

Electron-Neutron Detector Array (ENDAs), Status and Coincidence with the LHAASO Detectors

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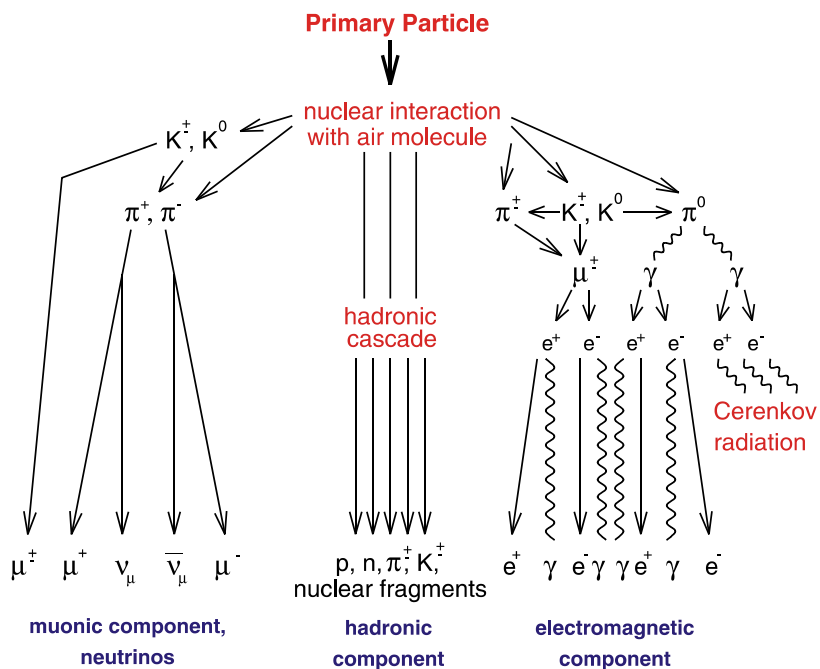
Hebei Normal University

outline

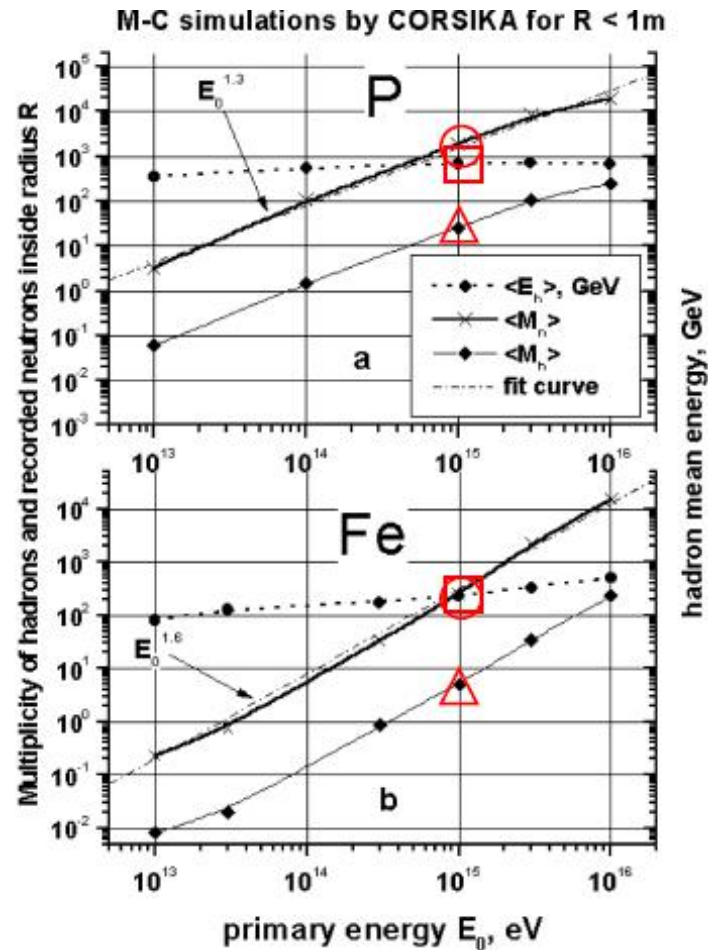


- 1. Physical motivation**
- 2. EN-detector**
- 3. Early study at high altitude**
- 4. ENDA in LHAASO**
- 5. Summary**

1. Physical motivation

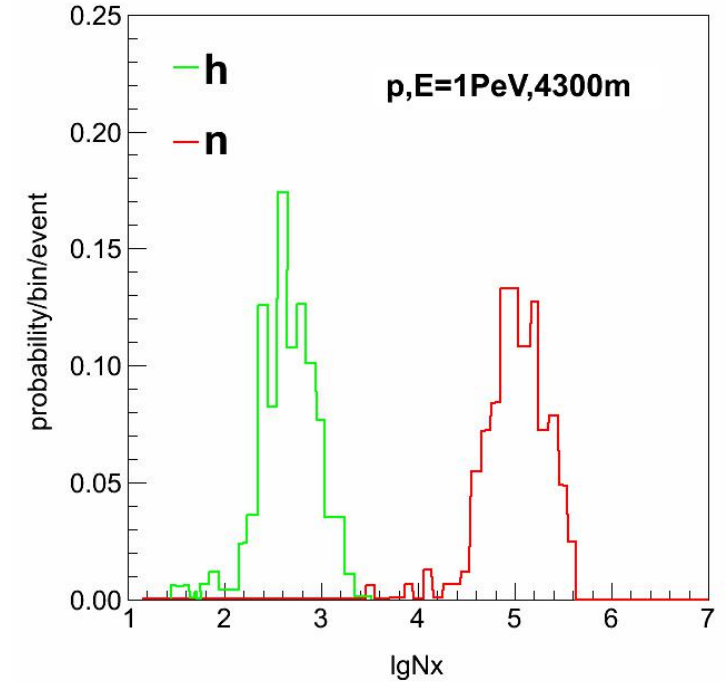
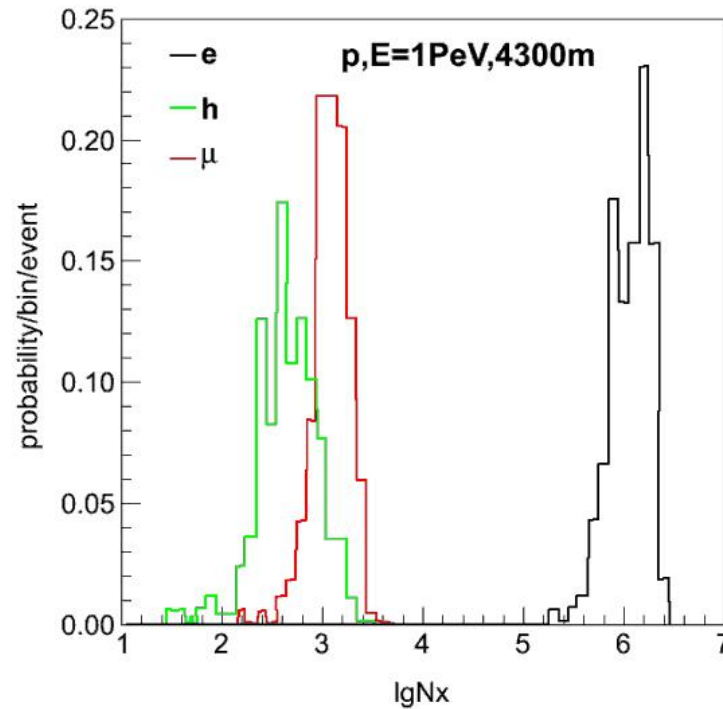
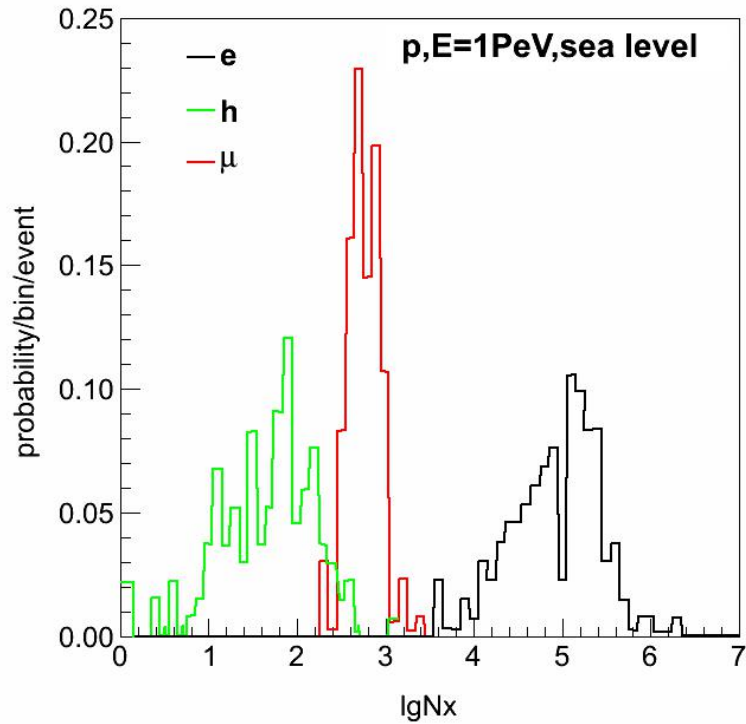


- Hadrons are the backbone of the shower development, very sensitive to primary composition.
- Hadrons can generate amount of fast neutrons in ground media (soil, building, etc.). Fast neutrons are moderated to thermal neutrons.



At the same primary energy, thermal neutrons generated by light components (such as proton) are one order more than one by heavy components (such as iron). It is very good for primary component separation.

