

Atmospheric μ data vs MC with KM3NeT detectors & prompt μ analysis

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on behalf of the KM3NeT Collaboration



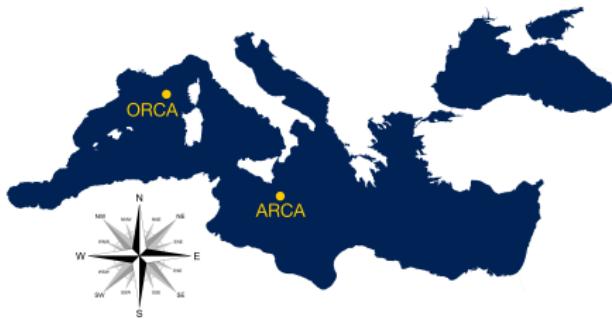
Flash talk
Poster #210



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KM3NeT basics

KM3NeT – **km³ Neutrino Telescope(s)**



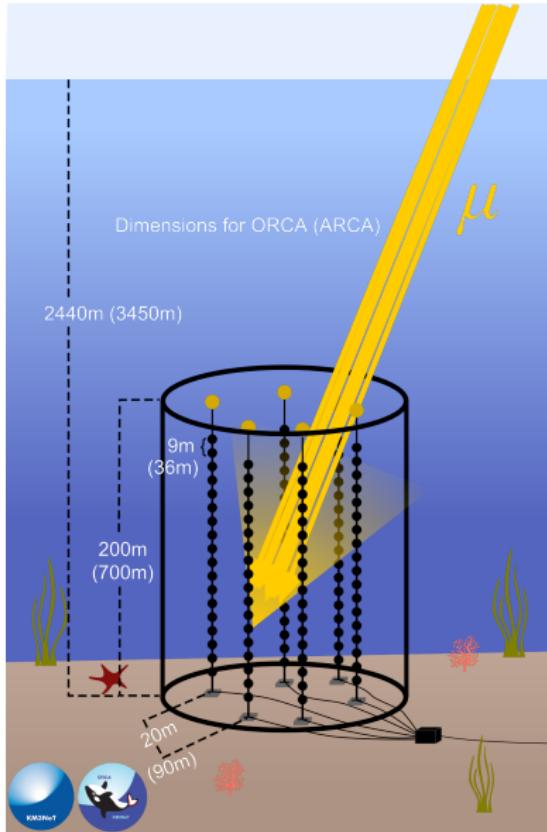
Detection Unit (**DU**): string with 18 DOMs

Example: ORCA4 has 4 DUs



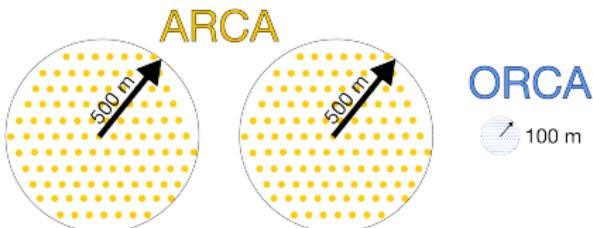
Digital Optical Module (**DOM**)
(31 PMTs + electronics etc.)

KM3NeT detectors



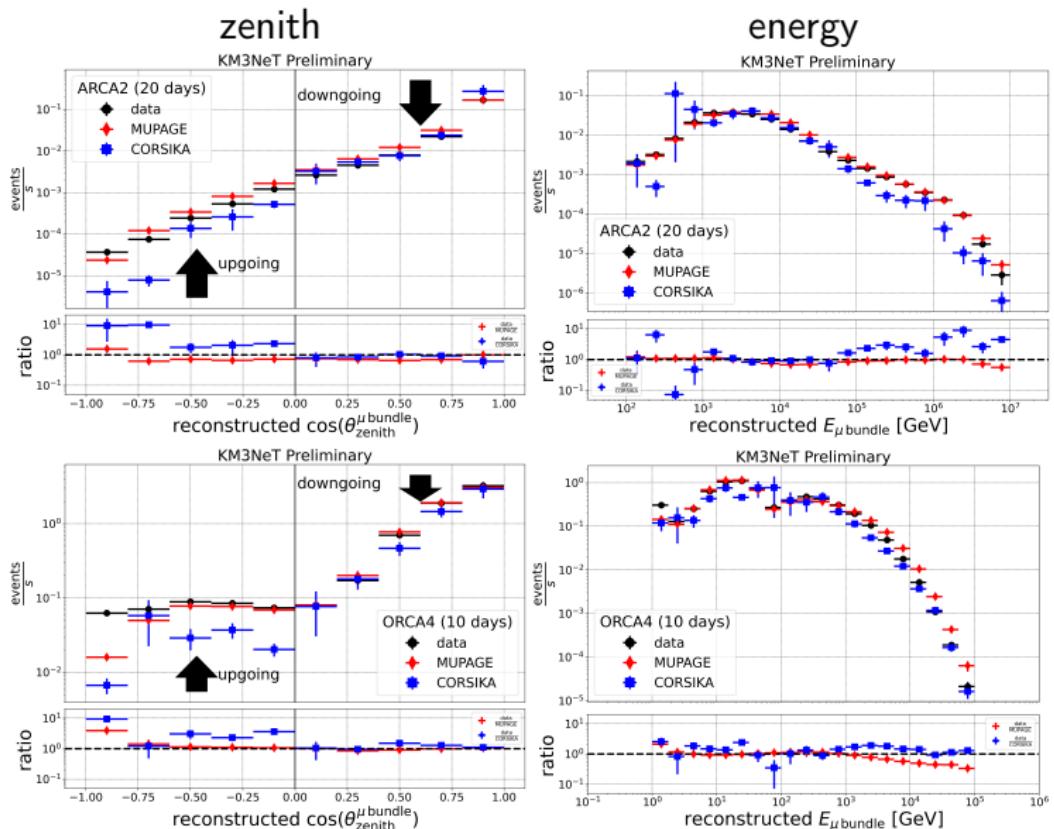
ORCA: Oscillation Research with Cosmics in the Abyss
(main goal: m_ν hierarchy)

