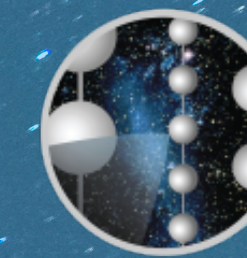


Direction Reconstruction

using a CNN for GeV-scale Neutrinos in IceCube

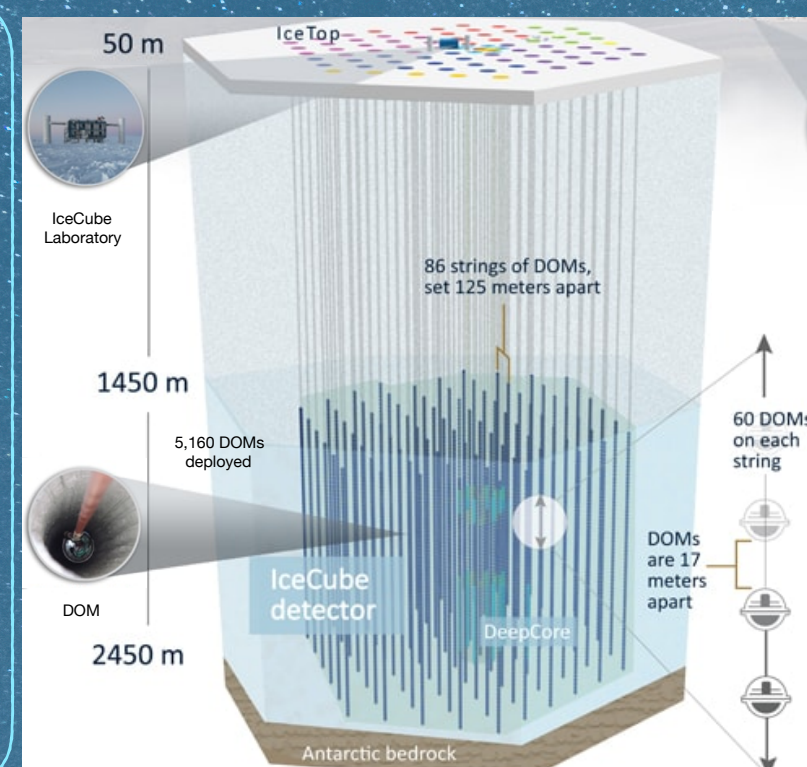
For the IceCube Collaboration



Shiqi Yu, Michigan State University

Introduction

- **IceCube Neutrino Observatory:** a Cherenkov detector observes neutrinos interacting deep within ice by detecting Cherenkov photons.
- **DeepCore (DC):**
 - subdetector at lower center of IceCube (IC)
 - lowers observable neutrino energy to GeV-scale
- **Precisely measuring neutrino direction (zenith) is critical in studying neutrino oscillations**
- **Reconstruct value of zenith using a CNN**

