



UNIVERSITÀ DEGLI STUDI DI TORINO

Damping of self-generated Alfvén waves in a partially ionized medium and the grammage of cosmic rays in the proximity of supernova remnants

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Overview

- CRs escaping from SNRs
- excitation of Alfvén waves by resonant streaming instability
 - * suppression of diffusion coefficient in the source region
 - self-confinement
- damping of Alfvén waves
 - ion-neutral friction
 - turbulent damping
 - non-linear Landau damping
- CR grammage accumulated in the source region?
- Need a revision of the standard picture of CR Galactic propagation?

CR escape from SNRs: a challenge

- a self consistent theory of acceleration and escape of CRs is still missing
 - time-dependent problem
 - active role of CRs non linearity
 - broad range of spatial scales and propagation regimes involved
 - difficult with simulations

SOURCE

- strong B amplif. $\delta B \sim B_0$
- Bohm diffusion $\lambda \sim R_L$
- $P_{CR} > P_{B}$



<u>ISM</u>

- $\delta B \ll B_0$
- $\lambda \gg R_L$
- $P_{CR} \sim P_{B}$

CR resonant streaming instability

- CR density gradient
 - streaming CRs transfer momentum to waves
 - \star generate resonant waves $k \sim 1/R_L$
 - CRs scatter on self-generated waves, more effective diffusion
- Geometry of CR propagation
 - ID flux-tube
 - 3D isotropic

X

affects the CR gradient



Wave damping

Ion-neutral friction

- ionization fraction of the ISM
- * species of the colliding ion and neutral

Turbulent damping

- interaction with counter-propagating Alfvén waves
- per-existing Alfvénic turbulence
- Non-linear Landau damping
 - interaction of background thermal ions with the beat of two Interfering Alfvén waves

Focus on WIM and WNM, which have a total filling factor in the ISM ~ 50%

Wave damping

WIM and WNM: H ions and H, He neutrals

f = ionization fraction $\chi = He fraction$

ion-neutral damping turbulent damping (FG)



Escape radius and age

Half-time of the CR cloud

- cloud of initial radius R
- t_{1/2} is the time after which half of the CRs has escaped the initial cloud of radius R
- t_a is the age of a SNR of a given radius
- * the escape radius is taken such that $t_{1/2}(R) = t_a(R)$
- t_{1/2} > t_a: SNR expansion is faster that CR cloud expansion
- $\mathbf{t}_{1/2} < \mathbf{t}_a$: SNR expansion is slower that CR cloud expansion
- R_{esc}(E), T_{esc}(E) are typically decreasing functions of the particle energy

Suppression of D

- t~t_{1/2}: CR overdensity at small radii
- t >> t_{1/2}: solutions approaches test particle
- t_{1/2}: timescale over which waves can grow, CRs confined
- CR overdensity for t >> t_{1/2}
- important suppression of the CR diffusion coefficient
- R_{esc}(E) and T_{esc}(E) decrease with energy
- high energy particles are less confined and escape/diffuse faster

Results: escape time and radius

Strong dependence on the ISM phase due to ion-neutral damping

- f = ionization fraction
- $\chi = He fraction$

High energy particles are less confined...



Results: suppression of D

WNM: spectrum and D suppression at the shock and at 50pc from the SNR at different ages



Results: suppression of D

WIM: spectrum and D suppression at the shock and at 50pc from the SNR at different ages



Results: residence time and grammage

- damping processes, especially ion-neutral damping
 - * there is an effect of self-generated waves on the confinement time
 - but the resulting source grammage is found to be negligible compared to observations
 - region of ~ 100pc around the source



Conclusions

- the escape of CRs from sources is still an open issue
- CR propagation in the source region can be highly non-linear
 - Streaming instability and suppression of D
 - Limitations due to damping mechanisms
 - * dependence on the ISM phase in which the SNR is embedded
- CR source grammage
- possibility of producing secondaries in the source region
- implications for the observed CR spectrum
- interpretation of gamma-ray data...

We find a negligible source grammage in the WNM and WIM