	ARCA	ORCA
Depth	3.5 km	2.5 km
Volume	1 km ³ (1 Gton)	0.007 km ³ (7 Mton)
#DU	6 / 2x115	6 / 115



ARCA: Astroparticle Research with Cosmics in the Abyss
(main goal: v_{astro})

data vs MC

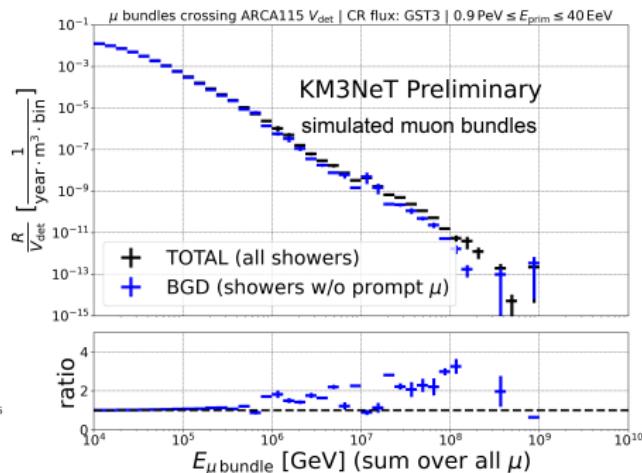
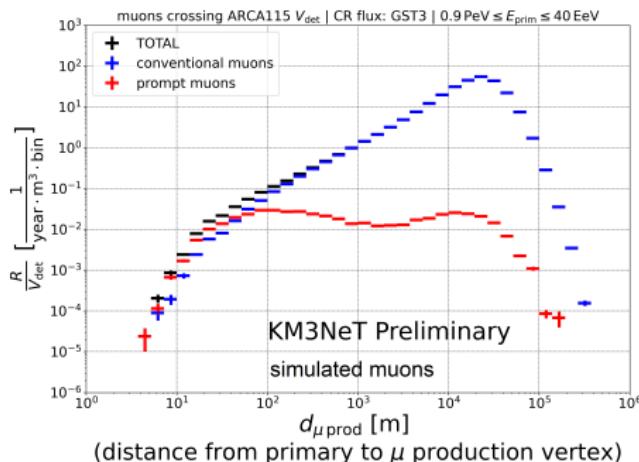
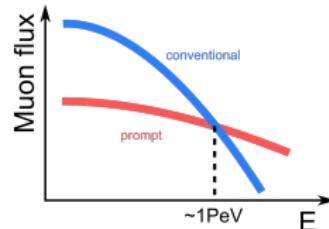


MC generators
CORSIKA v7.6400
MUPAGE v3r6
HE hadronic model
SIBYLL 2.3c
CR model
GST3
particles
μ bundles
stage
reco level
data livetime
ARCA2: 20 d
ORCA4: 10 d

Prompt muon analysis [MC-only]

Categories commonly used for μ flux:

- **conventional**: mostly π and K decays
- **prompt**: mostly heavy hadron decays



MC generators	HE hadronic model	CR model	particles	stage
CORSIKA v7.7400	SIBYLL 2.3d	GST3	μ bundles	detector level

Summary

data vs MC:

- data and MC are consistent
- next steps:
 - ▶ fix known issues
 - ▶ add systematic uncertainties
 - ▶ redo for ARCA6 and ORCA6

prompt muon analysis:

- first results encouraging
- next steps:
 - ▶ reconstruction level
 - ▶ sensitivities
 - ▶ compare with IceCube & theory



Fiona

Mecenas

Enjoy ICRC!