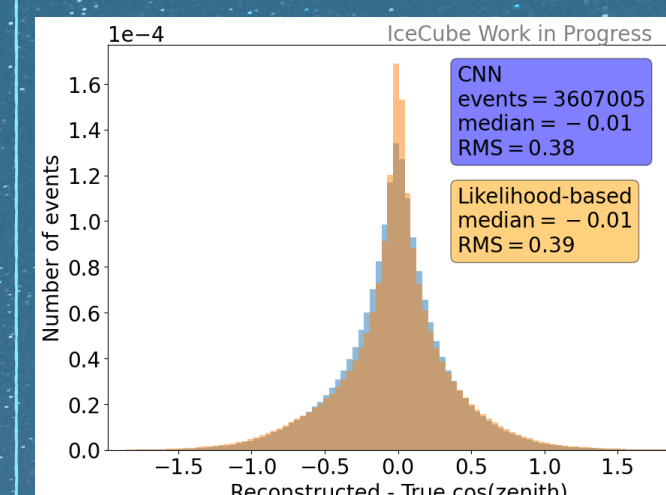


Siamese Tower

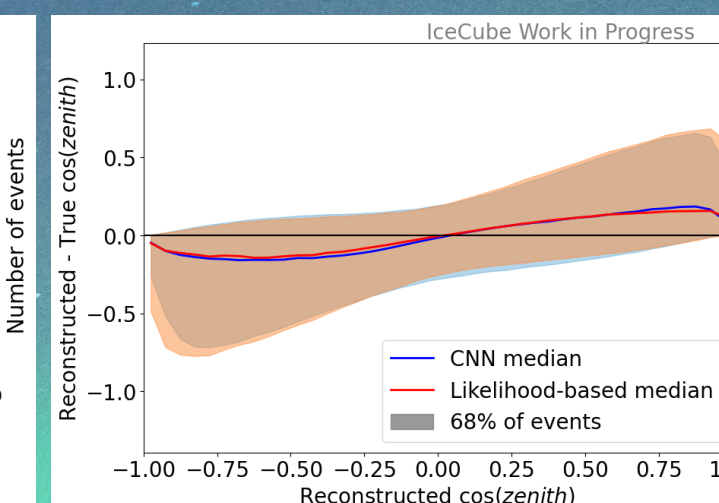
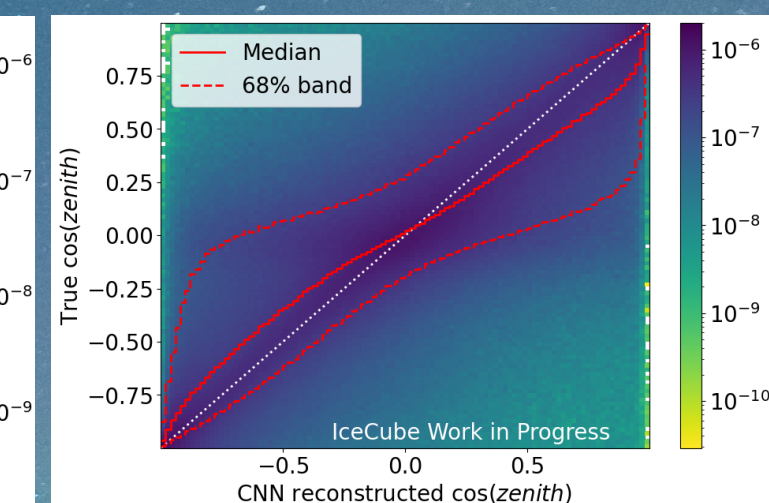
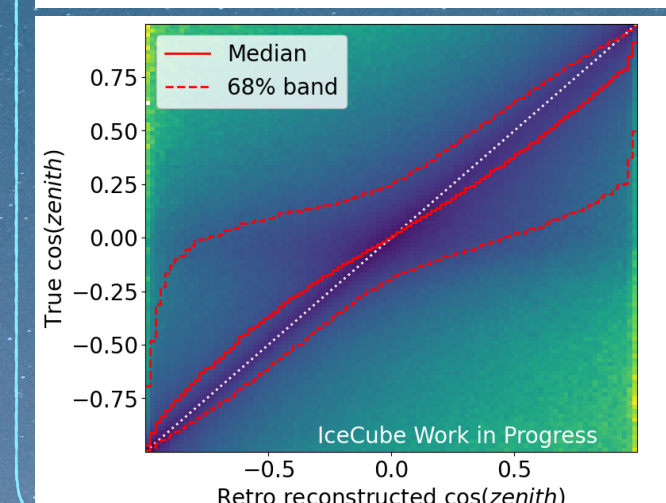
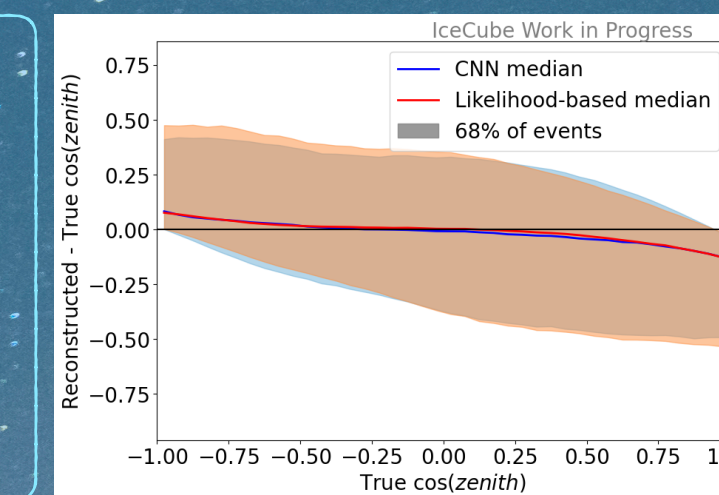


Performance

NuMu CC with current likelihood-based reconstruction ("Retro") cuts:
 $5 \text{ GeV} < E < 300 \text{ GeV}$; $-500 \text{ m} < z < -200 \text{ m}$; $\rho_{36} < 300 \text{ m}$



