## Latest Results from the Daya Bay Experiment

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#### The Daya Bay Collaboration

#### ~200 Collaborators from 41 institutions

#### Asia (24):

Beijing Normal Univ., CGNPG, CIAE, Congqing Univ., Dongguan Univ. Tech., ECUST, GXU, IHEP, Nanjing Univ., Nankai Univ., NCEPU, NUDT, Shandong Univ., Shanghai Jiao Tong Univ., Shenzhen Univ., Tsinghua Univ., USTC, Xian Jiaotong Univ., Zhongshan (Sun Yat-sen) Univ., Chinese Univ. of Hong Kong, Univ. of Hong Kong, National Chiao Tung Univ., National Taiwan Univ., National United Univ. **Europe (2):** Charles University, JINR Dubna

#### North America (15):

Brookhaven Natl Lab, Illinois Institute of Technology, Iowa State, Lawrence Berkeley Natl Lab, Princeton, Siena College, Temple University, UC Berkeley, Univ. of Cincinnati, Univ. of California Irvine, UIUC, Univ. of Wisconsin, Virginia Tech, William & Mary,<sup>2</sup> Yale

#### The Daya Bay Experiment Overview Daya Bay

- The experiment is located in south China in proximity to Daya Bay & Ling Ao nuclear power plants
- 6 nuclear reactors serve as an intense and pure source of electron antineutrinos
- 8 functionally identical antineutrino detectors (ADs) located in 3 experimental halls (EHs)
- Placed underground for suppression of cosmic rays



Far hall location optimized in order to measure neutrino oscillation at the ~2 km baseline 3



### Antineutrino Detector (AD)

- Each AD consists of 3 nested cylindrical regions:
  - 20 t of liquid scintillator doped with gadolinium (GdLS)
  - 22 t of pure liquid scintillator (LS)
  - 40 t of mineral oil (MO)
- Scintillation light collected by 192 x 8" PMTs
- ADs are submerged in instrumented water pool passive shielding and active muon detector (veto)





#### **Antineutrino Detection**

- Inverse beta decay (IBD):
  - Prompt signal: e<sup>+</sup> loses energy and annihilates
  - Delayed signal: *n* thermalizes and is captured on Gd (nGd) or H (nH) – two independent data sets
  - Temporal (and spatial) correlation of prompt and delayed signal leads to great suppression of background
  - Antineutrino energy can be deduced from prompt energy:





- Looking at neutrino flux as a function of distance and energy
- Disappearance of some  $\overline{\nu}_{e}$ 's due to neutrino oscillation:

$$P(\overline{\nu}_e \to \overline{\nu}_e) = 1 - \frac{\sin^2 2\theta_{13} \left(\cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E}\right)}{\text{Short baseline}} - \frac{\sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \frac{\Delta m_{21}^2 L}{4E}}{\text{Medium baseline}}$$

• Daya Bay optimized for short baseline measurement of  $\theta_{13}$  where:

$$P(\overline{\nu}_e \to \overline{\nu}_e) \simeq 1 - \sin^2(2\theta_{13})\sin^2\left(\frac{\Delta m_{ee}^2 L}{4E}\right)$$

$$\sin^2 \frac{\Delta m_{ee}^2 L}{4E} \simeq \cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \frac{\Delta m_{32}^2 L}{4E}$$



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### Oscillation Measurement with nGd

• Measurement of  $\sin^2 2\theta_{13}$  with world-leading 3.4% precision:

 $\sin^2(2\theta_{13}) = 0.0856 \pm 0.0029$ 

 Measurement of Δm<sup>2</sup><sub>ee</sub> with 2.8% precision, comparable to accelerator experiments:

$$\Delta m^2_{ee} = (2.522^{+0.068}_{-0.070}) \times 10^{-3} \,\mathrm{eV}^2$$

 $\Delta m_{32}^2 = (2.471^{+0.068}_{-0.070}) \times 10^{-3} \,\text{eV}^2 \text{ (Normal mass ordering)}$  $\Delta m_{32}^2 = -(2.575^{+0.068}_{-0.070}) \times 10^{-3} \,\text{eV}^2 \text{ (Inverted mass ordering)}$ 

• Statistics contribute 60% (50%) to the total uncertainty in the  $\sin^2 2\theta_{13}$  ( $\Delta m_{ee}^2$ ) measurement.