**Modern Physics Letters A, Vol.
17, No. 26 (2002) 1745-1751**

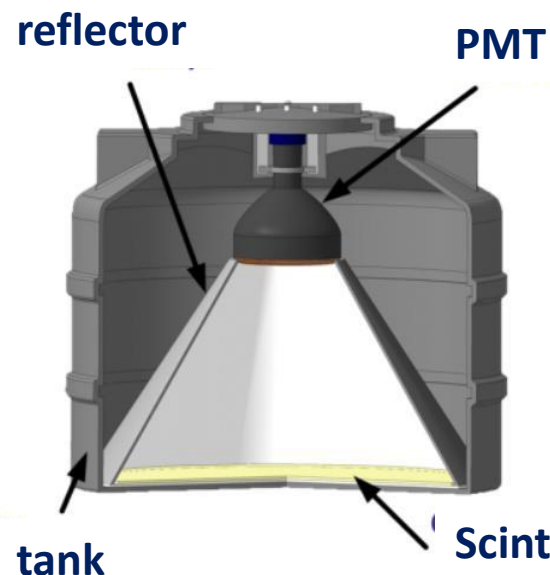


- Thermal neutrons are 2-3 orders more than hadrons.
- thermal neutrons are 1-2 orders at high altitude more than one at sea level

2. EN-detector

EN-detector (electron-neutron detector), developed by Yuri Stenkin et al., can detect both thermal neutrons and “charged” components.

Nuclear Physics B (Proc. Suppl.) 196 (2009) 293–296

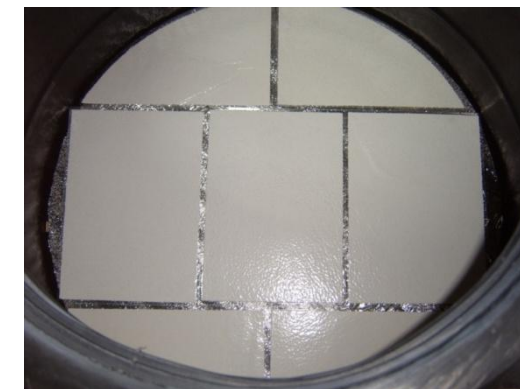


Thermal neutron recording efficiency ~20%.
Scintillator effective thickness 30 mg/cm².



PRISMA(PRImary Spectrum Measurement Array)

Nucl. Phys. B (Proc. Suppl.), 196, (2009), p. 293-296.

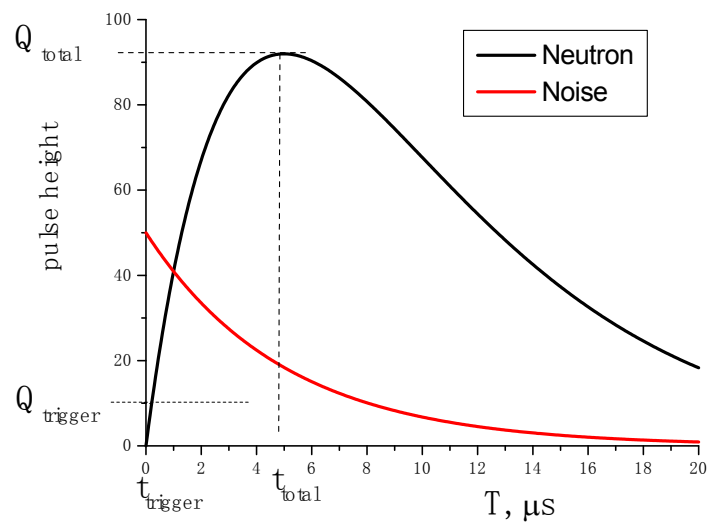


$\text{ZnS(Ag)}+{}^6\text{LiF}$

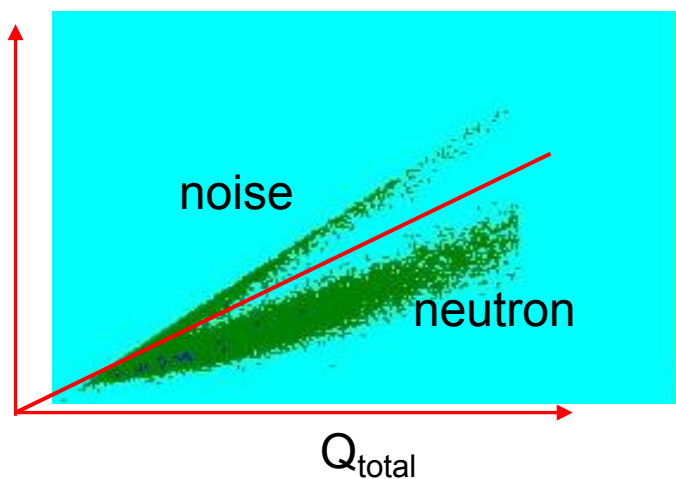


$\text{ZnS(Ag)}+{}^{10}\text{B}_2\text{O}_3$

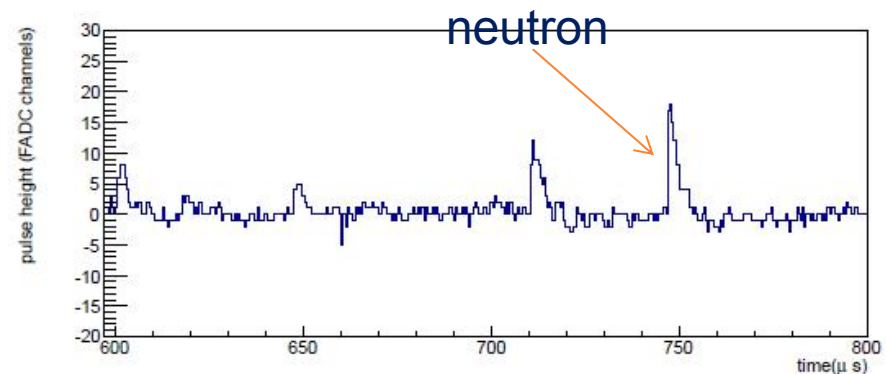
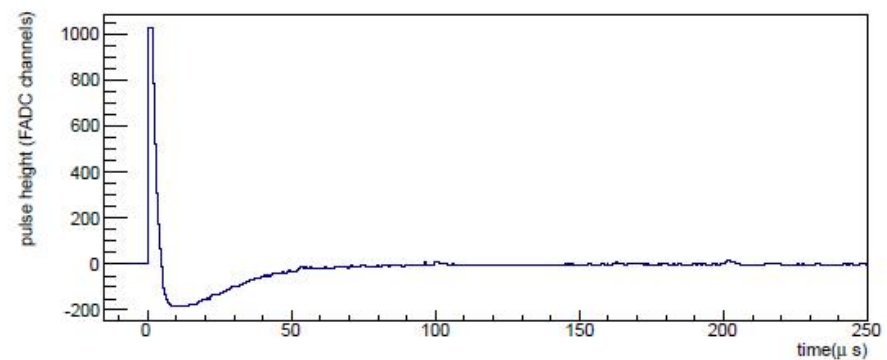
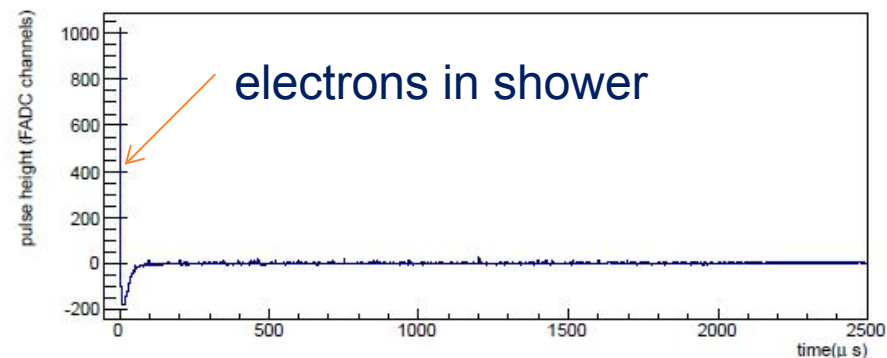
neutron / noise separation



$Q_{trigger}$



electrons / neutron separation in one shower



3. Early study at high altitude

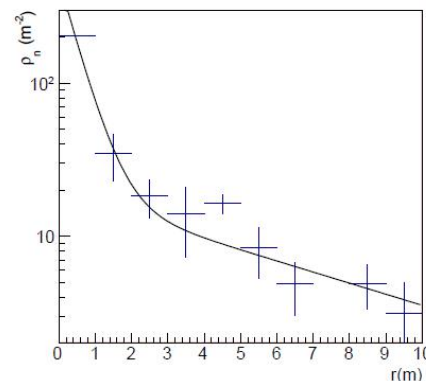
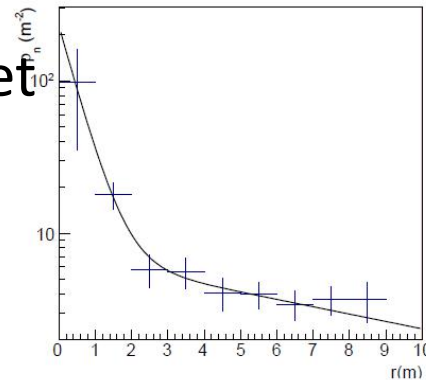
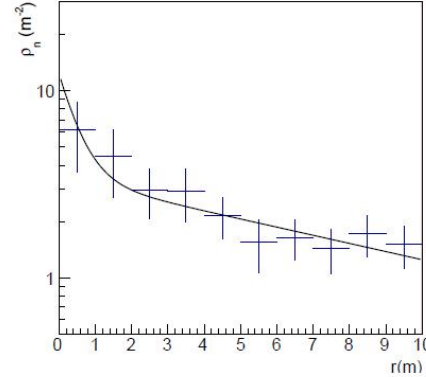
PRISMA-YBJ:

4 EN-detectors

4300m a.s.l.

Yangbajing Tibet, China,
from Jan. 2013.

