

MWL observations in 2019-2020 of a new VHE γ -ray emitter: the FSRQ QSO B1420+326

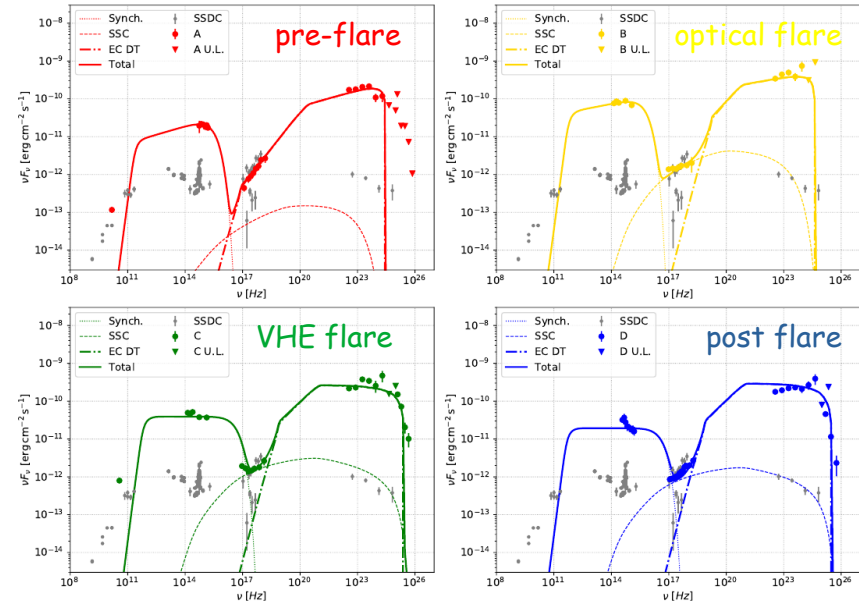
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MWL campaign allow us to trace the broadband emission of the source through different phases of the flaring activity around the VHE emission detected by MAGIC.

The ejection of a new radio knot and the rotation of optical EVPA were detected close in time to VHE emission.



Acciari et al. 2021, A&A, 647, A163



SED at different epochs around the VHE detection can be explained in EC scenario on DT photons + SSC with electron energy distribution limited by an interplay of acceleration, dynamic and cooling time scales.