World's largest data set of of almost 4 million antineutrino candidates collected (1958 days)



[Phys. Rev. Lett. 121, 241805 (2018)]

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### Oscillation Measurement with nH

- Largely independent measurement from nGd:
  - Different statistics
  - Mostly independent (and larger) systematics
  - Larger background
- Rate-only analysis:

 $\sin^2(2\theta_{13}) = 0.071 \pm 0.011$ 

- Based on 621 days data set
- Work on rate & shape analysis is in progress

It will likely be one of the leading measurements of  $\sin^2 2\theta_{13}$  in the world



[Phys. Rev. D 93, 072011 (2016)]

## Global Comparison: $\sin^2 2\theta_{13} \& \Delta m_{ee}^2$





- Motivated (among other things) by LSND & MiniBooNE observed excess of  $v_e$  ( $\overline{v}_e$ ) in  $v_\mu$  ( $\overline{v}_\mu$ ) beam inconsistent with 3v oscillation
- Daya Bay results consistent with  $3\nu$  model, no additional oscillation driven by  $\Delta m_{41}^2$  and  $\sin^2 2\theta_{14}$  observed (1230 days)



Combined analysis with Bugey-3, MINOS, MINOS+ excluded LSND and MiniBooNE 99% C.L. allowed regions at 99% CL<sub>s</sub> for  $\Delta m_{41}^2 < 1.6 \text{ eV}^2$ 

[Phys. Rev. Lett. 125, 07180 (2020)] <sup>10</sup>



- Flux: Daya Bay consistent with previous experimental results, but  $\sim$ 1.8  $\sigma$  below the best prediction up to date (Huber-Mueller model)
- Spectrum: comparison with Huber-Mueller model shows a total >5 $\sigma$  deviation, especially significant in 4-6 MeV 'bump' region (>6 $\sigma$ )





- Reactor  $\overline{\nu}_{e}$  come from fission of 4 main isotopes: <sup>235</sup>U, <sup>238</sup>U, <sup>239</sup>Pu, <sup>241</sup>Pu
- Clear change of total  $\overline{v}_{e}$  flux with reactor fuel composition observed





Result points to <sup>235</sup>U being mainly responsible for the deficit of observed  $\overline{\nu}_{e}$ 's compared to Huber-Mueller model



[Phys. Rev. Lett. 118, 251801 (2017)]



- Individual prompt spectra of <sup>235</sup>U and <sup>239</sup>Pu extracted using the evolution of the prompt spectrum as a function of the isotope fission fractions
  - First measurement of <sup>235</sup>U and <sup>239</sup>Pu spectra from a commercial reactors
- Joint analysis with the PROSPECT experiment done
  - <sup>235</sup>U spectra consistent
  - <sup>235</sup>U spectrum obtained with better precision
  - Correlation between <sup>235</sup>U and <sup>239</sup>Pu spectra reduced



[arXiv:2006.15386] [Phys. Rev. Lett. 123, 111801 (2019)]

# $\overline{\nu}_{e}$ 's Associated with Gravitational Waves

- No significant  $\overline{\nu}_e$  event excess within ±10/500/1000 s time window of gravitational wave event observed
  - → GW150914
  - → GW151012
  - → GW151226
  - → GW170104
  - → GW170608
  - → GW170814
  - → GW170817
  - Upper limits (90% C.L.) on  $\overline{\nu}_e$  fluence determined:





• The latest Daya Bay reactor neutrino oscillation results:

 $\sin^2(2\theta_{13}) = 0.0856 \pm 0.0029$ 

$$\Delta m_{ee}^2 = (2.522^{+0.068}_{-0.070}) \times 10^{-3} \,\mathrm{eV}^2$$

- Daya Bay imposed the most stringent limits on sterile neutrino mixing angle  $\sin^2 2\theta_{14}$  for  $\Delta m^2_{41} < 0.2 \text{ eV}^2$ 
  - Together with MINOS/MINOS+ excluded most of LSND / MiniBooNE allowed region
- Reactor antineutrino flux and spectrum measured as well as individual yields and spectra of <sup>235</sup>U and <sup>239</sup>Pu, joint analysis with PROSPECT
- No  $\bar{\nu}_{e}$  excess associated with gravitational waves observed, limits on  $\bar{\nu}_{e}$  fluence placed
- Daya Bay finished data taking at the end of 2020
  - Final results will include the best determination of  $\sin^2 2\theta_{13}$  for the foreseeable future (uncertainty < 3%) and more



### Thank you for your attention!