In March 2016 move to Tibet
University in Lhasa



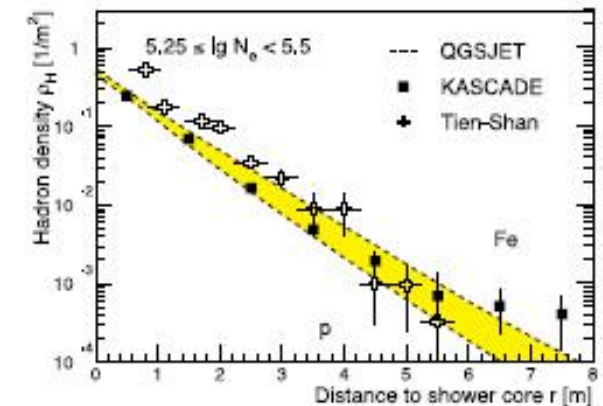
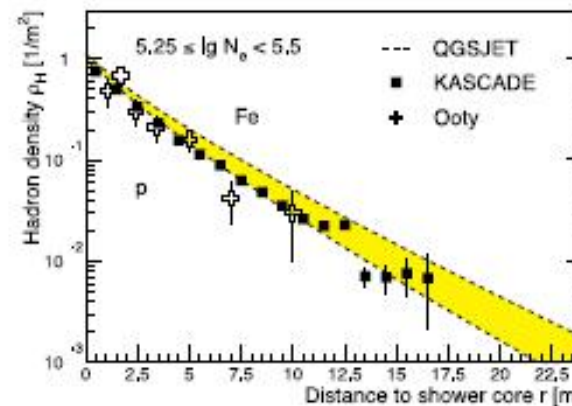
thermal neutron lateral distribution

Astroparticle Physics 81 (2016) 49–60

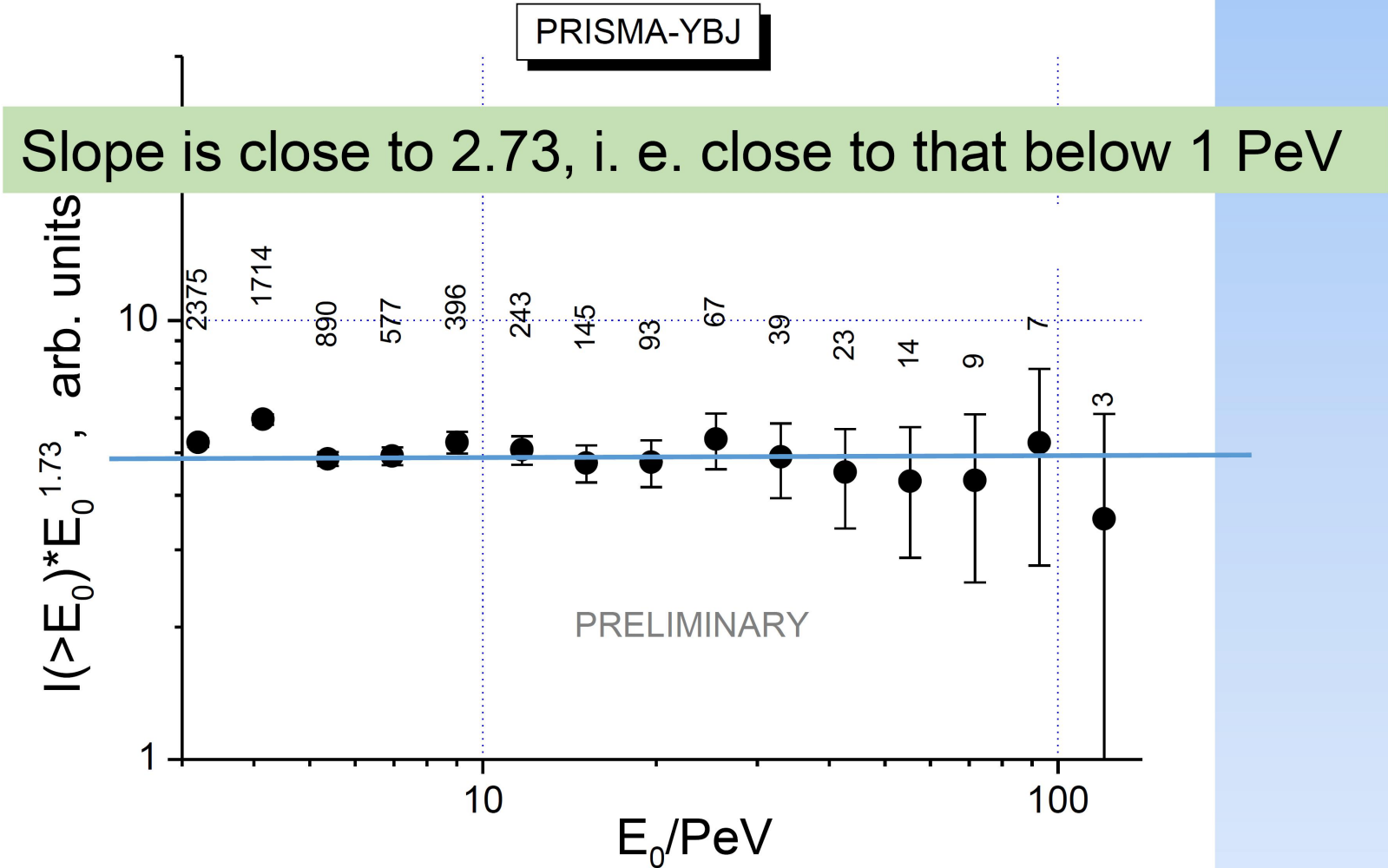
$$\rho_n(r) = \rho_0 \times e^{-(r/r_0)} + \rho_1 \times e^{-(r/r_1)}$$

N_{p10} intervals	χ^2/ndf	$\rho_0(m^{-2})$	$\rho_1(m^{-2})$
$\lg(N_{p10}) < 4.8$	2.44/8	9.0 ± 6.8	3.41 ± 0.32
$4.8 < \lg(N_{p10}) < 5.4$	2.69/7	222 ± 65	7.17 ± 0.65
$\lg(N_{p10}) > 5.4$	20.1/7	456 ± 230	18.7 ± 2.3

hadron lateral distribution, KASCADE HCAL, sea level

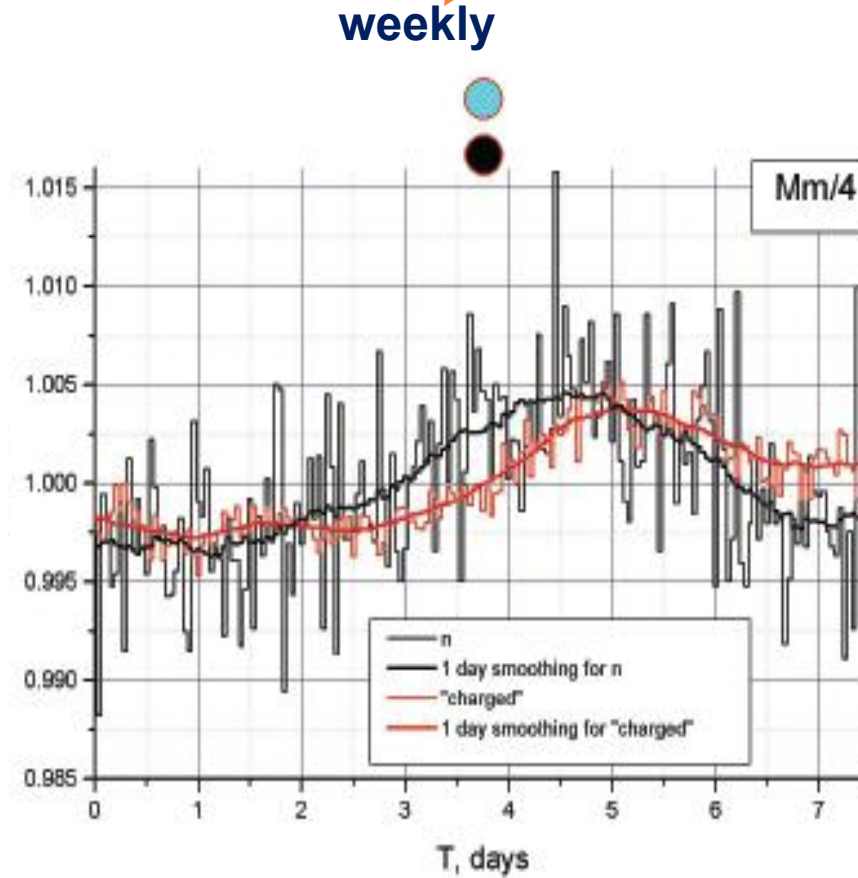
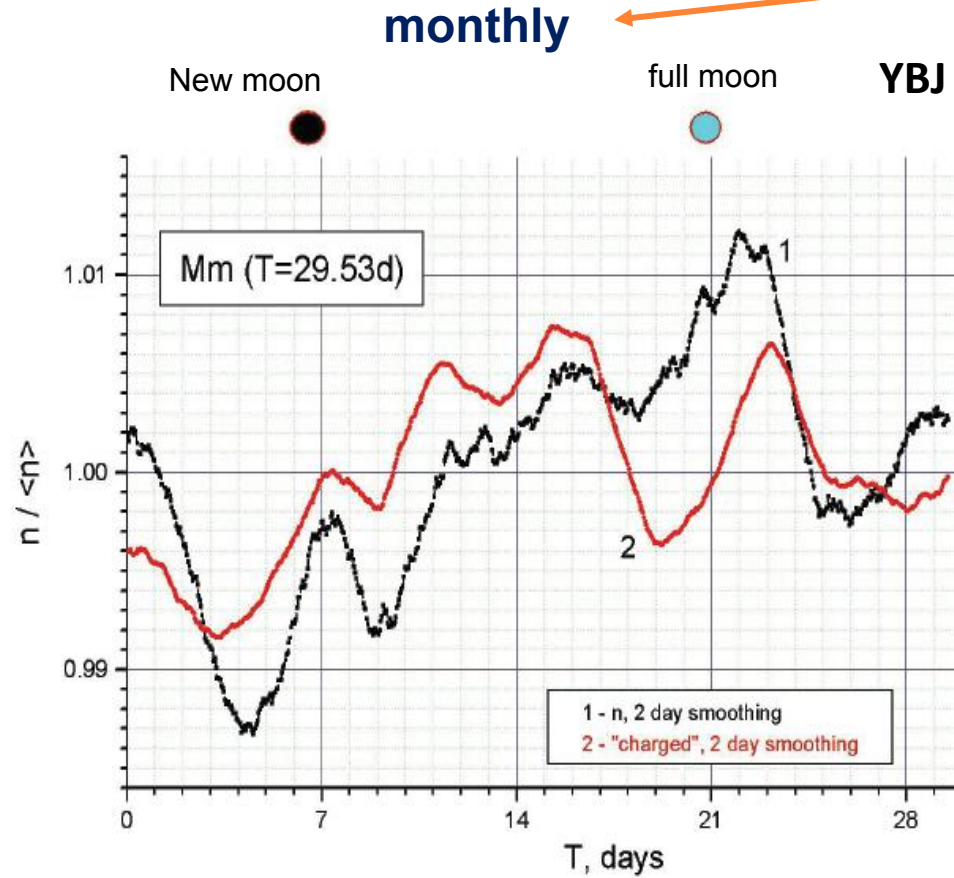
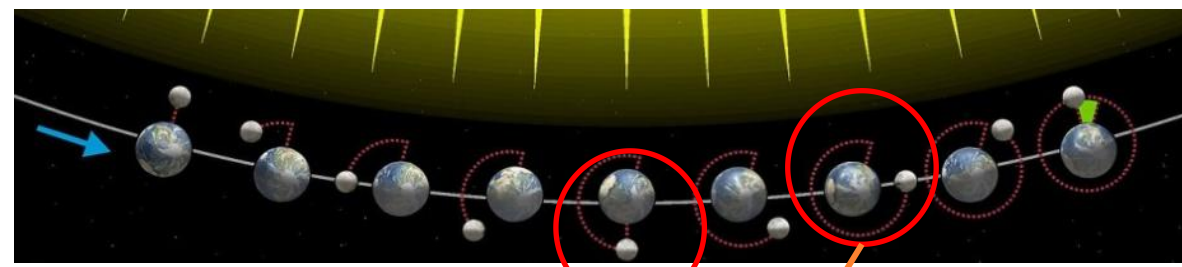


