Effects of galactic magnetic field on the UHECR anisotropy studies

Ryo Higuchi (ICRR, the University of Tokyo)

25

scale [deg]

UHECRs events (w/o coherent deflection by GMF)



true value

North

South

Correlation between

- events(w/o GMF)
- flux (w/o GMF)
- When we do not consider the coherent deflection by GMF, estimated parameters are centered at input value in any datasets.

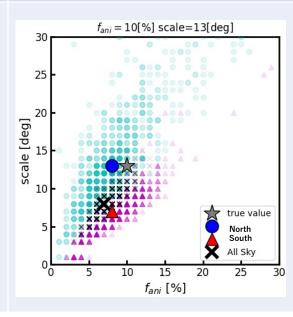
Va

Model CR flux pattern (on the earth) F_{earth}

Questions:

- 1. How much bias in the parameter estimation w/ and w/o GMF, North and South?
- 2. Can we reduce the bias by considering GMF effect in the analysis?
- =>To investigate the GMF effect, we conduct likelihood analysis with mock events & model flux for these patterns:
 - A. events w/o GMF & flux w/o GMF
 - B. events w/ GMF & flux w/o GMF
 - C. events w/ GMF & flux w/ GMF

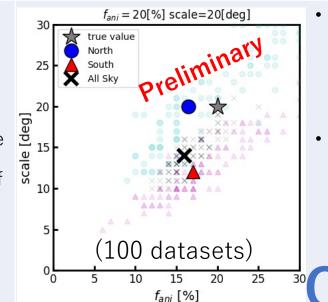
UHECRs events (w/ coherent deflection by GMF)



f_{ani} [%]

 $f_{ani} = 10[\%] \text{ scale} = 13[\text{deg}] \text{ (NO GMF)}$

- Correlation between
 - events(w/ GMF)
 - flux (w/o GMF)
- When we consider the GMF effect only for the events, there would be systematic decrease of estimated parameters in south-sky & all-sky datasets.



- Correlation between
 - events(w/ GMF)
 - flux (w/ GMF)
- When we consider the GMF effect for both events & flux, we may reduce the systematic decrease of anisotropic fraction due to GMF.