#### The CALET experiment

The CALorimetric Electron Telescope (CALET), operating on board the International Space Station since October 2015, is an experiment dedicated to high-energy astroparticle physics. The primary scientific goal of the experiment is the measurement of the electron+positron flux up to the multi-TeV region. This is possible thanks to the excellent detector performances, having a geometrical factor of 0.1 m<sup>2</sup>sr, an electromagnetic shower energy resolution of 2% and a proton rejection factor of 10<sup>5</sup>. These performances are obtained by an accurate design of the CALET calorimeter, made by a CHarge Detector (CHD), an IMaging Calorimeter (IMC) and a Total AbSorption Calorimeter (TASC).



## (a) Selections

Group of selections in the order they are applied: (1) Offline trigger confirmation - select a flat region in the efficiency curve of the trigger discriminators

(2) Geometrical condition - select tracks inside acceptance



(3) Track quality - ensure an accurate track reconstruction (4) Charge selection - remove contamination from He and nuclei (5) Longitudinal shower likelihood - suppress contamination from proton (6) Lateral shower concentration - suppress contamination from proton and events outside acceptance

The electron analysis strategy is divided in two main steps: (a) a group of selections to obtain a well reconstructed sample of electron candidates, removing contamination from events outside acceptance and particles with charge Z>1 ◆Above 30 GeV, selection efficiency is higher than 95% for electrons and smaller than 1% for protons (b) a proton rejection cut to further suppress the contaminating proton background (necessary since, in cosmic rays, protons are more abundant than electrons by a factor 100-1000) • The residual proton contamination (<5 % for E < 1 TeV and <20% for E > 1 TeV [<10% for the optimized BDT analysis using 13 parameters]) is subtracted from the final measurement



The analysis strategy for the measurement of the electron flux with CALET on the **International Space Station** Eugenio Berti<sup>1,2</sup>, Lorenzo Pacini<sup>2</sup> and Yosui Akaike<sup>3</sup> for the CALET collaboration (1) Florence University, Italy (2) INFN -Florence, Italy (3) Waseda University, Japan

#### Analysis strategy

#### **Systematic Uncertainties**

Systematic uncertainties can be divided in two groups: • Normalization uncertainties, *i.e.* detector acceptance, longterm stability, radiation environment, and live time, for a total of 3.2% • Energy-dependent uncertainties, *i.e.* trigger efficiency, track



## (b) Proton rejection - single cut

Exploiting the different longitudinal and lateral development of the electromagnetic and hadronic showers, protons can be rejected using a single cut based on K =  $\log_{10}(F_E) + 0.5 \times R_E$  [cm]:

### (b) Proton rejection - MVA cut

To further suppress the proton background, a multivariate algorithm based on Boosted Decision Tree (BDT) is used. BDT estimator is built employing 9 variables:  $F_E$ ,  $R_E$ , variables from longitudinal fit in IMC ( $p_0$ ,  $p_1$  and goodness of fit) and in TASC (shower maximum  $\alpha/b$ , attenuation constant b, 5% shower depth, goodness of fit)

**Longitudinal Fit in IMC**  $\frac{dE}{dt} = e^{p_0 + p_1 t}$ 

**Longitudinal Fit in TASC**  $b^{(\alpha+1)}$ dE $F = E_0 \frac{1}{\Gamma(\alpha + 1)}$ \_\_\_\_\_ dt

• F<sub>E</sub> is the fraction of energy deposited in the last TASC layer respect to the total energy deposited in TASC

• R<sub>E</sub> is the second moment of the lateral energy-deposit distribution in the TASC first layer computed with respect to the shower axis

# $-t^{\alpha}e^{-bt}$

All these variables were chosen considering their e/p discrimination power and their level of data/MC agreement. After a careful optimization of the algorithm we opted for a BDT made of 100 trees with a depth of 20. This detailed study leads to very stable performances at all energies, except above 1 TeV due to

the limited statistics. In an optimized BDT analysis that is currently under study at high energies, 4 additional variables are used to build