Result of PRISMA-YBJ from Nn measurement



Our preliminary result indicates that no significant slope changing above 3PeV

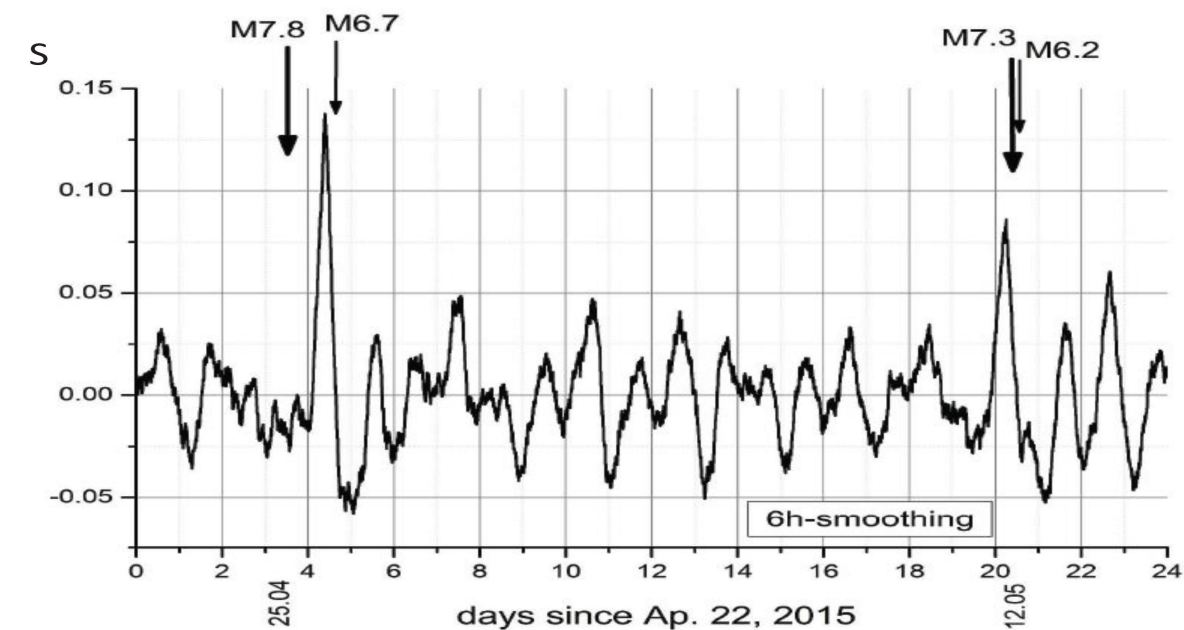
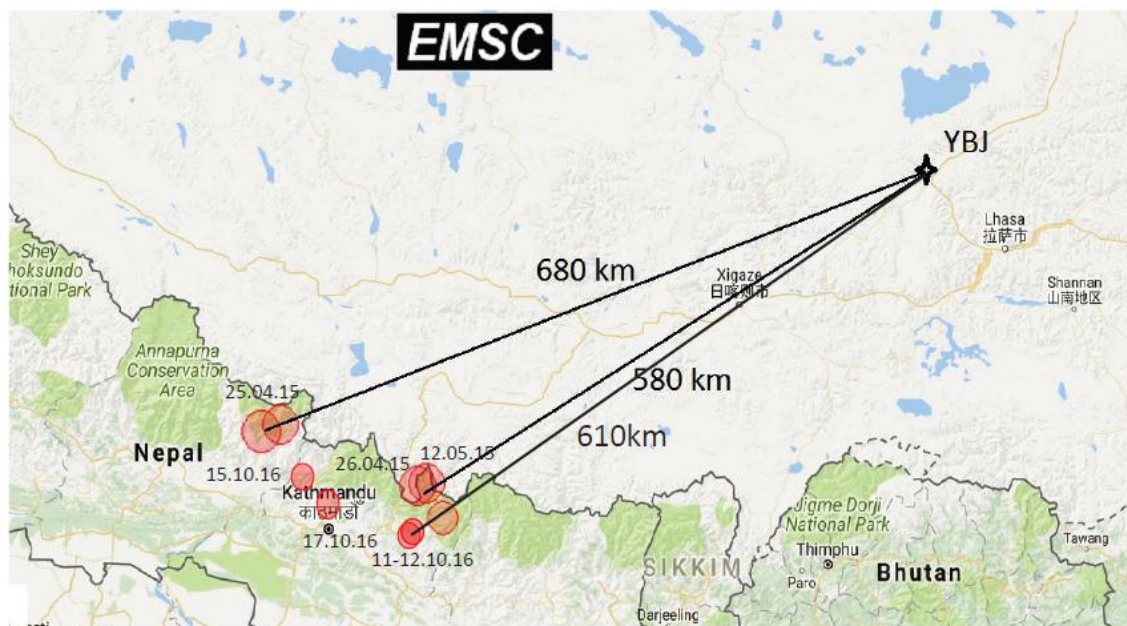
lunar tidal effect



superimposed epoch analysis

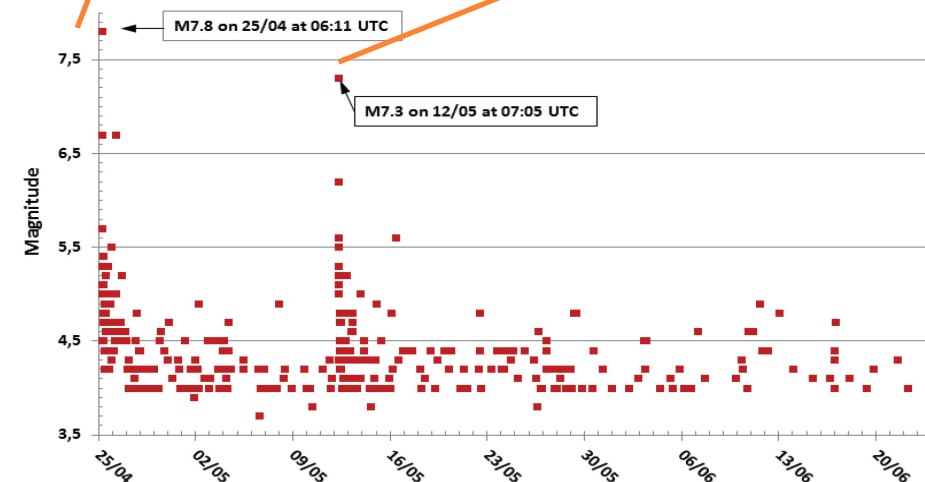
Pure And Appl. Geophys. 174 (2017) 2763–2771

