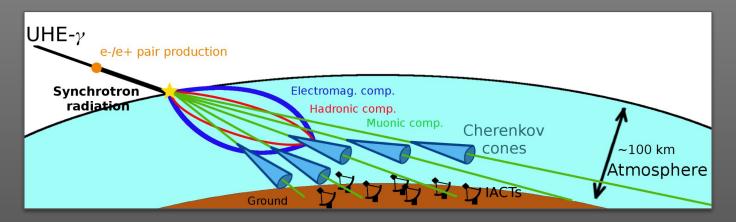
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- At ultra-high energy, photon/hadron separation retrieved in nearly-horizontal directions for IACTs.
  - → preshower effect must be taken into account.
- Cherenkov light emitted by the muonic component and recorded by IACTs.

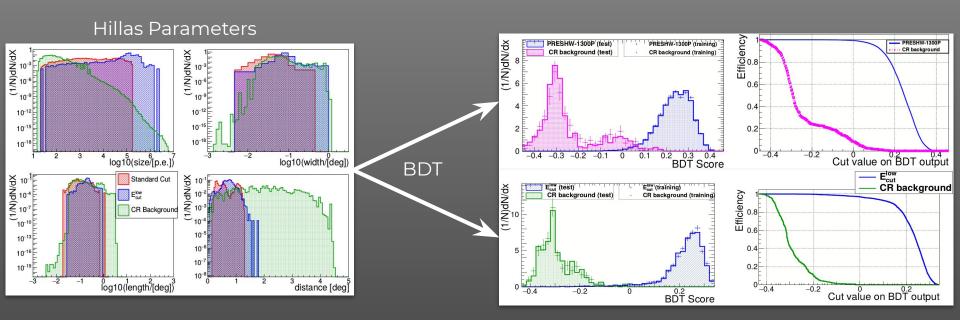






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**Hillas parameters** used to discriminate between preshower and cosmic-ray background events.



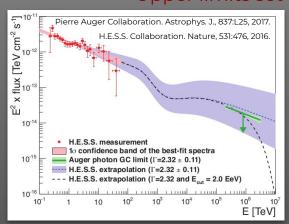


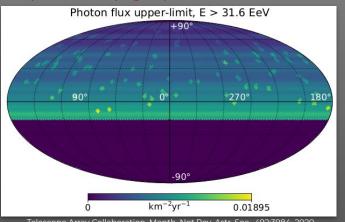


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#### Upper limits set by Auger (left) and TA (right)





Telescope Array Collaboration. Month. Not.Roy. Astr. Soc., 492:3984, 2020.

	AUGER <sub>point</sub>	$\langle \mathrm{TA}_{E>31.6\mathrm{EeV}} \rangle$	$\max(\mathrm{TA}_{E>31.6\mathrm{EeV}})$
$\phi_{\gamma_{-p.}}$ (40 EeV) [km <sup>-2</sup> yr <sup>-1</sup> ]	0.034	0.0073	0.019
$N_{\text{preshw}}$ – non-transient ( $R = 1$ )	$2.7 \times 10^{-4}$	$5.7 \times 10^{-5}$	$1.5 \times 10^{-4}$
-R = 5	$1.4 \times 10^{-3}$	$2.9 \times 10^{-4}$	$7.6 \times 10^{-4}$
-R = 652	0.17	0.037	0.09



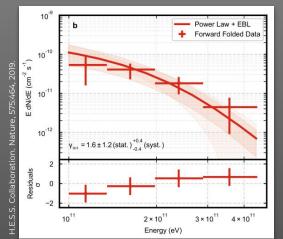


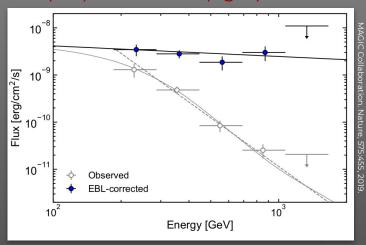


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### GRB observations by H.E.S.S. (left) and MAGIC (right)





	AUGER <sub>point</sub>	$\langle \mathrm{TA}_{E>31.6\mathrm{EeV}} \rangle$	$\max(\mathrm{TA}_{E>31.6\mathrm{EeV}})$
$\phi_{\gamma_{-p.}}$ (40 EeV) [km <sup>-2</sup> yr <sup>-1</sup> ]	0.034	0.0073	0.019
$N_{\text{preshw}}$ – non-transient ( $R = 1$ )	$2.7 \times 10^{-4}$	$5.7 \times 10^{-5}$	$1.5 \times 10^{-4}$
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