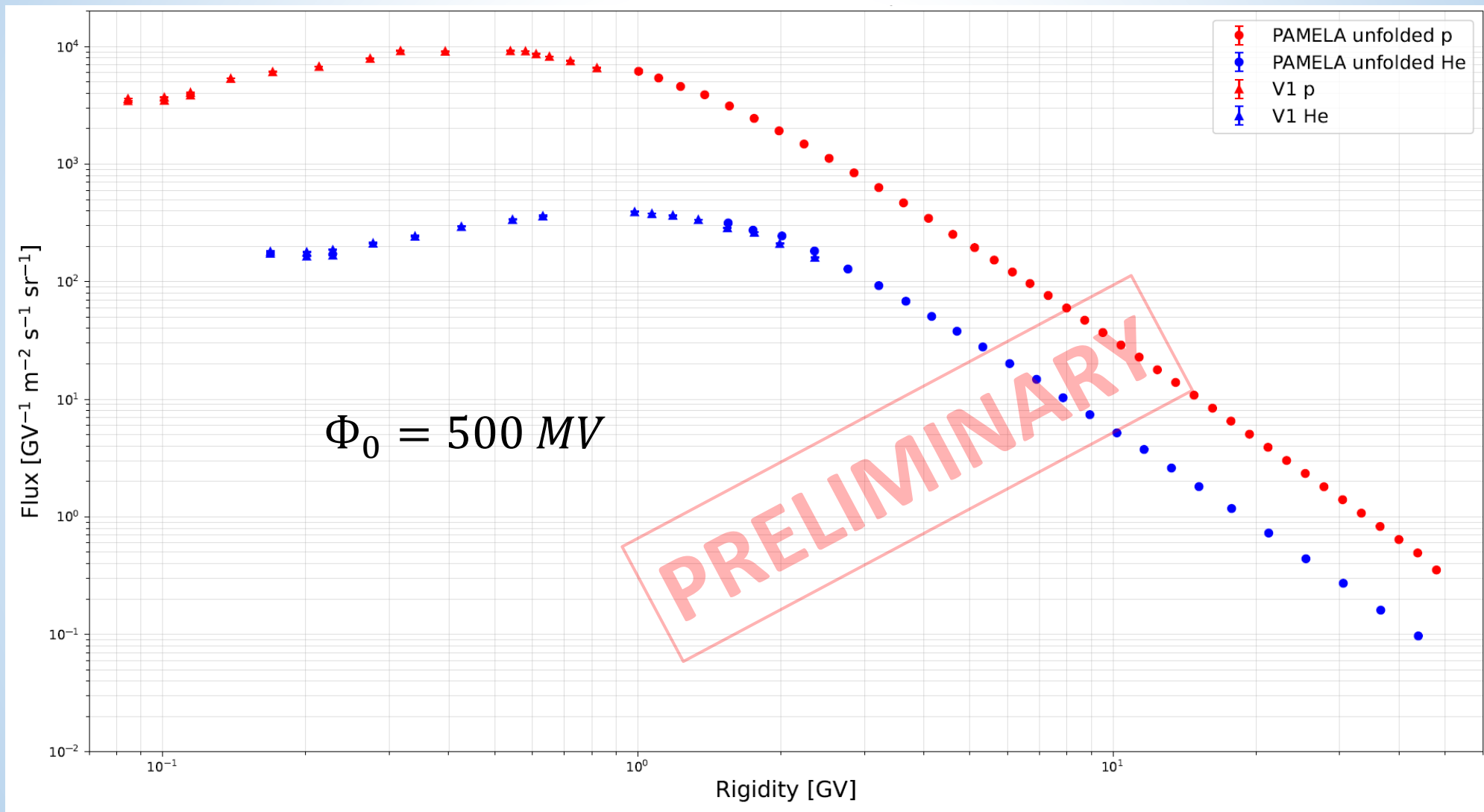
$\Delta\Phi$ from measured fluxes



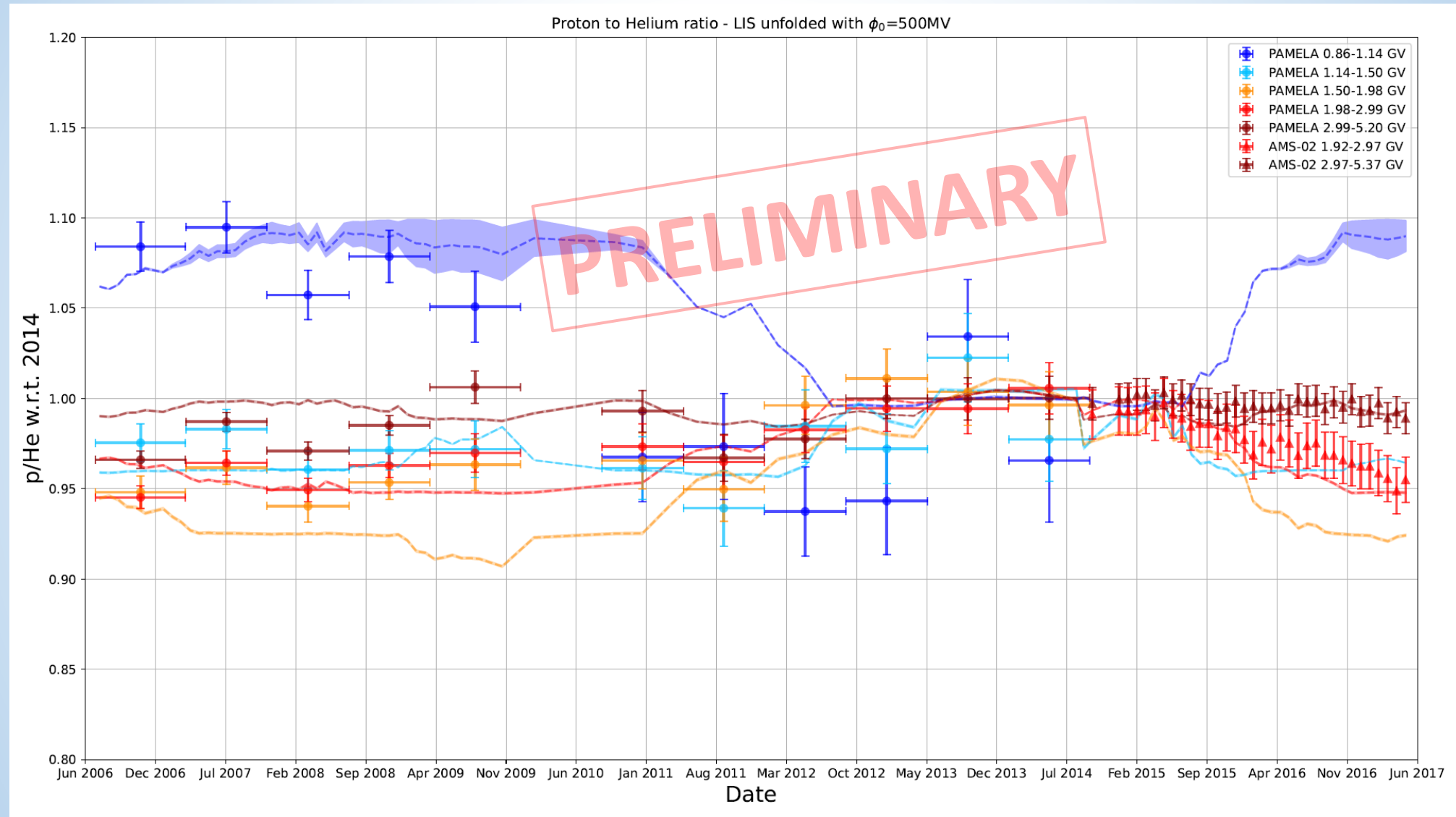
$\Delta\Phi$ from measured fluxes



Unfolded fluxes



Modeled p/He ratio



Conclusion

- The spectra for proton and helium-nuclei measured by the PAMELA experiment over a nearly complete solar cycle (July 2006 – September 2014) were presented
- An indication of time dependence in the p/He ratio is observed, suggesting either a difference in the processes involved in solar modulation effects and/or a difference in the LIS for the two species
- The force-field approximation for solar modulation was applied to relate the observed dependencies to the different shapes of the local interstellar proton and helium-nuclei spectra
- More accurate models will be developed to have a complete picture of the cause of these dependencies