Response of PRISMA-YBJ to 2015 Nepal earthquakes



M7.8 in Nepal on 25/04/2015
Earthquakes time sequence

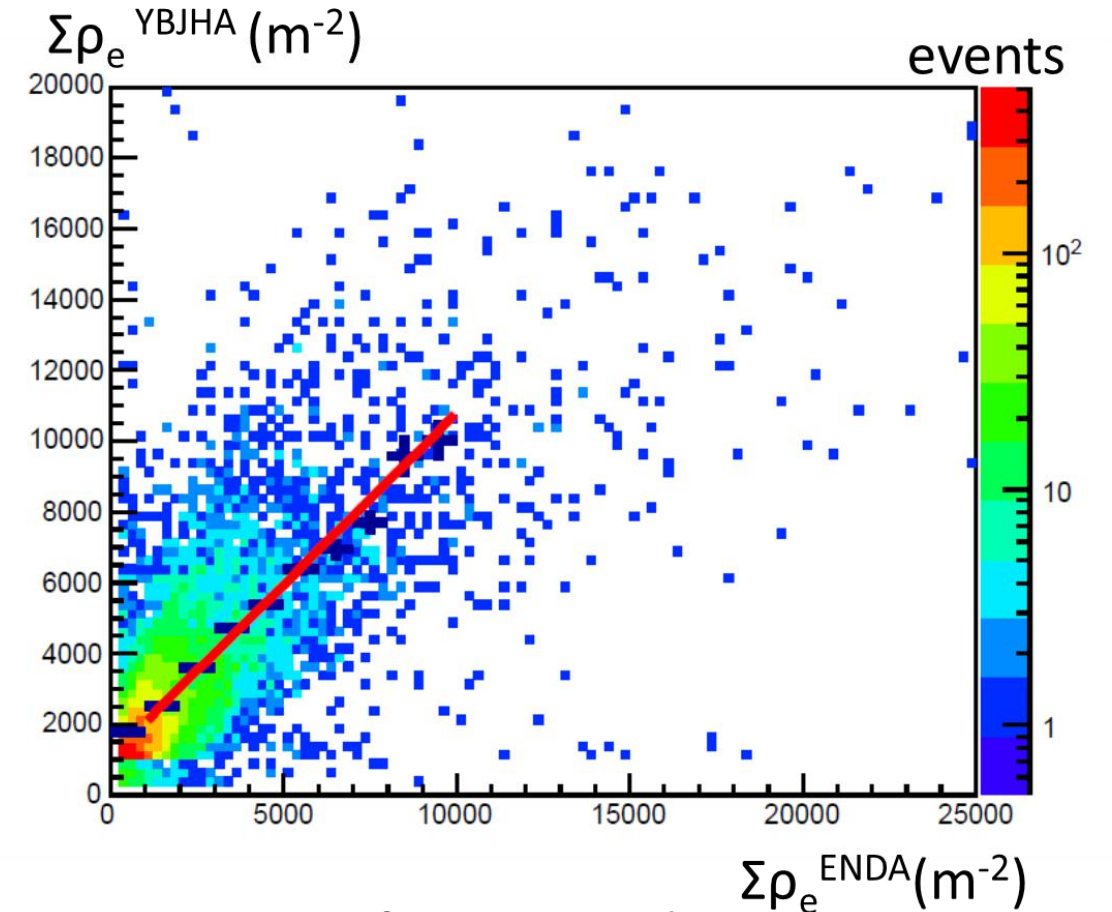
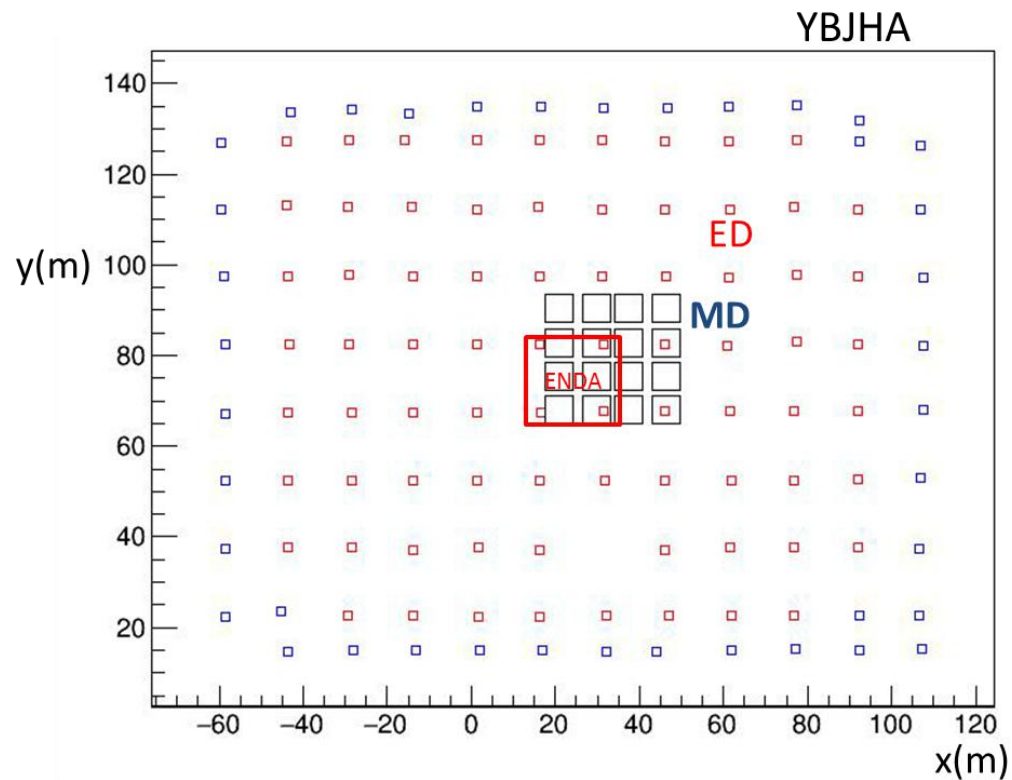
earthquakes magnitude