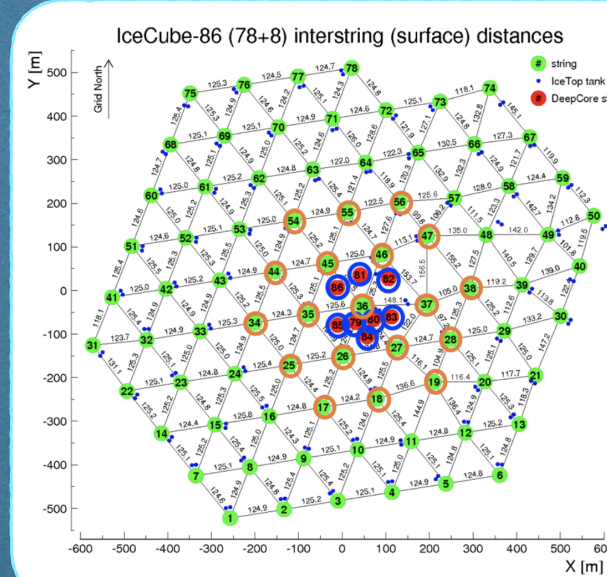
- **Smaller overall RMS**
- **Comparable bias against both true and reconstructed $\cos(\text{zenith})$**



Convolutional Neural Network (CNN)

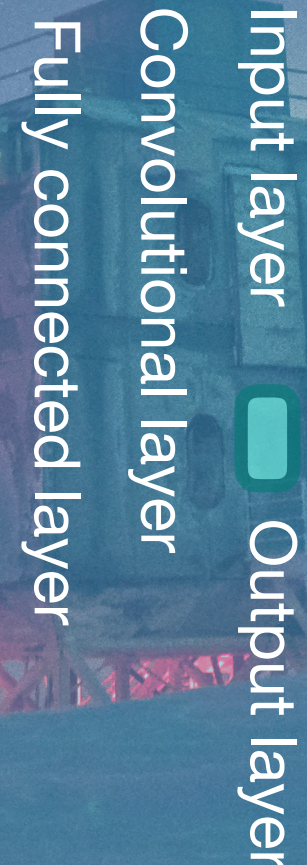
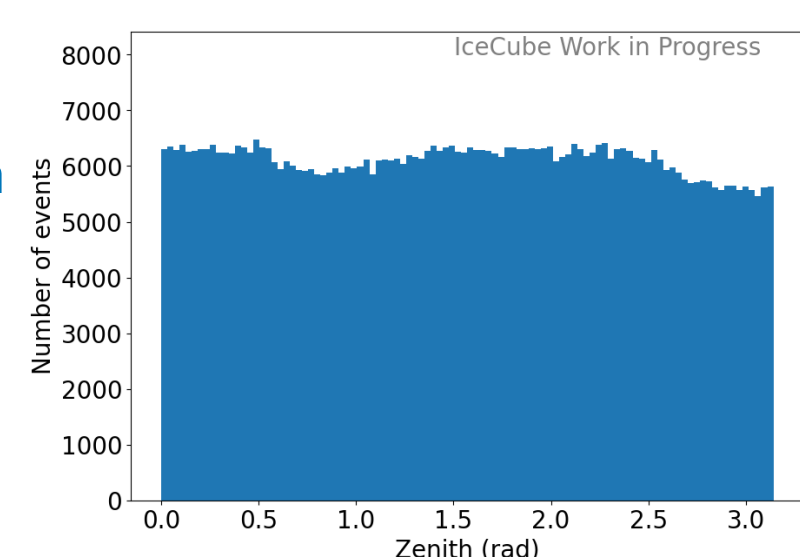
- **8 DC and 19 surrounding IC string as separate inputs:**
 - 60 digital optical modules (DOMs) per string
 - 5 variables calculated from the pulse series per DOM: sum of charges, time of first (last) pulse, charge weighted mean (standard deviation) of time of pulses

Training



- **5~300 GeV NuMu CC**
- **True vertex contained:**
 - vertex z in $(-505, 500) \text{ m}$
 - $\rho_{36} < 260 \text{ m}$
- **Flat zenith distribution**

Training set: 4 million;
 Validation set: 1 million



Summary & Future

Processing Speed

Time (s) /Event	GPU	CPU
CNN	0.0044	0.108
Retro	—	44.97

- **Improved RMS by 2.5%**
- **Up to 10,000 times faster than current processing**
- **Comparable bias**
- **Investigating improvements to systematic uncertainties**

Acknowledgement

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