**Journal of Environmental Radioactivity 208-209 (2019)
105981**

PRISMA-16 at Tibet University and Yangbajing

2017 JINST 12 P12028

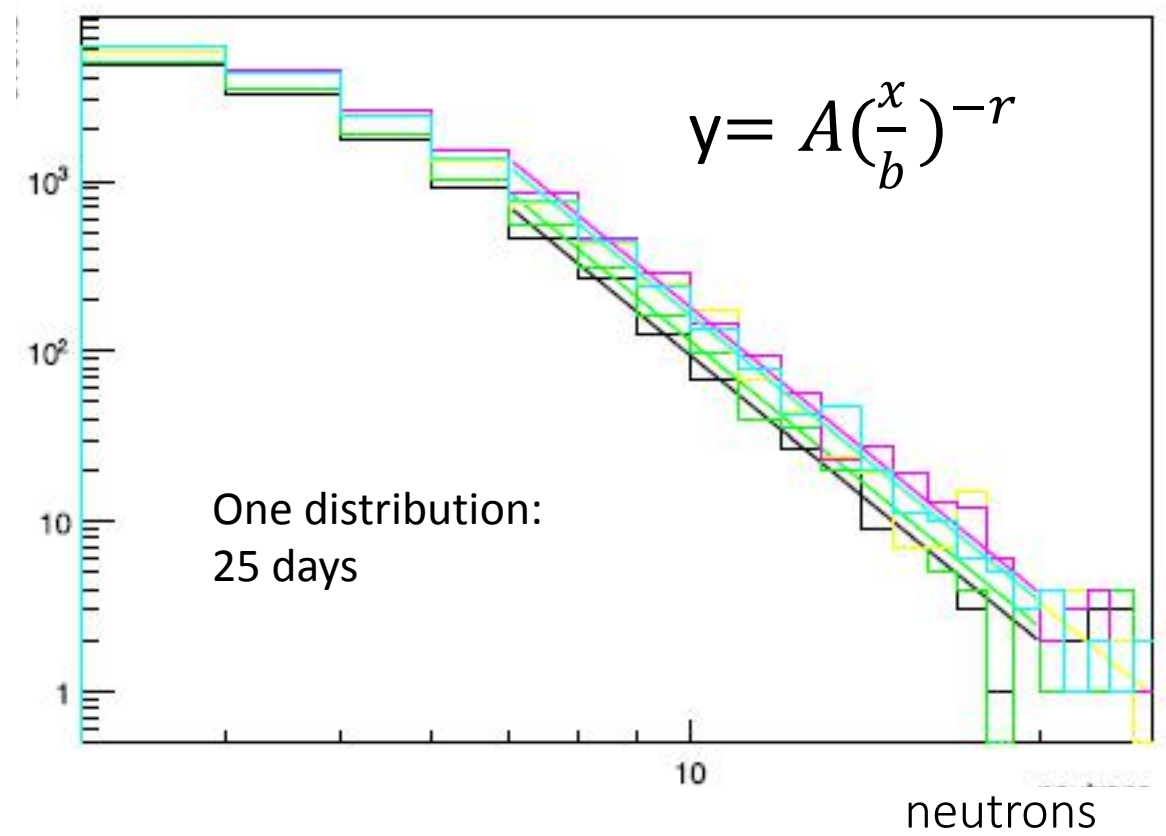


Linear fitting $y=a+bx$

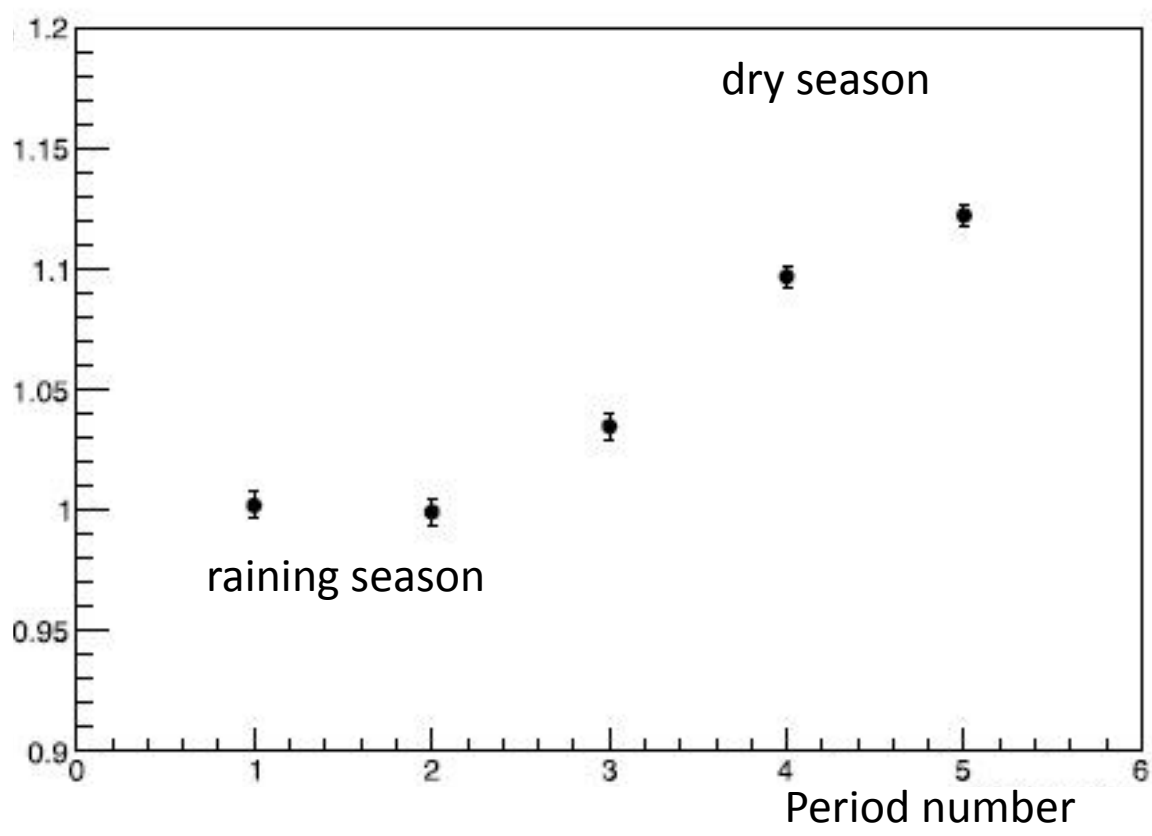
$$b=0.97 \pm 0.02$$

$$a=1099 \pm 43$$

events



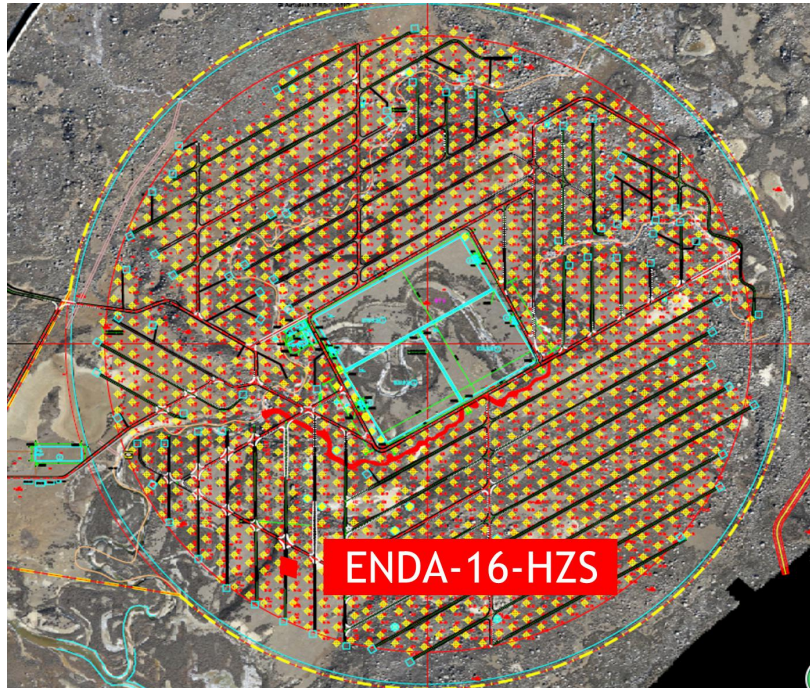
b: neutrons relative change



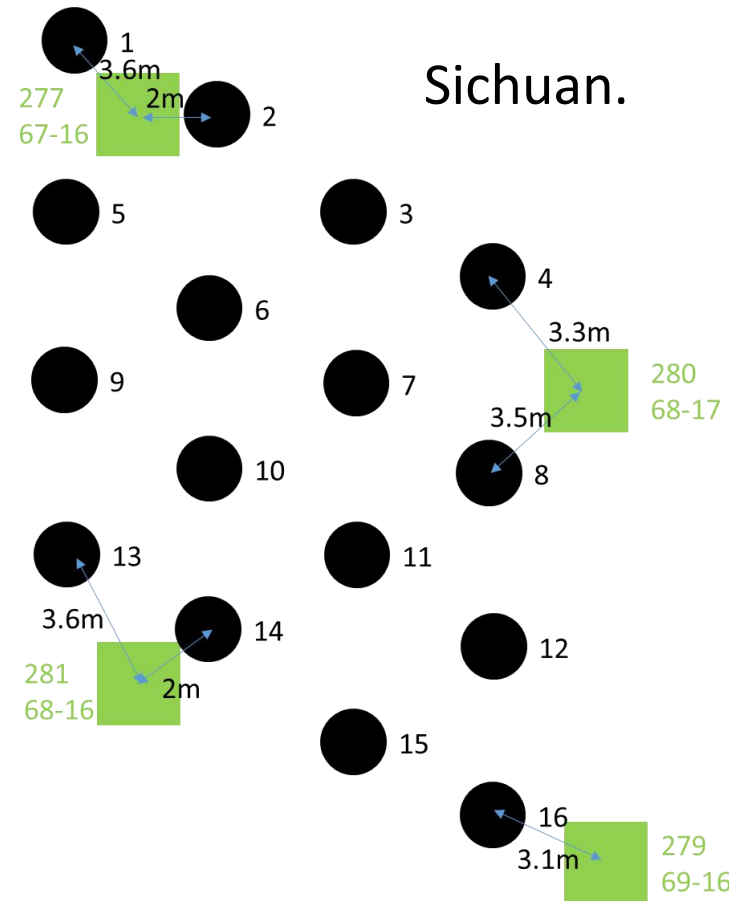
Neutrons in rain season are 10% less than ones in dry season because more water in soil.

Astrophysics Space Science (2020) 365:123

4. ENDA (EN-Detector Array) in LHAASO

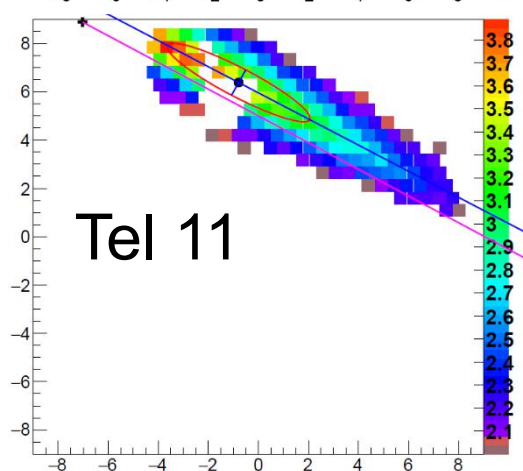
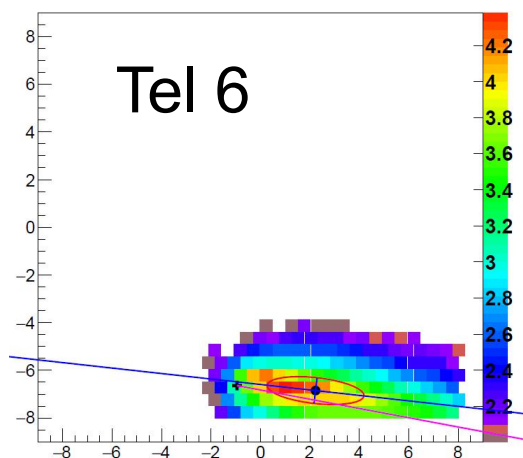


In 2019, a cluster so called ENDA-16-HZS was installed at LHAASO in Haizishan (HZS), Sichuan.

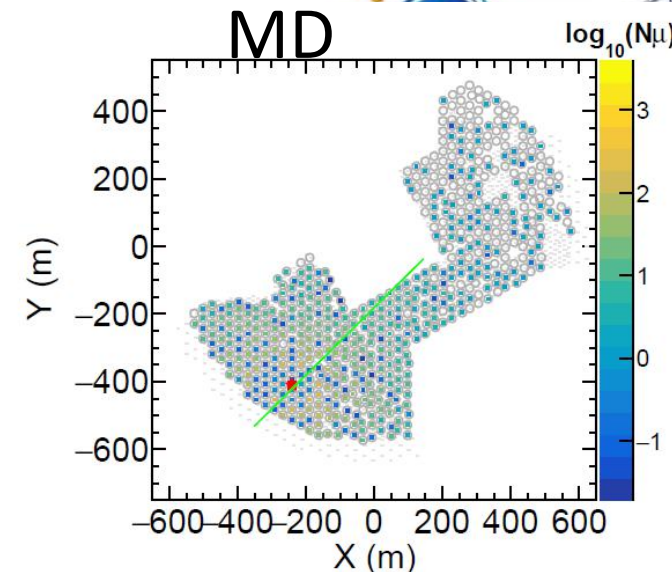
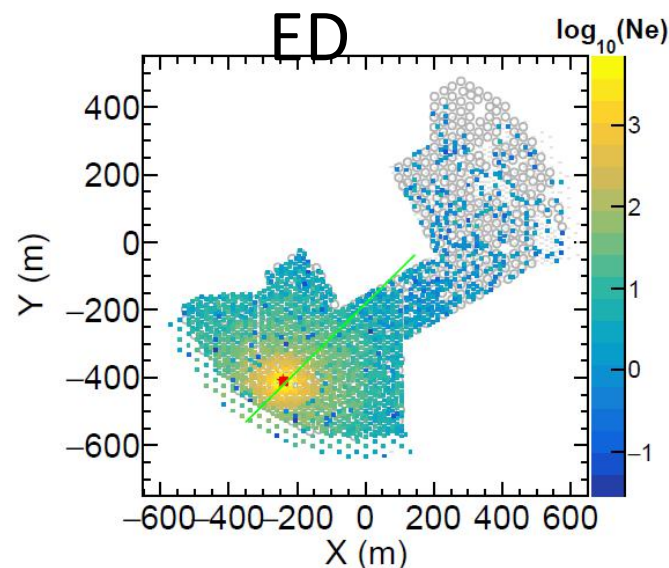
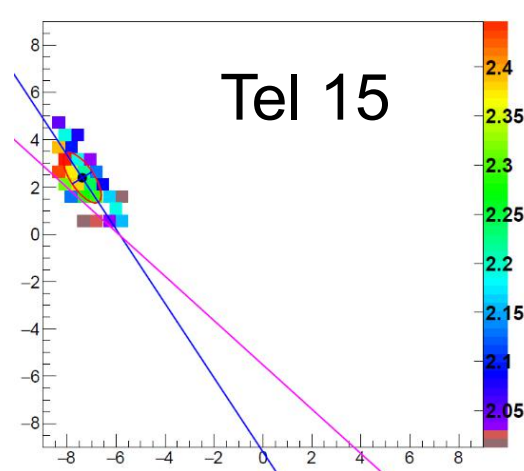


One Coincident event 1

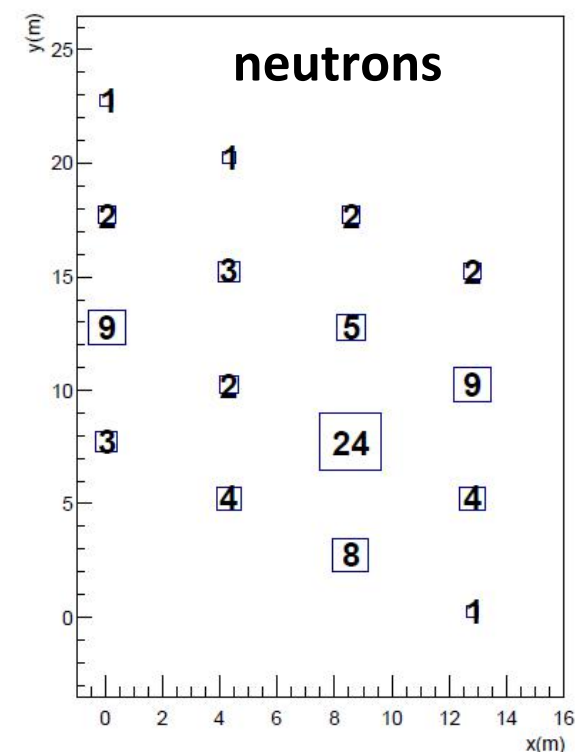
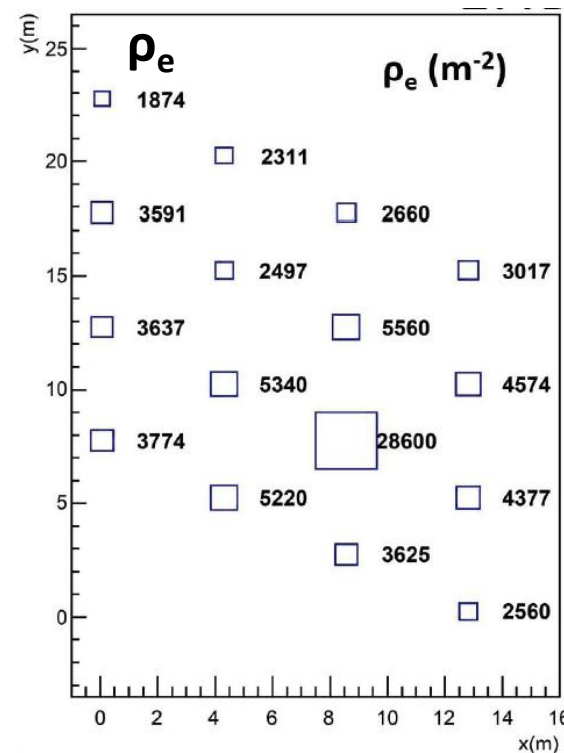
20210205_event4531

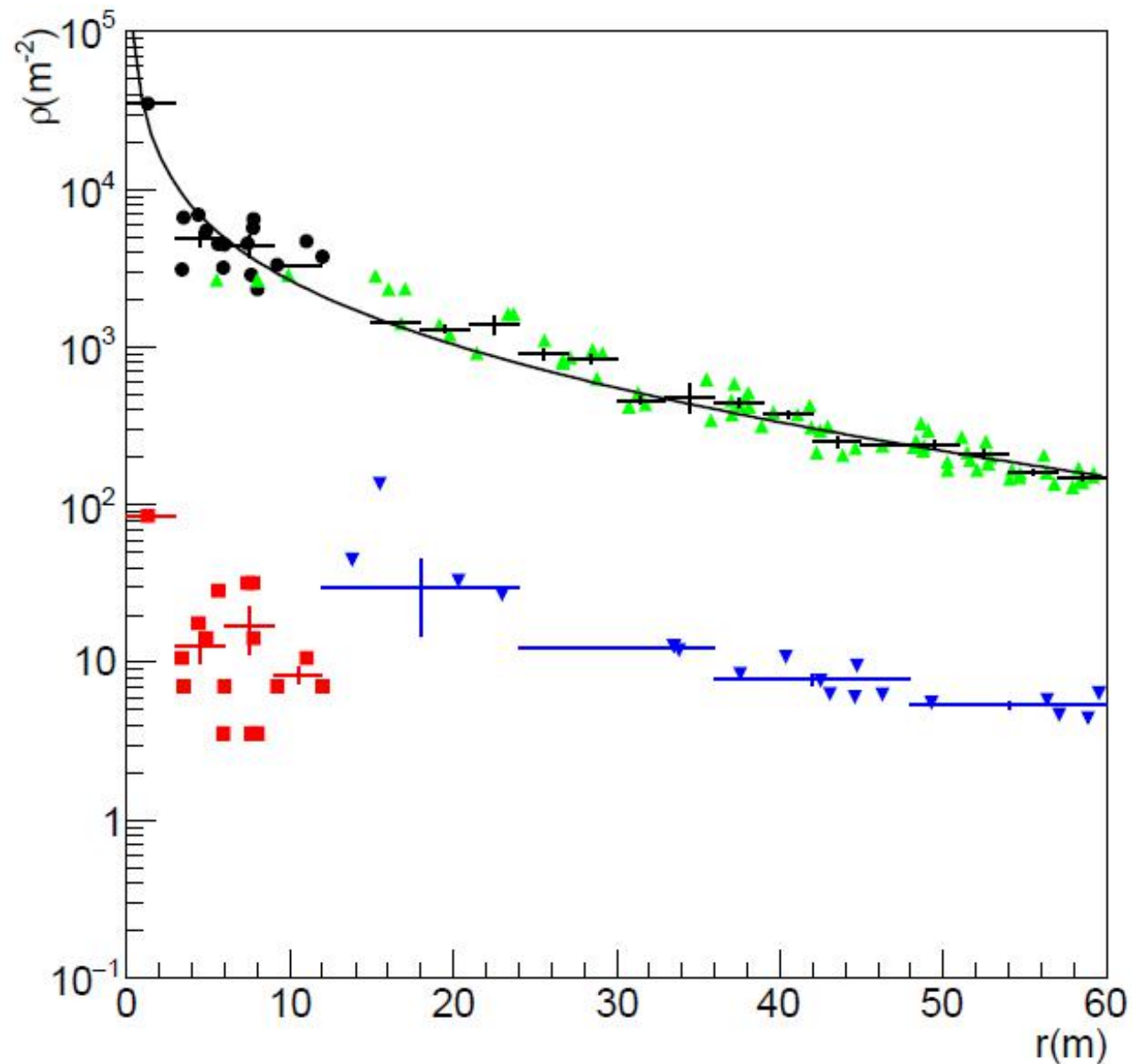


WFCTA



ENDA-16-HZS





Lateral distribution of the coincident event.

Red square: neutrons by ENDA.

Blue triangle: muons By KM2A-MD.

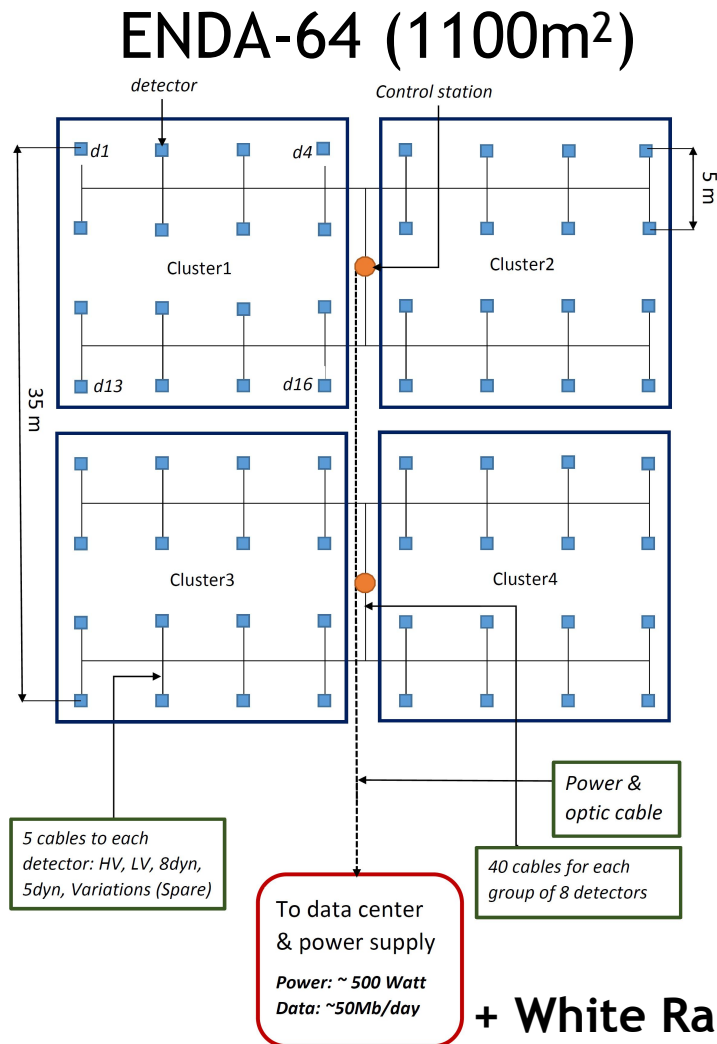
Black dot: electrons by ENDA.

Green triangle: electrons by KM2A-ED.

The black line is NKG function fitting.

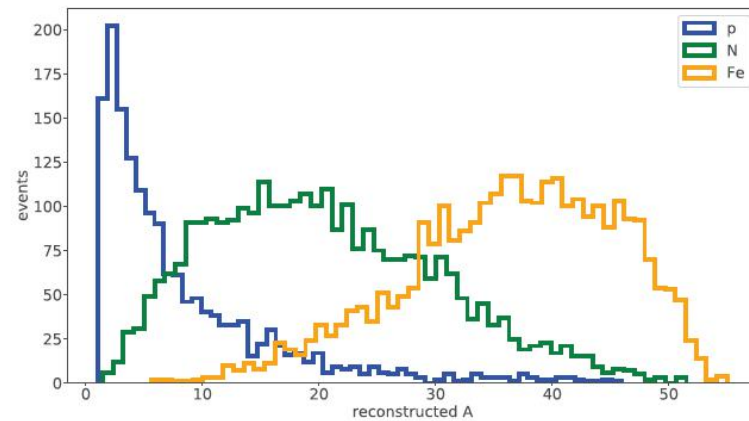
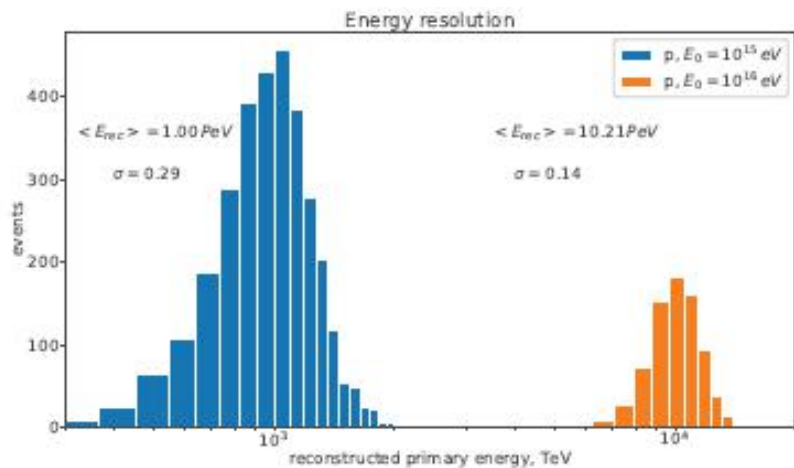
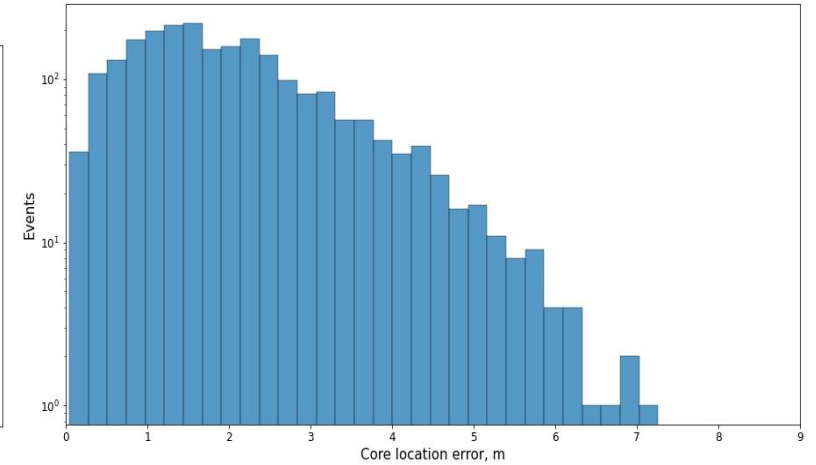
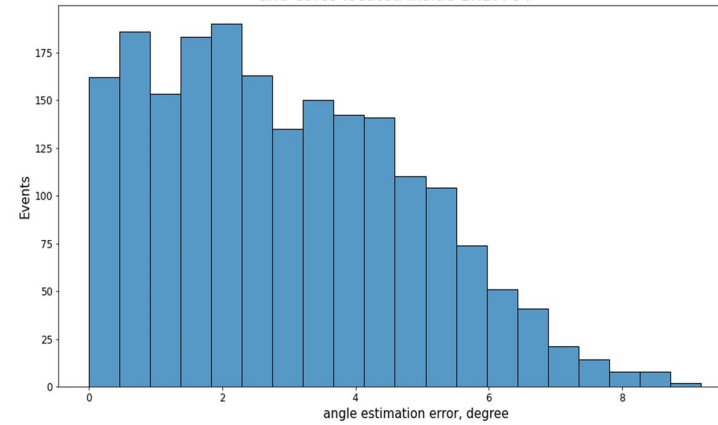
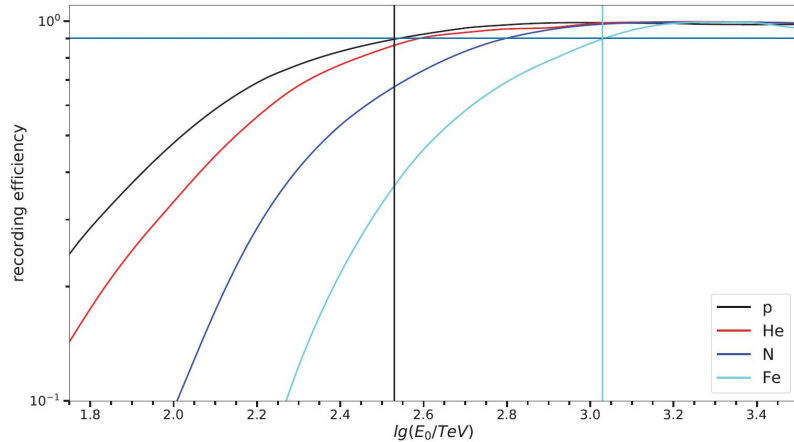
64 EN-detectors are made and will be added into LHAASO.

**Bulletin of the Russian Academy of Sciences: Physics, 2021,
Vol. 85, No. 4, pp. 405–407**



**EN-detectors at Hebei Normal University
FADCs are made by Sichuan University**

ENDA-64 simulation



At 1PeV

Efficiency: 90%

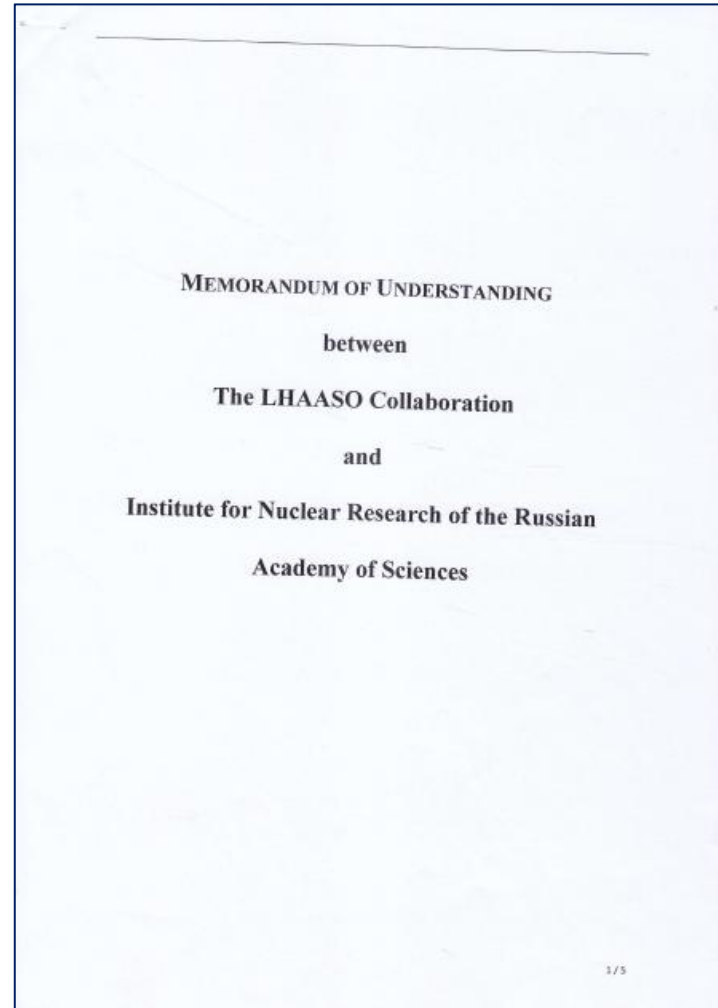
Angular resolution: 4°

Core position resolution: 3m

Energy resolution: 30%

ENDA-400: written into MoU

ENDA-400 (10000m²)

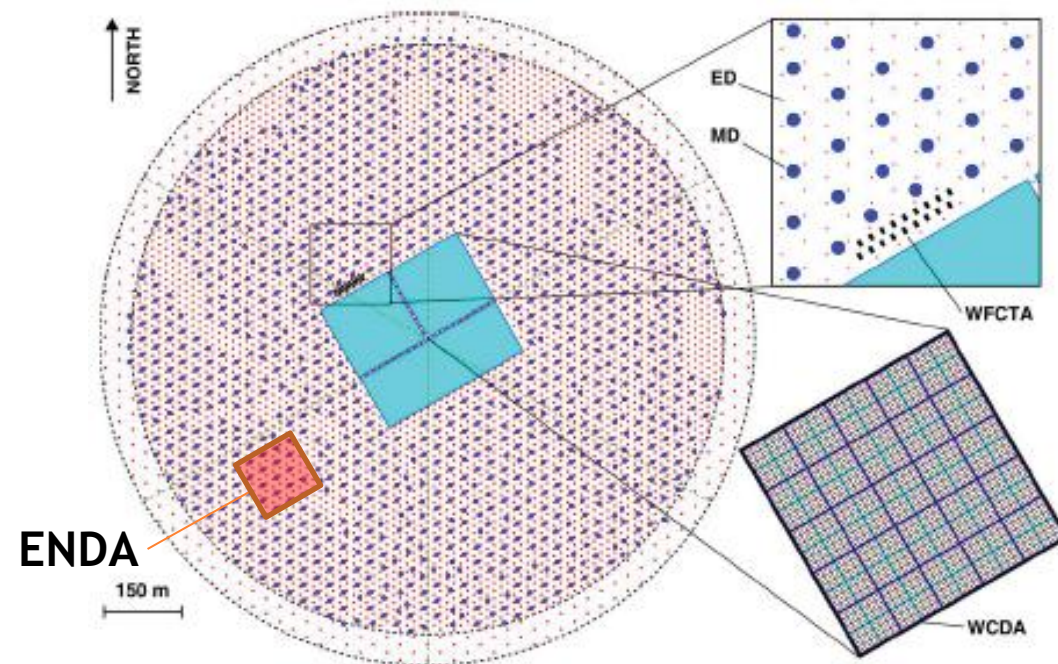


LHAASO at the knee region

- ED : e
- MD, WCDA: μ
- WFCTA: \check{C}
- WCDA++: γ family at core $\rightarrow h^0$
- ENDA: thermal neutrons $\rightarrow h^+h^-$

\rightarrow Full particle measurement of cosmic showers!

\rightarrow significant capability of component separation and energy determination!



Thanks!